Equivariant heat and Schrödinger flows from Euclidean space to complex projective space

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Abstract. We study the equivariant harmonic map heat flow, Schrödinger maps equation, and generalized Landau–Lifshitz equation from \mathbb{C}^n to \mathbb{CP}^n . By means of a careful geometric analysis, we determine a new, highly useful representation of the problem in terms of a PDE for radial functions from \mathbb{C}^n to \mathbb{S}^2 . Using this new representation, we are able to write explicit formulas for the harmonic maps in this context, and prove that they all have infinite energy. We show that the PDEs admit a family of self-similar solutions with smooth profiles; these solutions again have infinite energy, and give an example of regularity breakdown. Then, using a variant of the Hasimoto transformation applied to our new equation for the dynamics, we prove a small-data global well-posedness result when $n = 2$. This is, to the best of our knowledge, the first global well-posedness result for Schrödinger maps when the complex dimension of the target is greater than 1.

In the final section we study a special case of the harmonic map heat flow corresponding to initial data valued in one great circle. We show that the $n = 2$ case of this problem is a borderline case for the standard classification theory for PDEs of its type.

1. Introduction

The harmonic map heat flow and the Schrödinger maps equations are natural generalizations of the linear heat and Schrödinger equations where the domain and range of the functions considered are manifolds and the Euclidean partial derivatives are replaced by covariant derivatives. In this article we will be exclusively discussing the setting when the base space is some Euclidean space \mathbb{R}^d and the target is a Kähler manifold N with complex structure *J*. The energy of a map $u: \mathbb{R}^d \to N$ is defined by $\mathcal{E}(u) = (1/2) \int_{\mathbb{R}^d} |du|^2 dV$. The Euler Lagrange operator $\tau(u)$ corresponding to $\mathcal E$ is calculated, in coordinates, to be $\tau(u) = \sum_{k=1}^{d} D_k \partial_k u$, where the D_k operators are covariant derivatives on N. The *harmonic map heat flow* is then the Cauchy problem given by

$$
u_t = \tau(u) = \sum_k D_k \partial_k u, \quad u(0) = u_0,
$$
\n(1)

while the *Schrödinger maps equation* is the Cauchy problem given by

$$
u_t = J\tau(u) = J\sum_k D_k \partial_k u, \quad u(0) = u_0.
$$
 (2)

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One can also consider the *generalized Landau–Lifshitz (GLL) equation*, defined for $\alpha \in$ $[0, \infty)$ and $\beta \in \mathbb{R}$ by

$$
u_t = (\alpha + \beta J)\tau(u) = (\alpha + \beta J) \sum_k D_k \partial_k u, \quad u(0) = u_0;
$$
 (3)

this corresponds, when the range is \mathbb{C} , to the PDE $u_t = (\alpha + i\beta)\Delta u$. Let us emphasize that the linearity of the equations in the familiar case when the target is $\mathbb C$ is special: in general these problems are non-linear because of curvature.

The harmonic map heat flow is a well-known and extensively studied problem. It was introduced in [\[7\]](#page-39-0) as a tool for studying the existence of harmonic maps. These are maps which satisfy $\tau(u) = D_k \partial_k u = 0$ and correspond to stationary solutions of all of the problems above. Vast work has been done on the harmonic map heat flow in the subsequent years; see, for example, [\[16\]](#page-39-1) for a textbook treatment. We mention only that it has been shown that for general N uniqueness of the harmonic map heat flow does not hold, and that one way to demonstrate non-uniqueness is through studying self-similar solutions, as is done in $[9, 10]$ $[9, 10]$ $[9, 10]$. This approach is used to prove a non-uniqueness result for the case of the flow for maps from $\mathbb{C}^2 \cong \mathbb{R}^4$ to \mathbb{CP}^2 ; see Section [5](#page-26-0) below.

As opposed to the harmonic map heat flow, the Schrödinger maps equation [\(2\)](#page-0-0) has been much less studied in general. For the setting we are considering here, that of the flow for maps $u: \mathbb{R}^d \to N$, local well-posedness in the Sobolev space $H^l(\mathbb{R}^d; N)$ for integer $l > d/2 + 1$ is established in [\[17\]](#page-39-4). One can see by scaling that $\dot{H}^{d/2}$ is critical for the problem, and significant work has been done on proving global well-posedness in this and other critical spaces in the special case when the target is the sphere $N = \mathbb{S}^2 [1, 3, 4, 13]$ $N = \mathbb{S}^2 [1, 3, 4, 13]$.

The case of the sphere is particularly attractive for two reasons. First, given the usual embedding $\mathbb{S}^2 \subset \mathbb{R}^3$, the Schrödinger maps equation becomes quite explicit. In this framework, the complex structure at the point u is simply given by the cross product in \mathbb{R}^3 , $J w = u \times w$. The derivative term is calculated to be $\sum_{k} D_{k} \partial_{k} u = \Delta u + |\nabla u|^{2} u$, where Δ and ∇ are the Laplacian and gradient operators for functions from \mathbb{R}^d to \mathbb{R}^3 . The Schrödinger maps equation thus becomes

$$
u_t = u \times (\Delta u + |\nabla u|^2 u), \quad x \in \mathbb{R}^d, \ u(x) \in \mathbb{S}^2 \subset \mathbb{R}^3. \tag{4}
$$

The second reason this case of the Schrödinger maps equation is appealing is that it is physically relevant. Equation [\(4\)](#page-1-0) is used to describe the dynamics of ferromagnetic spin systems, and is known in the physics community as the Heisenberg model. It is a special case of the equation

$$
u_t = (\alpha + \beta u \times)(\Delta u + |\nabla u|^2 u), \quad x \in \mathbb{R}^d,
$$
 (5)

which is the Landau–Lifshitz–Gilbert equation and is used to study the direction of magnetism in a solid. (The survey article [\[15\]](#page-39-6) discusses the physical relevance of these equations.) Equation (5) corresponds precisely to the GLL equation (3) in the case of maps $u: \mathbb{R}^d \to \mathbb{S}^2$. The work on small-data existence and uniqueness in a critical space for the

Schrödinger maps equation in this case of the sphere culminated in [\[3\]](#page-38-1), which furnished a global critical small-data well-posedness result in the Sobolev space $\dot{H}^{d/2}$.

A large body of work has been devoted to the sphere problem when the domain is \mathbb{R}^2 . The critical space is \dot{H}^1 , so the problem in this dimension is energy critical. It is also tractable to study because one can make an *equivariant ansatz* and thereby study a subproblem of the flow as a whole. The equivariant ansatz involves studying solutions of the form $u(r, \theta) = e^{m\theta R} f(r)$, where $f(r) \in \mathbb{R}^3$, $m \in \mathbb{Z}$, and R is the generator of rotations about the z-axis and given by the matrix

$$
R = \begin{pmatrix} 0 & 1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}.
$$

The overall picture that has developed can be described in terms of the harmonic maps, which have finite energy in this context, and whose existence is generally seen as a barrier to global well-posedness. In the case of radial maps, $m = 0$, there are no non-trivial harmonic maps and a global existence result for arbitrarily sized data in $H²$ has been established [\[12\]](#page-39-7). In the case when $m = 1$, the lowest energy level of the non-trivial harmonic maps is 4π ; for initial data with energy strictly smaller than this, global existence has been shown to hold $[2]$. On the other hand, in $[18]$, a set of initial data with energy arbitrarily close to 4π is constructed which generates finite time blow-up solutions. (This paper resolved the long-standing question of whether finite energy initial data could lead to finite time blowup.) Finite time blow-up solutions are also constructed in [\[19\]](#page-39-9). For $m > 3$, it has been shown that if the initial data has energy close to that of the harmonic maps then the solution is, in fact, global [\[12\]](#page-39-7).

Still in dimension 2, the equivariant ansatz can be made under the more general assumption that the target N is a complex surface with an \mathbb{S}^1 symmetry. This was originally done in [\[5\]](#page-38-4), where a critical well-posedness theory for equivariant data small in \dot{H} ¹ was developed. Under the same equivariant ansatz, [\[11\]](#page-39-10) do not pursue the Sobolev theory, and instead study the self-similar solutions of the flow. These are solutions of the form $u(x,t) = \psi(x/\sqrt{|t|})$ for a profile ψ . A family of such solutions with C^{∞} profiles is constructed, giving an example of regularity breakdown: these solutions are smooth at all times $t \neq 0$ but not smooth at $t = 0$. The study of these self-similar solutions is supplemented with a global critical small-data well-posedness theorem in a Lorentz space that is shown to include the self-similar data.

When the dimensions of the range and domain are larger than 2, but the same, it is still possible to formulate an equivariant ansatz, as will be shown in detail below. For the case of the Schrödinger maps equation for maps $u: \mathbb{C}^n \to \mathbb{CP}^n$, this equivariant ansatz is considered in [\[6\]](#page-39-11), where the existence of self-similar solutions is established.

The primary purpose of the present paper is to expand upon this previous work on the equivariant \mathbb{C}^n to \mathbb{CP}^n case, with a particular interest in establishing a global wellposedness theorem. Our central result is a new equation for the dynamics in this case [\(7\)](#page-4-0). This new equation is similar in structure to the GLL equation for maps to the sphere [\(5\)](#page-1-1),

and thus immediately opens up the possibility of applying research ideas developed for the sphere problem to the present context. Our global well-posedness result in Section [4](#page-18-0) in the case $n = 2$ is an example of this in practice.

1.1. Overview of the results

1.1.1. The equivariant ansatz and derivation of the equation. We consider maps $v: \mathbb{C}^n \to \mathbb{CP}^n$, where \mathbb{CP}^n is equipped with the Fubini–Study metric, for $n \geq 2$. The $n = 1$ case is the usual problem of \mathbb{R}^2 to the sphere because \mathbb{CP}^1 with the Fubini–Study metric is isometric to \mathbb{S}^2 . In what follows, *n* is the complex dimension and $d = 2n$ is the real dimension.

Recall that \mathbb{CP}^n can be viewed in terms of the homogeneous coordinates as points $(z_0, z_1, \ldots, z_n) \in \mathbb{C}^{n+1}$ under the identification $[z_0, z_1, \ldots, z_n] = [\alpha z_0, \alpha z_1, \ldots, \alpha z_n]$ for all $\alpha \in \mathbb{C} \backslash \{0\}$. Given a complex isometry A of \mathbb{C}^n we can construct an isometry \tilde{A} of \mathbb{CP}^n by the formula $\tilde{A}[z_0, z_1, \ldots, z_n] = [z_0, A(z_1, \ldots, z_n)]$; that is, we let A act on the last *n* coordinates in the homogeneous representation. A map $v: \mathbb{C}^n \to \mathbb{CP}^n$ is said to be *equivariant* if $v(Az) = \tilde{A}v(z)$ for all isometries A of \mathbb{C}^n and all points $z \in \mathbb{C}^n$. This ansatz is formally conserved by the flow. This assumption is strong and, as we show, implies that v is in fact of the form $v(z) = v((z_1, \ldots, z_n)) = [z_0, f(r)z_1, \ldots, f(r)z_n],$ where $r = |z|$ and $f : \mathbb{R}^+ \to \mathbb{C}$. We observe that for any $x \in \mathbb{R}$ we have $v((x, 0, \ldots, 0)) =$ $[z_0, f(r)x, 0, \ldots, 0]$, or namely that

$$
v(\mathbb{R}^+e_1) \subset \{ [z_0, z_1, 0, \ldots, 0] : z_0, z_1 \in \mathbb{C} \} \simeq \mathbb{CP}^1,
$$

so the image of a real ray is contained in a complex line. The Fubini–Study metric of \mathbb{CP}^n restricts to the Fubini–Study metric on this \mathbb{CP}^1 , so in fact the image of $v(\mathbb{R}^+e_1)$ is contained in a manifold isometric to \mathbb{S}^2 . The idea, now, is to parameterize this sphere in the usual embedding $\mathbb{S}^2 \subset \mathbb{R}^3$ and determine an equation on $u(r) = v(re_1) \in \mathbb{S}^2$. From the equivariant ansatz we can recover v from u .

By a computation we determine that the energy of v is given in terms of $u: \mathbb{C}^n \to \mathbb{S}^2$ by the formula

$$
\mathcal{E}(u) = \frac{1}{2} \int_{\mathbb{R}^{2n}} \left(|u_r|^2 + \frac{u_1^2 + u_2^2 + (2n - 2)|u - e_3|^2}{r^2} \right) dx,\tag{6}
$$

where $|u - e_3|$ is the Euclidean distance in \mathbb{R}^3 between u and the north pole of the sphere e₃, and $|u_r|$ is the Euclidean norm in \mathbb{R}^3 of u_r . Observe that in the case $n = 1$, we recover the usual energy for the equivariant $\mathbb{R}^2 \to \mathbb{S}^2$ problem, as we would expect. (See, for example, [\[2,](#page-38-3) p. 2].) For $n \ge 2$, one determines that any function u with finite energy is continuous and has a limit as $r \to \infty$; by inspecting the energy one sees that this limit must be the north pole e_3 .

The harmonic map heat flow, the Schrödinger maps equation, and the GLL equation for this equivariant case are now determined by calculating the variation of the energy. We

find that the GLL equation is given by

$$
u_t = (\alpha P + \beta u \times) \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n - 1}{r} \frac{\partial u}{\partial r} + \frac{2n - 2 + u_3}{r^2} e_3 \right),\tag{7}
$$

where P is the projection onto the tangent space $T_u \mathbb{S}^2$ and $u_3 = \langle u, e_3 \rangle$. The harmonic map heat flow corresponds to $\alpha = 1$ and $\beta = 0$, while the Schrödinger maps equation corresponds to $\alpha = 0$ and $\beta = 1$. This representation of the problem appears to be new. Its similarity to the corresponding equation for maps to the sphere is precisely what makes it so useful: it immediately opens up the possibility of applying some of the techniques that have been developed for the case of the sphere to the present setting too.

By taking the stereographic projection from the north pole $f(r) = (u_1(r) + iu_2(r))/\sqrt{\frac{2}{\pi}}$ $(1 + u₃(r))$ we determine the stereographic representation of the problem

$$
f_t = (\alpha + \beta i) \Big(f_{rr} - \frac{2\bar{f}}{1 + |f|^2} + \frac{2n - 1}{r} f_r - \frac{2n - 1}{r^2} f + \frac{1}{r^2} \frac{2|f|^2 f}{1 + |f|^2} \Big), \quad (8)
$$

where the function here is a radial map $f : \mathbb{R}^{2n} \to \mathbb{C}$. From this representation we see right away that the harmonic maps – that is, the stationary solutions – are given explicitly in this context by $f(r) = \alpha r$ for any $\alpha \in \mathbb{C}$. In the terms of the sphere coordinates, the harmonic maps are given by a type of stereographic projection,

$$
u(r) = \frac{1}{1 + |\alpha|^2 r^2} (2 \operatorname{Re}(\alpha) r, 2 \operatorname{Im}(\alpha) r, 1 - |\alpha|^2 r^2).
$$
 (9)

Again, this is consistent with the $n = 1$ case, where the equivariant harmonic maps from \mathbb{R}^2 to \mathbb{S}^2 are known to be stereographic projections. What is remarkable is that the analytic expressions for the harmonic maps are independent of n . This seems to suggest that, from the perspective of the theory of harmonic maps, \mathbb{CP}^n is the natural higher-dimensional analog of \mathbb{S}^2 .

However there is a difference for $n \ge 2$: observe that from [\(9\)](#page-4-1) we have $\lim_{r \to \infty} u(r) =$ $-e_3$, and so we find, by previous remarks on the energy, that in this equivariant context all of the non-trivial harmonic maps have infinite energy.

1.1.2. Self-similar solutions. After deriving the equation describing the dynamics, we first study the self-similar solutions of the problem for $n > 2$, which are given by $u(r, t) =$ first study the self-similar solutions of the problem for $n \ge 2$, which are given by $u(r, t) = \psi(r/\sqrt{t})$ for a profile $\psi(r) = u(r, 1)$. By substituting this ansatz into [\(7\)](#page-4-0) we determine the following ODE system on ψ :

$$
0 = (\alpha P + \beta \psi \times) \left(\frac{\partial^2 \psi}{\partial r} + \left(\frac{2n - 1}{r} + \frac{r}{2} \right) \frac{\partial \psi}{\partial r} + \frac{2n - 2 + \psi_3}{r^2} e_3 \right),
$$
 (10)

$$
\psi(0) = e_3,
$$

$$
\psi'(0) = v = (v_1, v_2, 0) \in T_{e_3} \mathbb{S}^2.
$$

As mentioned previously, the self-similar solutions for the Schrödinger maps equation in this equivariant setting have already been studied in [\[6\]](#page-39-11). However, using the representation [\(10\)](#page-4-2) we are able to simplify the analysis significantly. We are also able to extend the analysis by gaining more information on the convergence of the self-similar profile, and by treating the general GLL equation as well as the Schrödinger maps equation.

Theorem 1.1. *Fix* $n \geq 2$, $\alpha \geq 0$, and $\beta \in \mathbb{R}$. *For every* $v \in T_{e_3} \mathbb{S}^2$ there is a unique global *solution to* [\(10\)](#page-4-2)*. The solution is smooth for* $r > 0$ *. In the non-trivial case, when* $v \neq 0$ *, the solution has the following properties:*

- (1) *For all* $r > 0$, $\psi(r) \neq e_3$.
- (2) If $\alpha > 0$ then $|\psi_r| \lesssim 1/r^3$. If $\alpha = 0$ then $r|\psi_r| \to 0$ as $r \to \infty$.
- (3) If $v \neq 0$, there exists a point $\psi_{\infty} \in \mathbb{S}^2$, $\psi_{\infty} \neq e_3$, such that $\lim_{r \to \infty} \psi(r) = \psi_{\infty}$. *Consequently,* $\mathcal{E}(\psi) = \infty$.
- (4) *The limit* ψ_{∞} *depends continuously on* v; *in particular,* $\lim_{v\to 0} \psi_{\infty} = e_3$.

Our proof shows that the theorem also holds for $n = 1$ if $\alpha > 0$.

Our proof shows that the theorem also holds for $n = 1$ if $\alpha > 0$.
Because of the convergence, we see that $u(r, t) = \psi(r/\sqrt{t})$ is a solution of the GLL flow corresponding to the initial data $u(r, 0) \equiv \psi_{\infty}$.

Notice that in the case $\alpha > 0$ – that is, when there is some dissipation – we are able to prove faster convergence to 0 of ψ_r . In the case of the Schrödinger maps equation ($\alpha = 0$) the rate of convergence of ψ_r is insufficient to guarantee the convergence of ψ , so an additional argument is needed.

1.1.3. Global critical well-posedness. We next illustrate how methods for proving wellposedness of the Schrödinger maps equation for the sphere may be adapted to prove wellposedness of [\(7\)](#page-4-0). We specifically adapt the Hasimoto transformation method from [\[5\]](#page-38-4). For a smooth solution $u(r, t)$ of [\(7\)](#page-4-0) and a fixed time t, the map $r \mapsto u(r, t)$ defines a curve on \mathbb{S}^2 starting at e_3 . Choose any element $e \in T_{e_3} \mathbb{S}^2$ and consider the parallel transport $e(r)$ of this curve along $r \mapsto u(r,t)$. Because the tangent space at the point $u(r,t)$ of the sphere is two-dimensional, it is spanned by $e(r)$ and $Je(r) = u \times e(r)$. We may therefore define a complex-valued function q by the formula

$$
Re(q)e(r) + Im(q)Je(r) = qe(r) = u_r,
$$
\n(11)

precisely as in [\[5\]](#page-38-4). This equation is known as the *Hasimoto transformation*. It is chosen so that the function q will satisfy a "nice" non-linear Schrödinger equation, namely, an equation where the non-linearity does not contain derivatives. We derive the equation on q for all n, and in the case $n = 2$ – that is, for the equivariant GLL maps equation from \mathbb{C}^2 to \mathbb{CP}^2 – we provide the necessary estimates to prove the following small-data critical global well-posedness result.

Theorem 1.2. *Fix* $p \in [1, 2]$ *. Define r by* $1/r = 1/2 - 1/6p$ *and the spaces* X *and* X_0 *by the norms*

 $||q||_X = ||\nabla q||_{L_t^{3p} L_x^r}$ and $||q||_{X_0} = ||e^{it\Delta}(aq)||_X$,

where $a(x) = x_1/r$. There exists $\varepsilon > 0$ such that if $u_0: \mathbb{R}^{2n} \to \mathbb{S}^2$ is radial, q_0 is defined *by* [\(11\)](#page-5-0)*, and* $||q_0||_{X_0} \leq \varepsilon$ *, there is a unique global solution of the GLL equation* [\(7\)](#page-4-0) *for* $\beta > 0$ for $n = 2$ with the derivative term q in the space X.

Some remarks:

- This is, to the best of our knowledge, the first global well-posedness result for the Schrödinger maps equation where the target manifold has complex dimension greater than 1.
- The space X is at the scaling level of the equation.
- Because $(3p, r)$ is an admissible exponent pair for the Strichartz estimates for the Schrödinger equation, we have $||q||_{X_0} \lesssim ||\nabla(aq)||_{L^2} \lesssim ||\nabla q||_{L^2}$ and hence data q_0 whose derivative is small in L^2 are included in the well-posedness result.
- For $n > 2$ we are unable to provide the estimates to close the argument in an elementary way. A global well-posedness result for arbitrary n , proved using the Hasimoto transform or another method adapted from the research on the Schrödinger maps equation for the sphere, would be very satisfactory.

1.1.4. The "real" heat flow case. We finally study an interesting subproblem of the general equation (7) corresponding to the harmonic map heat flow with an additional condition on the initial data. Recall that for the linear heat equation, if one starts with real-valued data then the solution will be real valued for all time. On the other hand, if one starts the linear Schrödinger equation with real-valued data then the solution will, in general, be complex valued for future times. This shows that in the heat flow case there is a lower-dimensional subproblem when one restricts to real-valued data.

In our context, the analogous fact is that if one starts the harmonic map heat flow [\(7\)](#page-4-0) with initial data valued in a great circle passing through the north pole, the solution will continue to be valued on the same great circle for future times. For the GLL flow this is not true: the solution will spread out to the whole sphere. For the harmonic map heat flow one can thus fix a great circle and consider the problem for initial data valued on that circle. One expects the analysis of this subproblem to be easier as the dimension of the problem is reduced. However, because both the harmonic maps and the self-similar solutions are solutions of this type, it is still an interesting case to consider.

By parameterizing the great circle by its spherical distance from the north pole, one finds that the "real" heat flow is given by the PDE

$$
g_t = \frac{\partial^2 g}{\partial r^2} + \frac{2n - 1}{r} \frac{\partial g}{\partial r} + \frac{\eta(g)}{r^2},\tag{12}
$$

where $\eta(g) = \sin(2g) + (2n - 2) \sin(g)$. Equations of this type, which arise in the study of the equivariant harmonic map heat flow on spherically symmetric manifolds, have been extensively studied [\[9,](#page-39-2) [10\]](#page-39-3). There is a general theorem which, based on the structure of η , classifies the PDE into a uniqueness regime or a non-uniqueness regime. Our primary purpose here is to show that for $n = 2$ – that is, the problem of maps from \mathbb{C}^2 to \mathbb{CP}^2 – the

PDE [\(12\)](#page-6-0) is a *borderline case* for this classification theorem. We find that the dynamics of the PDE share some of features of the uniqueness regime, and some of the features of the non-uniqueness regime, but ultimately that non-uniqueness holds.

- **Theorem 1.3.** (i) *For* $n = 2$ *there is a weak non-constant solution of* [\(12\)](#page-6-0) *corresponding to the initial data* $g_0(r) \equiv \pi$. This solution is distinct from the constant *solution* $g(r, t) \equiv \pi$.
	- (ii) In the case $n > 3$, for each initial data in L^{∞} and each $T > 0$, there is at most *one solution of* [\(12\)](#page-6-0) *in* $L^{\infty}([0, T], L^{\infty})$ *.*

2. The equivariant ansatz and derivation of the equation

2.1. The equivariant ansatz

We consider maps $v: \mathbb{C}^n \to \mathbb{CP}^n$. In order to rigorously describe the equivariant ansatz, we recall more carefully the construction of \mathbb{CP}^n . One begins with vectors $z = (z_0, z_0)$ $z_1, \ldots, z_n) \in \mathbb{C}^{n+1} \setminus \{0\}$ and first identifies points $z \sim \lambda z$ where $\lambda \in \mathbb{R} \setminus \{0\}$. The resulting equivalence classes can be identified with points on the sphere $\mathbb{S}^{2n+1} \subset \mathbb{C}^{n+1}$. This sphere has the usual metric induced from \mathbb{C}^{n+1} . Now one defines the equivalence relation $z \sim e^{i\theta} z$ for $\theta \in \mathbb{R}$, and defines $\mathbb{CP}^n = \mathbb{S}^{2n+1}/\sim$. The Fubini–Study metric is the metric induced from \mathbb{S}^{2n+1} .

To make the equivariant ansatz, we first construct a special class of isometries on \mathbb{CP}^n in the following way. Take any complex isometry A of \mathbb{C}^n , and define $\hat{A}: \mathbb{C}^{n+1} \to \mathbb{C}^{n+1}$ by

$$
\hat{A}(z_0, z_1, \ldots, z_n) = (z_0, A(z_1, \ldots, z_n));
$$

that is, A acts on the last n coordinates of a point in \mathbb{C}^{n+1} . If A is a complex isometry of \mathbb{C}^n , then \hat{A} is clearly a complex isometry of \mathbb{C}^{n+1} . Now define a map \tilde{A} on \mathbb{CP}^n through the homogeneous coordinates by

$$
\tilde{A}[z_0, z_1, \dots, z_n] = [\hat{A}(z_0, z_1, \dots, z_n)] = [z_0, A(z_1, \dots, z_n)].
$$
\n(13)

The map \tilde{A} is well defined because A commutes with complex scalar multiplication.

Lemma 2.1. If A is a complex isometry of \mathbb{C}^n then \tilde{A} defined by [\(13\)](#page-7-0) is an isometry of \mathbb{CP}^n .

Proof. We have

$$
d_{\mathbb{CP}^n}(\tilde{A}[v], \tilde{A}[w]) = d_{\mathbb{CP}^n}([\hat{A}v], [\hat{A}w]) = \min_{\alpha, \beta \in [0, 2\pi]} d_{\mathbb{S}^{2n+1}}(e^{i\alpha} \hat{A}v, e^{i\beta} \hat{A}w)
$$

\n
$$
= \min_{\alpha, \beta \in [0, 2\pi]} 2 \arcsin\left(\frac{1}{2} d_{\mathbb{C}^{n+1}}(e^{i\alpha} \hat{A}v, e^{i\beta} \hat{A}w)\right)
$$

\n
$$
= \min_{\alpha, \beta \in [0, 2\pi]} 2 \arcsin\left(\frac{1}{2} d_{\mathbb{C}^{n+1}}(e^{i\alpha}v, e^{i\beta}w)\right) = d_{\mathbb{CP}^n}([v], [w]),
$$

where in the second to last equality we used that \hat{A} commutes with $e^{i\theta}$ and that \hat{A} is an isometry of \mathbb{C}^{n+1} .

We say a map $v: \mathbb{C}^n \to \mathbb{CP}^n$ is *equivariant* if $v(Az) = \tilde{A}v(z)$ for all complex isometries A of \mathbb{C}^n . We now show that this assumption implies a strong rigidity on v. Take any $z \in \mathbb{C}^n$ and write $v(z) = [w_0, w]$ for some $w_0 \in \mathbb{C}$ and $w \in \mathbb{C}^n$. Now consider any isometry A that fixes z. By the equivariant ansatz and $Az = z$ we have

$$
[w_0, Aw] = \tilde{A}u(z) = u(Az) = u(z) = [w_0, w],
$$

which implies that $Aw = w$, so A also fixes w. Because A is an arbitrary isometry that fixes z, we must in fact have $w = f(z)z$ for some $f(z) \in \mathbb{C}$, and hence $v(z) = [w_0, f(z)z]$ for all z. Moreover, we have

$$
[w_0, f(Az)Az] = v(Az) = \tilde{A}v(z) = [w_0, A(f(z)z)] = [w_0, f(z)Az],
$$

so $f(Az) = f(z)$. Because this holds for all isometries A, $f(z)$ is in fact a radial function and hence

$$
v(z) = [w_0, f(|z|)z]
$$
 (14)

for some function $f: \mathbb{R}^+ \to \mathbb{C}$.

We now observe that if $r \in \mathbb{R}^+$ then $v(re_1) = [w_0, f(r)r, 0, \ldots, 0]$. In other words,

 $v(\mathbb{R}^+e_1) \subset \{ [w_0, w_1, 0, \ldots, 0] : w_0, w_1 \in \mathbb{C} \} \simeq \mathbb{CP}^1.$

The Fubini–Study metric on \mathbb{CP}^n restricts to the Fubini–Study metric on \mathbb{CP}^1 , and so this \mathbb{CP}^1 is isometric to the sphere \mathbb{S}^2 . Moreover, the complex structure of \mathbb{CP}^n restricts to the standard complex structure of \mathbb{CP}^1 . In the usual embedding $\mathbb{S}^2 \subset \mathbb{R}^3$ this is given, as is well known, by $Jw = u \times w$ at the point $u \in \mathbb{S}^2$ and for all $w \in T_u \mathbb{S}^2$. We next parameterize this sphere and determine an equation for the function $r \mapsto v(re_1) \in \mathbb{S}^2$.

2.2. Derivation of the energy

The isometric identification between \mathbb{CP}^1 (with the Fubini–Study metric) and $\mathbb{S}^2 \subset \mathbb{R}^3$ (with the metric from the standard embedding) can be made through the isometric invertible map

$$
\mathbb{S}^2 \ni (a_1, a_2, a_3) \mapsto \frac{1}{\sqrt{2}(1+a_3)^{1/2}} [1+a_3, a_1 + ia_2] \in \mathbb{CP}^1,
$$
 (15)

where in this case the north pole $e_3 = (0, 0, 1)$ is mapped to the point $[1, 0] \in \mathbb{CP}^1$. In this identification the complex structure on \mathbb{CP}^1 is mapped to the standard complex structure on the sphere. Given an equivariant map $v: \mathbb{C}^n \to \mathbb{CP}^n$, we wish to write it in a form so that $v(re_1) \in \mathbb{CP}^1$ has the representation $[1 + a_3, a_1 + ia_2, 0, \ldots, 0]$. In fact, we can write v in the form

$$
v(z) = \frac{1}{\sqrt{2}(1+u_3)^{1/2}} \Big[1 + u_3, (u_1 + i u_2) \frac{z}{r} \Big]
$$
 (16)

for $u(r) = (u_1(r), u_2(r), u_3(r))$ satisfying $|u|_{\mathbb{R}^3} = 1$. When we substitute $z = re_1$ we recover essentially the representation in (15) , and hence u parameterizes the sphere in the correct, isometric, way.

(To see that $v(z) = [w_0, f(r)z]$ in [\(14\)](#page-8-1) can be written as in [\(16\)](#page-8-2), observe that by scaling we can assume that $(w_0, g(r)z) \in \mathbb{S}^{2n+1}$, which means $|w_0|^2 + |g(r)|^2 r^2 = 1$. We can also assume by scaling that $w_0 > 0$. This means, in fact, that $w_0 \in [0, 1]$, and we can also assume by scaling that $w_0 > 0$. This means, in fact, that $w_0 \in [0, 1]$, and hence there is a unique $u_3(r) \in [-1, 1]$ such that $\sqrt{2}(1 + u_3(r))^{1/2} = w_0$. We then define hence there is a unique $u_3(r) \in [-1, 1]$ such that $\sqrt{2}(1 + u_3(r))^{1/2} = w_0$. We then define $u_1 + i u_2 = rg(r)\sqrt{2}(1 - u_3)^{1/2}$, and substituting this in gives the representation above. The condition $|w_0|^2 + |g(r)r|^2 = 1$ translates into $|u|_{\mathbb{R}^3} = 1$.)

Proposition 2.2. *The energy is given in the* u *coordinates by*

$$
\mathcal{E}(v) = \frac{1}{2} \int_{\mathbb{R}^{2n}} |dv|^2 \, dx = \frac{1}{2} \int_{\mathbb{R}^{2n}} \left[|u_r|^2 + \frac{1}{r^2} [1 - u_3^2 + 2(2n - 2)(1 - u_3)] \right] dx. \tag{17}
$$

Proof. In order to calculate the energy density $|dv|^2$ of $v(z)$ we have to fix a basis for $T_z \mathbb{C}^n$, which will be 2*n*-dimensional, and calculate first derivatives of v with respect to this basis. For concreteness we view v as being valued in the sphere \mathbb{S}^{2n+1} ,

$$
v(z) = \frac{1}{\sqrt{2}(1+u_3)^{1/2}} \Big(1 + u_3, (u_1 + iu_2)^{\frac{z}{r}} \Big) \in \mathbb{S}^{2n+1} \subset \mathbb{C}^{n+1},\tag{18}
$$

and perform the computation there. The only adjustment needing to be made is as follows. Given a point $p \in \mathbb{S}^{2n+1}$, all points $e^{i\theta} p$ are mapped to the same point $[p] \in \mathbb{CP}^n$. By differentiating with respect to θ , it is apparent that in $T_p \mathbb{S}^{2n+1}$ the tangent direction $i p \in T_p \mathbb{S}^{2n+1}$ is contracted under the identification $p \sim e^{i\theta} p$. Hence when calculating derivatives at the level of \mathbb{S}^{2n+1} we take the usual Euclidean derivatives in \mathbb{C}^{n+1} , project onto $T_p \mathbb{S}^{2n+1}$, and then factor out the real subspace spanned by ip. In fact, the last two parts of this process amount to taking the complex projection

$$
P w = w - \langle w, p \rangle_{\mathbb{C}^{n+1}} p \tag{19}
$$

of derivative terms w. We have, of course, $|Pv|^2 = |v|^2 - |\langle v, p \rangle|^2$.

Let $\partial/\partial z_k$ and $\partial/\partial \overline{z_k}$ be the usual basis for $T_z \mathbb{C}^n$. For any vector $w_0 \in \mathbb{C}^n$ define $\partial/\partial w_0 = \sum_{m=1}^n w_0^m \partial/\partial z_m$ and $\partial/\partial \overline{w_0} = \sum_{m=1}^n \overline{w_0^m} \partial/\partial \overline{z_m}$. If $\{w_k\}_{k=1}^n$ is an orthonormal basis of \mathbb{C}^n then the derivatives $\{\partial/\partial w_k, \partial/\partial \overline{w_k}\}$ are an orthogonal basis for the tangent space and so, by the expression for $|dv|^2$ local in coordinates,

$$
|dv|^2 = 4\sum_{k=1}^n \left| P\frac{\partial v}{\partial w_k} \right|^2 + \left| P\frac{\partial v}{\partial \overline{w_k}} \right|^2.
$$
 (20)

One verifies the formulas at the point $z \in \mathbb{C}^n$:

$$
\frac{\partial r}{\partial w_0} = \frac{\langle w_0, z \rangle}{2r}, \quad \frac{\partial r}{\partial \bar{w}} = \frac{\langle z, w_0 \rangle}{2r}, \quad \frac{\partial z}{\partial w_0} = \frac{w_0}{|w_0|}, \quad \frac{\partial z}{\partial \overline{w_0}} = 0.
$$
 (21)

We then set $w_1 = z/|z|$ and define $w_2(z), \ldots, w_n(z)$ locally so that $\{w_k(z)\}_{k=1}^n$ is an orthonormal basis of \mathbb{C}^n for each z. In this setup, w_1 is the radial direction and w_k derivatives for $k > 2$ will be independent of radial terms.

Hence, for $k \geq 2$ we compute and find

$$
\frac{\partial v}{\partial w_k} = \frac{1}{\sqrt{2}(1+u_3)^{1/2}} \Big(0, (u_1 + i u_2) \frac{w_k}{r} \Big) \quad \text{and} \quad \frac{\partial v}{\partial \overline{w_k}} = 0.
$$

We see from [\(18\)](#page-9-0) that $\partial v/\partial w_k$ is complex orthogonal to v and so

$$
\left| P \frac{\partial v}{\partial w_k} \right|^2 = \left| \frac{\partial v}{\partial w_k} \right|^2 = \frac{1}{2(1+u_3)} \frac{u_1^2 + u_2^2}{r^2} = \frac{1-u_3}{2r^2},
$$

where in the step we used $u_1^2 + u_2^2 + u_3^2 = 1$.

We now differentiate with respect to w_1 and $\overline{w_1}$. In this case the radial terms will also be differentiated. We note, however, that when differentiating we can ignore the scaling be differentiated. We note, however, that when differentiating we can ignore the scaling
term $1/(\sqrt{2}(1+u_3)^{1/2})$: when this is differentiated we simply get a scalar multiple of $v(z)$, which disappears under the projection [\(19\)](#page-9-1). Hence,

$$
P\frac{\partial v}{\partial w^1} = \frac{1}{\sqrt{2}(1+u_3)^{1/2}} P\frac{\partial}{\partial w^1} \Big[1+u_3, (u_1+iu_2)\frac{z}{r} \Big]
$$

=
$$
\frac{1}{\sqrt{2}(1+u_3)^{1/2}} P\Big(u_3'\frac{1}{2}, (u_1'+iu_2')\frac{z}{2r} - (u_1+iu_2)\frac{z}{2r^2}\Big),
$$

and similarly,

$$
P\frac{\partial v}{\partial w^1} = \frac{1}{\sqrt{2}(1+u_3)^{1/2}} P\left(u'_3\frac{1}{2}, (u'_1+iu'_2)\frac{z}{2r} + (u_1+iu_2)\frac{z}{2r^2}\right).
$$

The difference in sign gives rise to the simplification

$$
\left|P\frac{\partial v}{\partial w^1}\right|^2 + \left|P\frac{\partial v}{\partial w^1}\right|^2
$$

=
$$
\frac{1}{4(1+u_3)}\left[\left|P\left(u_3'(u_1'+iu_2')\frac{z}{r}\right)\right|^2 + \left|P\left(0,(u_1+iu_2)\frac{z}{r^2}\right)\right|^2\right].
$$

Finally, a computation using the relations $u_1^2 + u_2^2 + u_3^2 = 1$ and $u_1u_1' + u_2u_2' + u_3u_3' = 0$ reveals that

$$
\begin{aligned}\n\left| P\left(u_3', (u_1' + iu_2')\frac{z}{r} \right) \right|^2 \\
&= \left| \left(u_3', (u_1' + iu_2')\frac{z}{r} \right) \right|^2 \\
&- \left| \left| \left(u_3', (u_1' + iu_2')\frac{z}{r} \right), \frac{1}{\sqrt{2}(1+u_3)^{1/2}} \left(1 + u_3, (u_1 + iu_2)\frac{z}{r} \right) \right| \right|^2 \\
&= |u_r|^2 - \frac{1}{2(1+u_3)} |(1+u_3)u_3' + (u_1 - iu_2)(u_1' + iu_2')|^2\n\end{aligned}
$$

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$$
= |u_r|^2 - \frac{1}{2(1+u_3)}|u'_3 + i(u_1u'_2 - u'_1u_2)|^2
$$

\n
$$
= |u_r|^2 - \frac{1}{2(1+u_3)}[(u'_3)^2 + u_1^2(u'_2)^2 + (u'_1)^2u_2^2 - 2u_1u'_1u_2u'_2]
$$

\n
$$
= |u_r|^2 - \frac{1}{2(1+u_3)}[(1-u_3^2)|u_r|^2] = \frac{1+u_3}{2}|u_r|^2
$$

and

$$
\left| P\left(0, (u_1 + iu_2)\frac{z}{r^2}\right) \right|^2 = \left| P\left(\frac{-1 - u_3}{r}, 0\right) \right|^2
$$

= $\frac{1}{r^2} \Big[(1 + u_3)^2 - \frac{1}{2(1 + u_3)} (1 + u_3)^4 \Big]$
= $\frac{(1 + u_3)(1 - u_3^2)}{2r^2}.$

We have then, by substituting these expressions into (20) ,

$$
|du|^2 = \frac{|u_r|^2}{2} + \frac{1}{r^2} \left[\frac{1 - u_3^2}{2} + 2(n - 1)(1 - u_3) \right],
$$

and then

$$
\mathcal{E}(v) = \int_{\mathbb{R}^{2n}} |dv|^2 dx
$$

= $\frac{1}{2} \int_{\mathbb{R}^{2n}} \left[|u_r|^2 + \frac{1}{r^2} [1 - u_3^2 + 2(2n - 2)(1 - u_3)] \right] dx$,

which completes the computation.

By the relations $1 - u_3^2 = u_1^2 + u_2^2$ and $|u - e_3|^2 = u_1^2 + u_2^2 + (u_3 - 1)^2 = 2(1 - u_3)$, we can equivalently write the energy in an L^2 form as

$$
\mathcal{E}(v) = \frac{1}{2} \int_{\mathbb{R}^{2n}} \left[|u_r|^2 + \frac{1}{r^2} [u_1^2 + u_2^2 + (2n - 2)|u - e_3|^2] \right] dx.
$$
 (22)

With this representation we determine the following result.

Proposition 2.3. There holds $||u_r||_{L^2}^2 \lesssim \mathcal{E}(u) \lesssim ||u_r||_{L^2}^2$.

Proof. The lower bound is obvious. For the upper bound, we observe that $u_1^2 + u_2^2 \leq$ $|u - e_3|^2$ and hence that

$$
\mathcal{E}(u) \leq \frac{1}{2} \Big(\|u_r\|_{L^2}^2 + (2n-1) \Big\| \frac{u - e_3}{r} \Big\|_{L^2}^2 \Big),
$$

and the result follows from the Hardy inequality $\|\phi/r\|_{L^2} \lesssim \|\phi_r\|_{L^2}$ for functions $\phi: \mathbb{R}^d \to$ \mathbb{R}^3 (see Theorem [A.1](#page-35-0) in the appendix). \blacksquare

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2.3. Variation of the energy, and the flow PDEs

In order to find the PDEs corresponding to the harmonic map heat flow, the Schrödinger maps equation, and the GLL equation, we need to calculate the variation of the energy, given by the formula

$$
\int_{\mathbb{R}^{2n}} \langle \tau(u), w \rangle_{T_u \mathbb{S}^2} \, dx = -\frac{d}{d\varepsilon} \Big|_{\varepsilon=0} \mathcal{E}(u + \varepsilon w)
$$

for all radial maps $w: \mathbb{R}^{2n} \to T\mathbb{S}^2$ such that $w(r) \in T_u \mathbb{S}^2$.

Proposition 2.4. *We have*

$$
\tau(u) = P_{T_u \mathbb{S}^2} \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n-1}{r} \frac{\partial u}{\partial r} + \frac{2n-2+u_3}{r^2} e_3 \right).
$$

Proof. Using the representation [\(17\)](#page-9-3) we find for $w \in T_u \mathbb{S}^2$,

$$
\frac{d}{d\varepsilon}\Big|_{\varepsilon=0} \mathcal{E}(u+\varepsilon w) = \frac{1}{2} \int_{\mathbb{R}^{2n}} 2\langle u_r, w_r \rangle_{\mathbb{R}^3} + \frac{1}{r^2} [-2u_3w_3 + 2(2n-2)(-w_3)]
$$

\n
$$
= - \int_{\mathbb{R}^{2n}} \langle u_{rr} + \frac{2n-1}{r} u_r, w \rangle_{\mathbb{R}^3} + \frac{1}{r^2} \langle (2n-2+u_3)e_3, w \rangle_{\mathbb{R}^3} dx
$$

\n
$$
= - \int_{\mathbb{R}^{2n}} \langle P_{T_u} g_2 \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n-1}{r} \frac{\partial u}{\partial r} + \frac{2n-2+u_3}{r^2} e_3 \right), w \rangle_{T_u} g_2 dx,
$$

and the formula follows.

In general, the harmonic map heat flow is given by $u_t = \tau(u)$, the Schrödinger maps equation is given by $u_t = J\tau(u)$, where J is the complex structure on the target, and the GLL equation is given by $u_t = (\alpha + \beta J) \tau(u)$ for $\alpha \geq 0$ and $\beta \in \mathbb{R}$. By the previous proposition, $\tau(u)$ is determined, while as discussed above, the complex structure in the u coordinates is precisely the usual complex structure on the sphere. We are therefore ready to write down the flow PDEs.

Definition 2.1. The equivariant generalized Landau–Lifshitz (GLL) problem from \mathbb{C}^n to \mathbb{CP}^n is the Cauchy problem for $u: \mathbb{R}^{2n} \to \mathbb{S}^2$ given by

$$
u_t(r,t) = (\alpha P + \beta u \times) \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n - 1}{r} \frac{\partial u}{\partial r} + \frac{2n - 2 + u_3}{r^2} e_3 \right),
$$
 (23)

$$
u(r,0) = u_0(r) \text{ with } u_0(0) = 0,
$$

for $\alpha \ge 0$ and $\beta \in \mathbb{R}$. We recall that P is the projection operator onto the tangent space at u. The case $\alpha = 1$ and $\beta = 0$ is the harmonic map heat flow. The case $\alpha = 0$ and $\beta = 1$ is the Schrödinger maps equation.

Note, by rescaling time we can always assume that $\alpha^2 + \beta^2 = 1$, which we do from now on.

By taking the stereographic projection $f(r) = (u_1 + iu_2)/(1 + u_3)$, with inverse given by

$$
(u_1, u_2, u_3) = \frac{1}{1 + |f|^2} (2 \text{ Re } f, 2 \text{ Im } f, 1 - |f|^2), \tag{24}
$$

we can determine the *stereographic representation* of the problem. With this stereographic projection, the north pole is mapped to the origin.

Proposition 2.5. *The GLL equation is given in the stereographic coordinates by*

$$
f_t = (\alpha + i\beta) \Big[f_{rr} - \frac{2\bar{f}f_r^2}{1 + |f|^2} + \frac{2n - 1}{r}f_r - \frac{2n - 1}{r^2}f + \frac{1}{r^2}\frac{2|f|^2f}{1 + |f|^2} \Big].
$$
 (25)

The proof involves substituting the expression for the stereographic projection [\(24\)](#page-13-0) into the PDE [\(23\)](#page-12-0) and computing; we omit this standard computation.

2.4. Classification of the harmonic maps in this context

The equivariant harmonic maps from \mathbb{C}^n to \mathbb{CP}^n are the time-independent solutions of [\(23\)](#page-12-0). Because the PDE has one space dimension, the time-independent problem is an ODE. In all, ϕ is harmonic if and only if

$$
0 = \phi \times \left(\frac{d^2\phi}{dr^2} + \frac{2n-1}{r}\frac{d\phi}{dr} + \frac{(2n-2) + \phi_3}{r^2}e_3\right),\tag{26}
$$

with the boundary conditions given by $\phi(0) = e_3$ and $\phi'(0) = v = (v_1, v_2, 0) \in T_{e_3} \mathbb{S}^2$. Writing the harmonic function ϕ in the stereographic coordinates as g, the ODE is

$$
0 = g_{rr} - \frac{2\bar{g}g_r^2}{1+|g|^2} + \frac{2n-1}{r}g_r - \frac{2n-1}{r^2}g + \frac{1}{r^2}\frac{2|g|^2g}{1+|g|^2}
$$

and the boundary conditions are $g(0) = 0$ and $g_r(0) = v_1 + iv_2$. Remarkably, we can solve this ODE explicitly with the linear function $g(r) = (v_1 + iv_2)r$. Moreover, because it is an ODE for which we have a uniqueness theory, $g(r) = (v_1 + iv_2)r$ is the unique solution. (See Theorem [A.3](#page-37-0) in the appendix for a local well-posedness theory for ODEs of this type.) Using the stereographic projection we can write the harmonic map in the sphere coordinates as

$$
\phi(r) = \frac{1}{1+|v|^2 r^2} (2rv_1, 2rv_2, 1-|v|^2 r^2)
$$

$$
= \frac{1}{1+|v|^2 r^2} (2rv + (1-|v|^2 r^2)e_3); \tag{27}
$$

in fact, ϕ is just a version of the stereographic projection itself. This is consistent with the well-known fact that the harmonic maps in the sphere $(n = 1)$ case are stereographic projections; what is interesting is that when n is incremented in ODE [\(26\)](#page-13-1), the new terms still cancel under this expression.

Qualitatively speaking, the harmonic maps in our context are quite simple: they start, when $r = 0$, at the north pole and, as r increases, move monotonically away from the north pole, converging to the south pole in the limit $r \to \infty$. By way of comparison, in the case of equivariant harmonic maps from the *d*-dimensional ball B^d to \mathbb{S}^d the situation is different [\[14\]](#page-39-12). For $3 \le d \le 6$ the harmonic maps oscillate about the south pole, while for $d \geq 7$ the harmonic maps approach the south pole monotonically, as here. In general, one finds that the equivariant harmonic maps usually fall into either an oscillatory regime or a monotonic regime [\[11\]](#page-39-10).

Finally, we note that while the expressions above for the harmonic maps are independent of n, there is a difference when $n > 2$. In the case of the sphere, $n = 1$, the energy of the stereographic projection is 4π . (This may be verified by substituting [\(27\)](#page-13-2) into the energy [\(17\)](#page-9-3) with $n = 1$, or by consulting [\[2\]](#page-38-3).) However, for $n \ge 2$ the energy is infinite. To see this it is sufficient to observe that $\lim_{r\to\infty} \phi(r) = -e_3$ and to use the following lemma.

Lemma 2.6. *Suppose that* $\mathcal{E}(u) < \infty$ *and* $n \ge 2$ *. Then* $\lim_{r \to \infty} u(r)$ *exists and equals e3.*

Proof. For any $r_2 > r_1 > 0$ we have

$$
|u(r_2) - u(r_1)| = \left| \int_{r_1}^{r_2} u_r(r) dr \right| \le \left(\int_{r_1}^{r_2} |u_r|^2 r^{2n-1} dr \right)^{1/2} \left(\int_{r_1}^{r_2} \frac{1}{r^{2n-1}} dr \right)^{1/2}
$$

 $\le C \mathcal{E}(u) r_1^{-n+1},$

which, because $n \geq 2$, shows that $\lim_{r \to \infty} u(r)$ exists. This means that in the energy [\(22\)](#page-11-0), the rightmost term in the integrand, $(1/r^2)(2n-2)|u(r) - e_3|^2r^{2n-1}$ converges as $r \to \infty$. For the energy to be finite, the limit must be 0. As $n \ge 2$, this implies that $\lim_{r \to \infty} u(r) = e_3.$

Corollary 2.7. When $n \geq 2$, the equivariant harmonic maps from \mathbb{C}^n to \mathbb{CP}^n all have *infinite energy.*

3. Self-similar solutions for $n \geq 2$

In this section we study the self-similar solutions for $n \geq 2$, which are solutions of the In this section we study the self-similar solutions for
form $u(r, t) = \psi(r/\sqrt{t})$ for a profile $\psi(r) = u(r, 1)$.

To determine a convenient equation for the profile, we take the GLL flow PDE [\(23\)](#page-12-0) and multiply both sides by $(\alpha u \times + \beta P)$. Using the relationship

$$
(\alpha u \times + \beta P)(\alpha P + \beta u \times) = (\alpha^2 + \beta^2)u \times = u \times
$$

(compare to $(\alpha i + \beta)(\alpha + \beta i) = i$) we may equivalently write the PDE as

$$
\alpha u \times u_t + \beta u_t = u \times \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n-1}{r} \frac{\partial u}{\partial r} + \frac{2n-2+u_3}{r^2} e_3\right).
$$
 (28)

We now substitute in $u(r, t) = \psi(r/\sqrt{t})$ to determine the ODE for the profile.

Definition 3.1. The self-similar problem for the GLL flow is given by the ODE

$$
-\frac{r}{2}\left(\alpha\psi\times\psi_r+\beta\psi_r\right)=\psi\times\left(\frac{\partial^2\psi}{\partial r^2}+\frac{2n-1}{r}\frac{\partial\psi}{\partial r}+\frac{(2n-2+\psi_3)}{r^2}e_3\right),\qquad(29)
$$

where $\psi: [0, \infty) \to \mathbb{S}^2$, subject to the initial conditions $\psi(0) = e_3$ and $\psi'(0) = v =$ $(v_1, v_2, 0) \in T_{e_3} \mathbb{S}^2$.

In the following sequence of lemmas we will prove Theorem [1.1,](#page-5-1) as stated in the introduction.

Lemma 3.1. *For every* $v = (v_1, v_2, 0) \in T_{e_3} \mathbb{S}^2$ *there is a unique global solution to* [\(29\)](#page-15-0)*. For* $r > 0$ *this global solution is smooth and, if* $v \neq 0$ *, satisfies* $\psi(r) \neq e_3$ *.*

Proof. Local existence and uniqueness in a neighborhood of the singular point $r = 0$ follows from Theorem [A.3](#page-37-0) in the appendix. For $r > 0$, ODE [\(29\)](#page-15-0) is smooth and local existence, uniqueness, and smoothness come from standard ODE theory. In order to prove global existence we establish an a priori bound on the derivative of ψ .

Define the function $A(r) = r^2 |\psi_r|^2$. We have

$$
A'(r) = 2r|\psi_r|^2 + 2r^2\psi_{rr} \cdot \psi_r.
$$
 (30)

In order to calculate $\psi_{rr} \cdot \psi_r$, we take the inner product of ODE [\(29\)](#page-15-0) with $\psi \times \psi_r$. Using the fact that if v or w is orthogonal to u, then $(u \times v) \cdot (u \times w) = v \cdot w$, and also the relation $v \cdot (u \times v) = 0$, we determine that

$$
-\frac{\alpha r}{2}|\psi_r|^2 = \psi_{rr} \cdot \psi_r + \frac{2n-1}{r}|\psi_r|^2 + \frac{2n-2+\psi_3}{r^2}e_3 \cdot \psi_r,
$$

and hence by solving for $\psi_{rr} \cdot \psi_r$ and substituting this into [\(30\)](#page-15-1) we find

$$
A'(r) = 2r|\psi_r|^2 - \left(\frac{2n-1}{r} + \frac{\alpha r}{2}\right)2r^2|\psi_r|^2 - 2(2n - 2 + \psi_3)(\psi_3)_r
$$

=
$$
-\left(\frac{2n-2}{r} + \frac{\alpha r}{2}\right)2A(r) - \frac{d}{dr}[(4n-4)\psi_3 + \psi_3^2].
$$
 (31)

Integrating this equation gives

$$
A(r) + \int_0^r \left(\frac{2n-2}{s} + \frac{\alpha s}{2}\right) 2A(s) \, ds = (4n-4)(1-\psi_3) + 1 - (\psi_3)^2. \tag{32}
$$

To bound $A(r)$, we observe that the integral on the left-hand side is non-negative because $A(s) \ge 0$ and $n \ge 2$, and so the left-hand side is bounded below by $A(r)$. On the other hand, we have $\psi_3 \in [-1, 1]$ and hence the right-hand side is bounded above by 8n. This then gives $A(r) \leq 8n$, and $|\psi_r| \leq 4n/r$. This proves global existence. (The constants 8n and $4n$ are, of course, not optimal; they are noted merely to show that the constants may be chosen independently of ψ .)

:

To prove that $\psi(r) \neq e_3$ for $r > 0$ we observe that the integral on the left-hand side in [\(32\)](#page-15-2) is increasing in r. In the non-trivial case $v \neq 0$, it is strictly increasing in a neighborhood of $r = 0$ because $A'(r) = r^2 |\psi_r|^2 \ge \varepsilon r^2$ in a neighborhood of $r = 0$. Hence in this case the integral is strictly positive for $r > 0$. Because $A(r) \ge 0$ we see that the left-hand side of [\(32\)](#page-15-2) is strictly positive and so

$$
(4n-4)(1-\psi_3)+1-(\psi_3)^2>0
$$

for $r > 0$. This gives $\psi_3(r) \neq 1$, which means $\psi(r) \neq e_3$.

Lemma 3.2. If $\alpha > 0$ we have $|\psi_r| \lesssim 1/r^3$.

Proof. Recall the bound $A(r) \leq 4n$. Using equation [\(31\)](#page-15-3) we have

$$
A'(r) \le -\alpha r A(r) - 2(2n - 2 + \psi_3)(\psi_3)_r
$$

\n
$$
\le -\alpha r A(r) + 2 \Big| \frac{2n - 2 + \psi_3}{r^{3/2} \sqrt{\alpha/2}} \Big| \cdot \Big| r^{3/2} \sqrt{\alpha/2} (\psi_3)_r \Big|
$$

\n
$$
\le -\alpha r A(r) + \Big(\frac{8n^3}{r^3 \alpha} + \frac{r^3 \alpha}{2} |\psi_r|^2 \Big) = -\frac{\alpha r}{2} A(r) + \frac{8n^3}{r^3 \alpha}
$$

Integrating this equation then gives $A(r) \lesssim A(1)e^{-\alpha r^2/8} + 1/r^4 \lesssim 1/r^4$ and $|\psi_r| \lesssim 1/r^3$. (The details of how this integration may be performed are given in Proposition [A.4](#page-37-1) in the appendix.)

Lemma 3.3. *There exists a point* $\psi_{\infty} \in \mathbb{S}^2$, $\psi_{\infty} \neq e_3$, such that $\lim_{r \to \infty} \psi(r) = \psi_{\infty}$. We *have the convergence rate inequality* $|\psi_{\infty} - \psi(r)| \leq 40n^2/r^2$. The profile ψ has infinite *energy.*

Proof. For $\alpha > 0$, the bound $|\psi_r| \lesssim 1/r^3$ implies convergence of ψ in the limit $r \to \infty$. In the case $\alpha = 0$, when there is no heat flow contribution, the decay on the derivative is less strong, and so a different argument is needed. However, in the proof we consider the general case as it is useful to know that the constant in the rate of convergence equation may be chosen independently of ψ .

We first multiply ODE [\(29\)](#page-15-0) by $(-\alpha \psi \times +\beta P)$. We have the relations $(-\alpha \psi \times +\beta P)$ $(\alpha \psi \times + \beta P) = (\alpha^2 + \beta^2)P = P$ and $(\psi \times)(\psi \times) = -P$ (compare to $(-\alpha i + \beta)(\alpha i +$ β) = 1 and $(i)(i) = -1$). We can thus write the equation as

$$
-\frac{r}{2}\psi_r = (\alpha P + \beta \psi \times)(\psi_{rr} + \frac{2n-1}{r}\psi_r + \frac{2n-2+\psi_3}{r^2}e_3),
$$

= $(\alpha + \beta \psi \times)(\frac{1}{r^{2n-1}}\frac{\partial}{\partial r}(r^{2n-1}\psi_r) + |\psi_r|^2\psi + \frac{2n-2+\psi_3}{r^2}Pe_3),$

where in the second equality we have moved the projection P inside and expanded $P \psi_{rr} = \psi_{rr} + |\psi_r|^2 \psi$. We divide through by r and integrate over $[r_1, r_2]$ to determine

that

$$
-\frac{1}{2}(\psi(r_2) - \psi(r_1))
$$

= $\int_{r_1}^{r_2} (\alpha + \beta \psi \times) \left(\frac{1}{r^{2n}} \frac{\partial}{\partial r} (r^{2n-1} \psi_r) + \frac{|\psi_r|^2 \psi}{r} + \frac{2n - 2 + \psi_3}{r^3} P e_3 \right) dr.$

Now integrating by parts in the first term yields

$$
-\frac{1}{2}(\psi(r_2) - \psi(r_1)) = \frac{[\alpha + \beta\psi(r_2) \times]\psi_r(r_2)}{r_2} - \frac{[\alpha + \beta\psi(r_1) \times]\psi_r(r_1)}{r_1} - \int_{r_1}^{r_2} (\alpha + \beta\psi \times)\left(\frac{-2n}{r^{2n+1}}r^{2n-1}\psi_r\right) dr + \int_{r_1}^{r_2} \left(\alpha\frac{|\psi_r|^2\psi}{r} + \frac{2n-2+\psi_3}{r^3}(\alpha P + \beta\psi \times)e_3\right) dr.
$$

Now using the bounds $|\psi(r)| = 1$ and $|\psi_r(r)| \leq 2n/r$ yields

$$
\frac{1}{2}|\psi(r_2)-\psi(r_1)| \leq \frac{2n}{r_2^2}+\frac{2n}{r_1^2}+\int_{r_1}^{r_2}\frac{4n^2}{r^3}dr+\int_{r_1}^{r_2}\left(\alpha\frac{4n^2}{r^3}+\frac{2n}{r^3}\right)dr \leq \frac{20n^2}{r_1^2},
$$

which implies the solution converges with the rate given in the statement of the lemma.

To see that the limit ψ_{∞} cannot be e_3 we consider equation [\(32\)](#page-15-2) again. As discussed previously, the integral in [\(32\)](#page-15-2) is strictly positive and non-increasing for $r > 0$. If $\delta > 0$ denotes the value of the integral at $r = 1$ we then have, for all $r > 1$,

$$
\delta \le \int_0^r \Big(\frac{2n-2}{s}+\frac{\alpha s}{2}\Big)2A(s)\,ds \le (4n-4)(1-\psi_3(r)) + 1-(\psi_3(r))^2.
$$

We therefore have

$$
\delta \le (4n-4)(1-\psi_3(\infty))+1-(\psi_3(\infty))^2,
$$

which gives $\psi_{\infty} \neq e_3$.

Because the limit is not e_3 , the profile has infinite energy by Lemma [2.6.](#page-14-0)

Lemma 3.4. When $\alpha = 0$ we have $\lim_{r \to \infty} r |\psi_r| = 0$.

Proof. It is sufficient to show that $\lim_{r\to\infty} A(r) = 0$. In the $\alpha = 0$ case, equation [\(32\)](#page-15-2) reads

$$
A(r) + \int_0^r \left(\frac{2n-2}{s}\right) 2A(s) \, ds = (4n-4)(1-\psi_3) + 1 - (\psi_3)^2.
$$

We know from the previous lemma that ψ_3 converges as $r \to \infty$. The integral also converges simply because it is non-decreasing; moreover, because it is bounded above (by 4n) it converges to a real number. We then have that $A(r)$ converges as $r \to \infty$. By examining the integral, which is finite in the limit, we see that we must have $\lim_{r \to \infty} A(r) = 0$. \blacksquare

Lemma 3.5. The limit ψ_{∞} is a continuous function of the initial data v. In particular, as $v \rightarrow 0$ *we have* $\psi_{\infty} \rightarrow e_3$.

Proof. For convenience we will denote the self-similar profile corresponding to initial data v by $\psi_v(r)$, and we will let $\psi_v(\infty)$ denote its limit as $r \to \infty$.

The ODE local existence results give that for any $r_0 > 0$ the map $v \mapsto \psi_v(r_0)$ is continuous.

We have previously established the bound, for $r_1 < r_2$, $|\psi_v(r_2) - \psi_v(r_1)| \leq 60n^2/r_1^2$. This shows that the map $v \mapsto \psi_v(r)$ converges to the map $v \mapsto \psi_v(\infty)$ uniformly, and hence that the map $v \mapsto \psi_v(\infty)$ is continuous.

Finally, we note that $\psi_0(r) \equiv 0$, $\psi_0(\infty) = 0$, and so $\lim_{v \to 0} \psi_v(\infty) = 0$, by continuity.

With this lemma, the proof of Theorem [1.1](#page-5-1) is complete.

4. Global critical well-posedness in dimension 2

In this section we prove a global critical small-data well-posedness theorem for the Schrödinger maps equation for equivariant maps from \mathbb{C}^n to \mathbb{CP}^n when $n = 2$. The equation may be written in the sphere coordinates as

$$
u_t(r,t) = u \times \left(\frac{\partial^2 u}{\partial r^2} + \frac{2n-1}{r} \frac{\partial u}{\partial r} + \frac{2n-2+u_3}{r^2} e_3\right),\tag{33}
$$

or equivalently as

$$
-u \times u_t(r,t) = \frac{\partial^2 u}{\partial r^2} + |u_r|^2 u + \frac{2n-1}{r} \frac{\partial u}{\partial r} + \frac{2n-2+u_3}{r^2} P_u e_3, \tag{34}
$$

where $P_u e_3$ is the projection of the vector $e_3 = (0, 0, 1)$ onto the tangent space at u.

Our proof relies on techniques that have been developed for the Schrödinger maps equation for the sphere. Because of the structural similarity between that equation and [\(34\)](#page-18-1), such techniques can be adapted here. We first use a form of the Hasimoto transform to determine an equation on a derivative term of u that has a simpler non-linearity. We then formulate the fixed point argument, and determine necessary estimates on the nonlinearity for the fixed point argument to be carried through. We conclude by proving these estimates in the case $n = 2$, thereby establishing Theorem [1.2.](#page-5-2)

We present our work in terms of the Schrödinger maps equation $(\alpha = 0)$; however, our proof is valid for the general GLL case when $\beta > 0$ because all the same estimates (in particular the Strichartz estimates) still apply.

4.1. Derivation of the PDE through the Hasimoto transform

The Hasimoto transform is an extensively used tool for proving well-posedness of the Schrödinger maps equations when the target is the sphere or a general complex surface. In geometric terms, it arises as follows. For fixed t , a smooth solution of [\(33\)](#page-18-2) will satisfy $u(0,t) = e_3$. The function $r \mapsto u(r,t)$ thus defines a curve in \mathbb{S}^2 starting at e_3 at $r = 0$. If one fixes a unit tangent vector $e(0) \in T_{e_3} \mathbb{S}^2$, one can consider the parallel transport $e(r)$ of this vector along the curve $r \mapsto u(r, t)$; the function $e(r)$ satisfies $D_r e(r) = \nabla_{u_r} e(r) = 0$. Now, because the tangent space at any point is two-dimensional, the vectors $e(r)$ and $Je(r)$ give a basis for the tangent space $T_{u(r)}\mathbb{S}^2$. Any derivative of u, or other element of the tangent space, can be expressed in terms of this basis. In our case, we define a complex-valued function q by the formula

$$
qe = (\text{Re }q + \text{Im }q \, J)e = u_r. \tag{35}
$$

We then determine an equation on q . The right-hand side is chosen so that q will satisfy a Schrödinger equation with a non-linearity that is easier to handle than that of [\(34\)](#page-18-1).

Lemma 4.1. *The function* q *satisfies the PDE*

$$
iq_t = q_{rr} + \frac{2n-1}{r}q_r - \frac{2n-1}{r^2}q + N(q),\tag{36}
$$

where the non-linear term $N(q)$ *is given by*

$$
N(q) = \frac{d}{dr} \bigg[-\frac{2n - 2 + u_3}{r^2} \int_0^r u_3(s) q(s) \, ds \bigg] + \alpha q, \tag{37}
$$

for a real-valued function ˛ *satisfying*

$$
\alpha_r = \text{Re}\bigg(\bar{q}q_r + \frac{|q|^2}{r} - \bar{q}\frac{2n-2+u_3}{r^2} \int_0^r u_3(s)q(s) \, ds\bigg). \tag{38}
$$

Proof. First, we recall that in the embedding $\mathbb{S}^2 \subset \mathbb{R}^3$ the covariant derivative of a vector field $v(r) \in T_{u(r)} \mathbb{S}^2$ is given by $D_r v = v_r + \langle u_r, v \rangle u$, where the inner product here is the usual inner product on \mathbb{R}^3 .

Now let p and q satisfy $pe = u_t$ and $qe = u_r$. We will determine three equations relating p, q , and u .

(1) Because *e* satisfies $D_r e = 0$ we have

$$
q_r e = D_r(qe) = D_r(u_r) = u_{rr} + |u_r|^2 u,
$$
\n(39)

which are the first two terms in the right-hand side of (34) . The next term in (34) is $((2n - 1)/r)$ qe. For the projection term we calculate, using $D_r e = 0$,

$$
\frac{d}{dr}\langle P_u e_3, e \rangle = \frac{d}{dr}\langle e_3 - \langle u, e_3 \rangle u, e \rangle = \langle D_r(e_3 - \langle u, e_3 \rangle u), e \rangle
$$

$$
= \left\langle \frac{d}{dr}(e_3 - \langle u, e_3 \rangle u) + \langle u_r, e_3 - \langle u, e_3 \rangle u \rangle u, e \right\rangle
$$

$$
= \langle -\langle u_r, e_3 \rangle u - \langle u, e_3 \rangle u_r + \langle u_r, e_3 \rangle u, e \rangle
$$

$$
= -u_3 \langle u_r, e \rangle = -u_3(r) \operatorname{Re} q(r).
$$

Using the fact that $u(0, t) = e_3$, so that $P_u e_3 = 0$ at $r = 0$, we have

$$
\langle P_u e_3, e \rangle = -\int_0^r u_3(s) \operatorname{Re} q(s) ds. \tag{40}
$$

An identical calculation for $\langle P_u e_3, Je \rangle$ gives, in total,

$$
P_u e_3 = -\left(\int_0^r u_3(s)q(s) \, ds\right) e(r).
$$

Plugging (39) and (40) into (34) then gives

$$
ip = q_r + \frac{2n-1}{r}q - \frac{2n-2+u_3}{r^2} \int_0^r u_3(s)q(s) ds.
$$
 (41)

(2) From the identity $D_r u_t = D_t u_r$ we find

$$
p_r e = D_r (pe) = D_r u_t = D_t u_r = D_t (qe) = q_t e + q D_t e. \tag{42}
$$

Because *e* is a parallel transport vector field, $|e|^2 = 1$ and so $0 = \frac{d}{dt}|e|^2 = 1$ $\langle D_t e, e \rangle$. The vector $D_t e$ is thus orthogonal to e. Because the tangent space is spanned by e and Je, we must have $D_t e = \alpha Je$ for some real-valued function α . Substituting this into [\(42\)](#page-20-1), we get $p_r e = q_t e + q \alpha Je$, or

$$
p_r = q_t + i\alpha q. \tag{43}
$$

T.

(3) To determine an equation on α we use the curvature relation $D_t D_r e = D_r D_t e +$ $R(u_t, u_t)e$, where R is the Riemann curvature tensor. On the sphere, $R(v, w)z =$ $\langle Jv, w \rangle$ *Jz*. Therefore, also using $D_r e = 0$, we find

$$
0 = D_r(\alpha Je) + \langle Ju_t, u_r \rangle Je = \alpha_r Je + \langle pJe, qe \rangle Je,
$$

which gives $\alpha_r = -\text{Im}(p\bar{q})$. Substituting the formula for p in [\(41\)](#page-20-2) gives equation [\(38\)](#page-19-1).

To determine an equation only on q we differentiate [\(41\)](#page-20-2) with respect to r, to find

$$
ip_r = q_{rr} + \frac{2n-1}{r}q_r - \frac{2n-1}{r^2}q + \frac{d}{dr}\bigg[-\frac{2n-2+u_3}{r^2}\int_0^r u_3(s)q(s) ds\bigg].
$$

Substituting the expression for p_r in [\(43\)](#page-20-3) gives equations [\(36\)](#page-19-2) and [\(37\)](#page-19-3).

4.2. Formulating the fixed point argument

We recall Theorem [1.2](#page-5-2) from the introduction.

Theorem (Theorem [1.2\)](#page-5-2). *Fix* $p \in [1, 2]$ *and define*

$$
\frac{1}{r} = \frac{1}{2} - \frac{1}{6p}
$$

and the spaces X *and* X_0 *given by the norms*

$$
\|q\|_X = \|\nabla q\|_{L_t^{3p}L_x^r} \quad \text{and} \quad \|q\|_{X_0} = \|e^{it\Delta}(aq)\|_X,
$$

where $a(x) = x_1/r$. There exists $\varepsilon > 0$ such that if $||q_0||_{X_0} \leq \varepsilon$ there is a unique global *solution of* [\(36\)](#page-19-2) *for* $n = 2$ *in the space* X.

We begin by determining a convenient Duhamel representation for the problem. Our Duhamel representation will be valid for all n , though we carry out the well-posedness argument for $n = 2$ only. In the following we will rely heavily on the Hardy inequalities given in Theorem [A.1](#page-35-0) and (70) in the appendix.

First, we absorb the linear term $-(2n-1)q/r^2$ into the Laplacian. To do this, we fix a function $a: \mathbb{S}^{2n-1} \to \mathbb{C}$ that satisfies $\Delta_{\mathbb{S}^{2n-1}} a = -(2n-1)a$. We may concretely choose $a(x) = x_1$. To see this, extend a to a function on all of \mathbb{R}^{2n} by $a(x/|x|)$. On the one hand, we have

$$
\Delta_{\mathbb{R}^{2n}}(ra(x/|x|)) = \Delta_{\mathbb{R}^{2n}}(x_1) = 0.
$$

Then, using the polar representation, $\Delta_{\mathbb{R}^{2n}} = \partial_{rr} + ((2n-1)/r)\partial_r + (1/r^2)\Delta_{\mathbb{S}^{2n-1}}$, we see that

$$
0 = \left[\partial_{rr} + \frac{2n-1}{r} \partial_r + \frac{1}{r^2} \Delta_{\mathbb{S}^{2n-1}}\right] (ra(x/|x|))
$$

= $\left[0 + \frac{2n-1}{r}\right] a(x/|x|) + \frac{r}{r^2} \Delta_{\mathbb{S}^{2n-1}} a(x/|x|),$

and so

$$
\Delta_{\mathbb{S}^{2n-1}}a(x/|x|) = -(2n-1)a(x/|x|).
$$

Now defining $w(x, t) = q(r, t) a(x/|x|)$, we see that

$$
\Delta_{\mathbb{R}^{2n}} w = \frac{\partial^2 q}{\partial r^2} a + \frac{2n-1}{r} \frac{\partial q}{\partial r} a - \frac{2n-1}{r^2} qa. \tag{44}
$$

This is exactly the Laplacian term in the PDE (36) multiplied by a.

In terms of estimates, we have the pointwise estimate $|\nabla a| \leq 1/r$, which is determined from a calculation. For Lebesgue estimates we have

$$
||w||_{L^p}^p = ||qa||_{L^p}^p = \int_0^\infty (|q(r)|^p \int_{rS^{2n-1}} |a(x/|x|)|^p dx) dr
$$

=
$$
\int_0^\infty (|q(r)|^p r^{2n-1} \int_{S^{2n-1}} |a(x/|x|)|^p dx) dr
$$

=
$$
C ||q||_{L^p}^p,
$$

where $C = ||a||_{L^p(\mathbb{S}^{2n-1})}/|\mathbb{S}^{2n-1}| < \infty$. We also have

$$
\|\nabla q\|_{L^p} \sim \|\nabla_r q\|_{L^p} \sim \|\nabla_r (aq)\|_{L^p} \lesssim \|\nabla_r (aq)\|_{L^p} + \|\nabla_\theta (aq)\|_{L^p} \sim \|\nabla w\|_{L^p},
$$

while

$$
\|\nabla w\|_{L^p} \sim \|\nabla(a)q\|_{L^p} + \|a\nabla_r q\|_{L^p} \lesssim \left\|\frac{1}{r}q\right\|_{L^p} + \|\nabla_r q\|_{L^p} \lesssim \|\nabla q\|_{L^p},
$$

so in conclusion $\|\nabla q\|_{L^p} \sim \|\nabla w\|_{L^p}$.

In the next lemma we rather carefully verify that we can recover solutions to the PDE for q from solutions to the PDE for w .

Lemma 4.2. *The PDE on* w *is given by*

$$
w_t = \Delta w + N(q)a \tag{45}
$$

or in Duhamel form by

$$
w(x,t) = e^{it\Delta}w(x,0) + i \int_0^r e^{i(t-s)\Delta} N(q(r,s))a(x/|x|,s) ds.
$$
 (46)

If the solution $w(x,t)$ *corresponding to initial data of the form* $w(x,0) = q(r)a(x/|x|)$ *is unique, then the solution is of the form* $w(x,t) = q(r,t)a(x/|x|)$ *, where q satisfies* [\(36\)](#page-19-2)*.*

Proof. To determine equation [\(45\)](#page-22-0) for w we simply multiply the PDE for q [\(36\)](#page-19-2) by a , and use expression [\(44\)](#page-21-0) for Δw . The Duhamel representation is standard.

We now show how solutions of [\(36\)](#page-19-2) may be recovered from solutions of the equation for w. Let w be a solution of [\(46\)](#page-22-1) and define $\tilde{w} = -(1/(2n - 1))\Delta_{\mathbb{S}^{2n-1}}w$. Assuming uniqueness we will show that $\tilde{w} = w$. We take the spherical Laplacian $-(1/(2n -$ 1)) $\Delta_{\mathbb{S}^{2n-1}}$ of [\(45\)](#page-22-0), noting that it commutes both with $\Delta = \Delta_{\mathbb{R}^{2n}}$ and $N(q)$, as $N(q)$ is radial. We then find that \tilde{w} satisfies the same PDE [\(45\)](#page-22-0) as w. Moreover, we have

$$
\widetilde{w}(x,0) = -\frac{1}{2n-1} \Delta_{\mathbb{S}^{2n-1}} w(x,0) = -\frac{1}{2n-1} \Delta_{\mathbb{S}^{2n-1}}[q(r)a(x/|x|)] = w(x,0),
$$

and so by uniqueness, $\tilde{w}(x,t) = -(1/(2n-1))\Delta_{\mathbb{S}^{2n-1}}w(x,t) = w(x,t)$. This means that w is a radial function times an eigenfunction of the Laplacian of the sphere of \mathbb{S}^{2n-1} with eigenvalue $-(2n - 1)$.

Let $T_k: \mathbb{R}^{2n} \to \mathbb{R}^{2n}$ be the linear map that multiplies the kth component of $x \in \mathbb{R}^{2n}$ by -1 and leaves the other components fixed. From the representation of a we see that for $k = 1$ we have $w_0(T_k x) = -w_0(x)$ while for $k \ge 2$ we have $w_0(T_k x) = w_0(x)$. By uniqueness, $x \mapsto w(x, t)$ inherits these properties also. But now the only eigenfunction of the Laplacian on the sphere with eigenvalue $-(2n - 1)$ with these symmetries is precisely a. Therefore $w(x, t) = q(r, t)a(x/|x|)$. Substituting this expression into the PDE (45) for w yields the PDE (36) for q. \blacksquare

By virtue of this lemma, we can perform the fixed point argument on w . The next lemma describes sufficient estimates for this fixed point argument to hold, and in the proof the fixed point argument is described.

4.3. Well-posedness when $n = 2$

For the remainder of this section we fix $n = 2$.

Before stating the lemma we fix some index notation. In the course of the proof we will need to handle Lebesgue space norms of quantities like q , q_r , q^2 , qq_r , and other quantities which scale like these. We are led to define the index

$$
\frac{1}{s(i,j)} = \frac{i+j}{4} - \frac{i}{6p}.
$$
 (47)

We will put items that scale like the product of i copies of q with a total of j derivatives in the space $L_x^{s(i,j)}$. For example, we will put q in $L_x^{s(1,0)}$, we will put q^2 in $L_x^{s(2,0)}$ and qq_r in $L_x^{s(2,1)}$. In this way, critical scaling is maintained throughout as, for example, $\|qq_r\|_{L_x^{s(2,1)}}$ is invariant under scaling.

The Strichartz inequality we will use is

$$
\left\| \int_0^t e^{i(t-s)\Delta} G \, ds \right\|_{L_t^{3p} L_x^r} \lesssim \|G\|_{L_t^p L_x^{s(3,1)}};
$$

this is classical: see, for example, [\[21\]](#page-39-13). The Hölder inequality is

$$
||fg||_{L_x^{s(i+k,j+m)}} \leq ||f||_{L_x^{s(i,j)}} \cdot ||g||_{L_x^{s(k,m)}},
$$

and the Sobolev, for $k < l$, is

$$
\|\nabla^k f\|_{L^{s(i,j)}_x} \lesssim \|\nabla^l f\|_{L^{s(i,j+l-k)}_x}.
$$

One verifies that these inequalities hold by checking the relevant exponent conditions.

Finally, note that $s(1, 1) = r$.

Lemma 4.3. *For Theorem* [1.2](#page-5-2) *to be true, it is sufficient that the following bounds hold:*

$$
\|\nabla N(q)\|_{L_x^{s(3,1)}} \lesssim \|\nabla q\|_{L_x^{s(1,1)}}^3,
$$
\n(48)

$$
\|\nabla(N(q_1)-N(q_2))\|_{L_x^{s(3,1)}} \lesssim \|\nabla(q_1-q_2)\|_{L_x^{s(1,1)}} (\|\nabla q_1\|_{L_x^{s(1,1)}}^2 + \|\nabla q_2\|_{L_x^{s(1,1)}}^2). \tag{49}
$$

Proof. Well-posedness follows by a fixed point argument for the operator

$$
Tw = e^{it\Delta}w(x,0) + i \int_0^t e^{i(t-s)\Delta} N(q(r,s))a(x,s) ds.
$$

We will show that T is a contraction mapping on a small ball around 0.

We first show that T maps a ball to itself. We have the bound

$$
||Tw||_X \le ||w_0||_{X_0} + \left\| \int_0^t e^{i(t-s)\Delta} \nabla(N(q)a) ds \right\|_{L_t^{3p} L_x^r}
$$

\$\leq ||w_0||_{X_0} + ||\nabla(N(q)a)||_{L_t^p L_x^{s(3,1)}}.\$

;

 \blacksquare

Considering the space norm of the integral, we have, by Hölder and Sobolev, and then conditions [\(48\)](#page-23-0),

$$
\|\nabla(N(q)a)\|_{L_x^{s(3,1)}} \lesssim \|\nabla(N)a\|_{L_x^{s(3,1)}} + \|N\nabla a\|_{L_x^{s(3,1)}} \le \|\nabla(N)\|_{L_x^{s(3,1)}} + \left\|\frac{1}{r}N\right\|_{L_x^{s(3,1)}} \le \|\nabla(N)\|_{L_x^{s(3,1)}} \lesssim \|\nabla q\|_{L_x^{s(1,1)}}^2 \lesssim \|\nabla w\|_{L_x^{s(1,1)}}^3
$$

and hence, as $r = s(1, 1),$

$$
||Tw||_X \leq ||w_0||_{X_0} + C ||w||_X^3.
$$

Now choose ε_0 so that $C \varepsilon_0^2 \le 1/2$, and let $\varepsilon \le \varepsilon_0$. Then, if $||w_0||_{X_0} \le \varepsilon/2$ and $||w||_X \le \varepsilon$, we have

$$
||Tw||_X \leq \frac{\varepsilon}{2} + C\varepsilon^3 \leq \frac{\varepsilon}{2} + (C\varepsilon_0^2)\varepsilon \leq \varepsilon,
$$

and so T maps every ε ball into itself, for ε sufficiently small, assuming the initial data satisfies the bound $||w_0||_{X_0} \leq \varepsilon/2$.

We next show that T is a contraction in a sufficiently small ball around 0. Let w_1 and w_2 be two solutions, with radial parts q_1 and q_2 respectively. We have

$$
Tw_1 - Tw_2 = \int_0^t e^{i(t-s)\Delta} (\nabla (N(q_1)a) - \nabla (N(q_2)a)) ds,
$$

which gives, using [\(49\)](#page-23-1),

$$
\begin{split}\n\|\nabla(N(q_1)a) - \nabla(N(q_2)a)\|_{L_x^{s(3,1)}} \\
&\lesssim \|\nabla(N(q_1) - N(q_2))a\|_{L_x^{s(3,1)}} + \|(N(q_1) - N(q_2))\nabla a\|_{L_x^{s(3,1)}} \\
&\lesssim \|\nabla(N(q_1) - N(q_2))a\|_{L_x^{s(3,1)}} + \left\|\frac{1}{r}(N(q_1) - N(q_2))\right\|_{L_x^{s(3,1)}} \\
&\lesssim (\|\nabla q_1\|_{L_x^{s(1,1)}}^2 + \|\nabla q_2\|_{L_x^{s(1,1)}}^2)\|\nabla(q_1 - q_2)\|_{L_x^{s(1,1)}} \\
&\lesssim (\|\nabla w_1\|_{L_x^{s(1,1)}}^2 + \|\nabla w_2\|_{L_x^{s(1,1)}}^2)\|\nabla(w_1 - w_2)\|_{L_x^{s(1,1)}},\n\end{split}
$$

and so

$$
||Tw_1 - Tw_2||_X \lesssim (||w_1||_X^2 + ||w_2||_X^2) ||w_1 - w_2||_X,
$$

and hence by choosing the ball small enough, T is a contraction.

Lemma 4.4. *When* $n = 2$ *the bounds* [\(48\)](#page-23-0) *and* [\(49\)](#page-23-1) *hold.*

Proof. Write

$$
N = \frac{d}{dr} \left(-\frac{2n-2+u_3}{r^2} \int_0^r u_3(s)q(s) \, ds \right) + \alpha q =: N_1 + N_2,
$$

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and recall

$$
\alpha_r = \text{Re}\bigg(q_r\bar{q} + \frac{|q|^2}{r} - \bar{q}\frac{2n-2+u_3}{r^2} \int_0^r u_3(s)q(s) \, ds\bigg). \tag{50}
$$

We will prove the bounds for N_1 first.

We have

$$
\|\nabla N_{1}\|_{s(3,1)} \lesssim \left\|\nabla^{2}\left(-\frac{2n-2+u_{3}}{r^{2}}\int_{0}^{r}u_{3}(s)q(s) ds\right)\right\|_{L_{x}^{s(3,1)}} \n\lesssim \left\|(\nabla^{2}u_{3})\frac{1}{r^{2}}\int_{0}^{r}u_{3}(s)q(s) ds\right\|_{L_{x}^{s(3,1)}} \n+ \left\|-\frac{2n-2+u_{3}}{r^{2}}\frac{1}{r^{2}}\int_{0}^{r}u_{3}(s)q(s) ds\right\|_{L_{x}^{s(3,1)}} \n+ \left\|-\frac{2n-2+u_{3}}{r^{2}}\nabla(u_{3}(r)q(r)) ds\right\|_{L_{x}^{s(3,1)}} \n=: A + B + C
$$

From the equation $u_{rr} + |u_r|^2 u = q_r e$, we have $|u_{rr}| \le |q|^2 + |q|$ pointwise. Therefore, for A,

$$
A \leq \left\| (|q|^2 + |q_r|) \frac{1}{r^2} \int_0^r u_3(s)q(s) ds \right\|_{L_x^{s(3,1)}} \n= \|q\|_{L_x^{s(1,0)}} \left\| \frac{q}{r} \right\|_{L_x^{s(1,1)}} \left\| \frac{1}{r} \int_0^r u_3(s)q(s) ds \right\|_{L_x^{s(1,0)}} \n+ \|q_r\|_{L_x^{s(1,1)}} \left\| \frac{1}{r^2} \int_0^r u_3(s)q(s) ds \right\|_{L_x^{s(2,0)}} \n= \|q\|_{L_x^{s(1,0)}} \|q/r\|_{L_x^{s(1,1)}} \|u_3(s)q(s)\|_{L_x^{s(1,0)}} + \|q_r\|_{L_x^{s(1,1)}} \left\| \frac{1}{r} u_3(s)q(s) ds \right\|_{L_x^{s(2,0)}} \n\lesssim \|q\|_{L_x^{s(1,0)}}^2 \|q_r\|_{L_x^{s(1,1)}} + \|q_r\|_{s(1,1)} \left\| \frac{1}{r} u_3(r)q(r) \right\|_{L_x^{s(2,0)}} \n\lesssim \|q\|_{L_x^{s(1,0)}}^2 \|q_r\|_{L_x^{s(1,1)}} + \|q_r\|_{s(1,1)} \left\| \frac{u_3}{r} \right\|_{L_x^{s(1,0)}} \|q(r)\|_{L_x^{s(1,0)}} \lesssim \|q_r\|_{L_x^{s(1,1)}}^3. \tag{51}
$$

For B , we have

$$
B \lesssim \left\| \frac{2n - 2 + u_3}{r} \right\|_{L_x^{s(1,0)}} \left\| \frac{1}{r^3} \int_0^r u_3(s) q(s) ds \right\|_{L_x^{s(2,1)}} \n\lesssim \|u_r\|_{L_x^{s(1,0)}} \left\| \frac{1}{r^2} u_3(r) q(r) \right\|_{L_x^{s(2,1)}} \n\lesssim \|u_r\|_{L_x^{s(1,0)}} \left\| \frac{u_3(r)}{r} \right\|_{s(1,0)} \left\| \frac{q(r)}{r} \right\|_{L_x^{s(1,1)}} \n\lesssim \|u_r\|_{L_x^{s(1,0)}}^2 \|q_r\|_{L_x^{s(1,1)}} \lesssim \|q_r\|_{L_x^{s(1,1)}}^3.
$$
\n(52)

For C , we have

$$
C \lesssim \left\| \frac{2n - 2 + u_3}{r^2} \nabla(u_3(r)q(r)) ds \right\|_{L_x^{s(3,1)}} \n\lesssim \left\| \frac{2n - 2 + u_3}{r} \right\|_{L_x^{s(1,0)}} \left(\left\| \frac{1}{r} \nabla(u_3)q \right\|_{L_x^{s(2,1)}} + \left\| \frac{1}{r}(u_3) \nabla q \right\|_{L_x^{s(2,1)}} \right) \n\lesssim ||q||_{L_x^{s(1,0)}} \left(\left\| \nabla u_3 \right\|_{L_x^{s(1,0)}} \left\| \frac{q}{r} \right\|_{L_x^{s(1,1)}} + \left\| \frac{u_3}{r} \right\|_{L_x^{s(1,0)}} \left\| \nabla q \right\|_{L_x^{s(1,1)}} \right) \n\lesssim ||q_r||_{s(1,1)}^3.
$$
\n(53)

The three estimates [\(51\)](#page-25-0), [\(52\)](#page-25-1), and [\(53\)](#page-26-1) together give the estimate $\|\nabla N_1\|_{L_x^{s(3,1)}} \lesssim$ $||q_r||^3$ $L_{x}^{s(1,1)}$.

As for N_2 , we have

$$
\begin{split} \|\nabla(\alpha q)\|_{L_{x}^{s(3,1)}} &\lesssim \|\alpha_r q\|_{L_{x}^{s(2,1)}} + \|\alpha q_r\|_{L_{x}^{s(3,1)}} \\ &\lesssim \|\alpha_r\|_{L_{x}^{s(2,1)}} \|q\|_{L_{x}^{s(1,0)}} + \|\alpha\|_{L_{x}^{s(2,0)}} \|q_r\|_{L_{x}^{s(1,1)}} \\ &\lesssim \|\alpha_r\|_{L_{x}^{s(2,1)}} \|q_r\|_{L_{x}^{s(1,1)}} . \end{split} \tag{54}
$$

Then, using the expression for α_r in [\(50\)](#page-25-2) and the fact that $u_3 \in L^{\infty}$,

$$
\|\alpha_r\|_{s(2,1)} \lesssim \|qq_r\|_{L_x^{s(2,1)}} + \left\|\frac{q^2}{r}\right\|_{L_x^{s(2,1)}} + \left\|\frac{2n-2+u_3}{r^2}q \int_0^r u_3(s)q(s) \, ds\right\|_{s(2,1)}
$$

\n
$$
\lesssim \|q\|_{L_x^{s(1,0)}} \Big(\|q_r\|_{L_x^{s(1,1)}} + \left\|\frac{q}{r}\right\|_{L_x^{s(1,1)}} \Big) + \left\|\frac{q}{r}\right\|_{L_x^{s(1,1)}} \left\|\frac{1}{r} \int_0^r u_3(s)q(s) \, ds\right\|_{L_x^{s(1,0)}}
$$

\n
$$
\lesssim \|q_r\|_{L_x^{s(1,1)}}^3 + \left\|\frac{q}{r}\right\|_{L_x^{s(1,1)}} \|q\|_{L_x^{s(1,0)}} \lesssim \|q_r\|_{L_x^{s(1,1)}}^3. \tag{55}
$$

Estimates [\(54\)](#page-26-2) and [\(55\)](#page-26-3) give $\|\nabla N_2\|_{L_x^{s(3,1)}} \lesssim \|q_r\|_{L_x^{s(3,3)}}^3$ $L_x^{s(1,1)}$ and hence [\(48\)](#page-23-0). Estimate [\(49\)](#page-23-1) follows from an identical argument. \blacksquare

Theorem [1.2](#page-5-2) is thus established.

5. The "real" heat flow case

In this section we will discuss what might be termed the "real" equivariant heat flow from \mathbb{C}^n to \mathbb{CP}^n . In the case when $\alpha = 1$ and $\beta = 0$, that is, for the harmonic map heat flow, it is possible to make an ansatz which further reduces the problem. In terms of the spherical coordinates,

$$
u_t = u_{rr} + u|u_r|^2 + \frac{2n-1}{r}u_r + \frac{2n-2+u_3}{r^2}(e_3 - u\langle u, e_3 \rangle),
$$
 (56)

$$
u(r, 0) = v(r) = (v_1(r), v_2(r), 0) \in T_{e_3} \mathbb{S}^2,
$$

Figure 1. In the case of the harmonic map heat flow, if the initial data takes values in one great circle (here the great circle spanned by v_0 and e_3), then the solution will be valued in the same great circle for future times. Both the harmonic maps and the self-similar solutions constructed in Section [3](#page-14-1) are of this type.

this ansatz involves assuming that the initial data is valued in one great circle passing through the north pole; that is, the initial data is of the form $c(r)e_3 + d(r)v_0$. (See Fig-ure [1.](#page-27-0)) In this case, for $t > 0$ the solution will continue to be valued in the same great circle. To see this, let $w_0 = v_0 \times e_3$ and let $a(r, t) = u(r, t) \cdot w_0$. By taking the inner product of equation [\(56\)](#page-26-4) with aw_0 we have

$$
aa_t = aa_{rr} - a^2|u_r|^2 + \frac{2n-1}{r}aa_r + \frac{2n-2+u_3}{r^2}(-a^2u_3)
$$

$$
\leq aa_{rr} + \frac{2n-1}{r}aa_r + \frac{2n-2}{r^2}a^2.
$$

We next integrate this equation. For $n \geq 2$ we use the Hardy inequality with best constant $4/d^2 = 4/(2n - 2)^2$ to determine that

$$
\frac{d}{dt}\frac{1}{2}\int_{\mathbb{C}^n} (a)^2 dx \le \sigma_{2n-1} \int_0^\infty a \frac{\partial}{\partial r} (r^{2n-1}a_r) dr + (2n-2) \left\| \frac{a}{r} \right\|_{L^2}^2
$$

= $-\sigma_{2n-1} \int_0^\infty (a_r)^2 r^{2n-1} dr + (2n-2) \frac{4}{(2n-2)^2} \|a_r\|_{L^2}^2 \le 0,$

and hence $a(r, t) = 0$ for all time. For $n = 1$ we obtain the same inequality because the last term is 0. The solution is therefore a linear combination of v_0 and e_3 .

In terms of the stereographic representation of the problem,

$$
f_t = f_{rr} - \frac{2\bar{f}f_r^2}{1+|f|^2} + \frac{2n-1}{r}f_r(r) + \frac{2n-1}{r^2}f(r) + \frac{2|f|^2f}{1+|f(r)|^2},
$$

$$
f(r, 0) = f_0(r),
$$

the ansatz is that the initial data is of the form $f(r, 0) = b(r)e^{i\theta}$ for some real-valued function $b(r)$ and a constant θ . The solution will then be of the form $f(r, t) = b(r, t)e^{i\theta}$ for the same constant θ and for some real-valued function $b(r, t)$. This motivates the terminology "real heat flow".

It is not surprising that this problem is simpler to analyze, and in fact with this assumption we are able to say more about the dynamics of the problem. On the other hand, this problem is still interesting because both the harmonic maps and the self-similar solutions constructed in Section [3](#page-14-1) fit into this context. In fact, the harmonic maps are given in stereographic coordinates by $f(r,t) = \alpha r = |\alpha| r e^{i\theta}$. The initial data for a self-similar solution is just a point, so the initial data is valued in the great circle passing through that point and the north pole.

We will now describe how, based on the ansatz just described, a simpler PDE on the solution may be determined. As the solution is valued on a great circle we can perform a change of variables, $u(r, t) = \cos(g)e_3 + \sin(g)v_0$, for an unknown real-valued g. Geometrically, g is the spherical distance between $u(r, t)$ and e_3 . We calculate $u_r =$ $g_r(-\sin(g)e_3 + \cos(g)v_0)$ and

$$
u_{rr} = g_{rr}(-\sin(g)e_3 + \cos(g)v_0) + g_r^2(-\sin(g)e_3 - \cos(g)v_0)
$$

= $g_{rr}(-\sin(g)e_3 + \cos(g)v_0) - u|u_r|^2$.

Substituting these into [\(56\)](#page-26-4) gives

$$
g_t(-\sin(g)e_3 + \cos(g)v_0) = \left(g_{rr} + \frac{2n-1}{r}g_r\right)(-\sin(g)e_3 + \cos(g)v_0) + \frac{2n-2 + \cos(g)}{r^2}(e_3 - \cos(g)(\cos(g)e_3 + \sin(g)v_0)).
$$

Taking the inner product of this equation with $-\sin(g)e_3 + \cos(g)v_0$ then yields the equation on g.

Definition 5.1. The real heat flow problem is the Cauchy problem

$$
g_t = g_{rr} + \frac{2n-1}{r}g_r - \frac{1}{r^2} \Big[(2n-2)\sin(g) + \frac{1}{2}\sin(2g) \Big] \tag{57}
$$

subject to the initial condition $g(r, 0) = g_0(r)$.

For convenience we let $\eta(x) = (2n - 2) \sin(x) + \sin(2x)/2$.

Definition 5.2. The stationary real heat flow problem is the ODE

$$
0 = \psi_{\alpha}''(r) + \frac{2n-1}{r} \psi_{\alpha}'(r) - \frac{1}{r^2} \eta(\psi_{\alpha}),
$$
 (58)

subject the initial conditions $\psi_{\alpha}(0) = 0$ and $\psi_{\alpha}'(0) = \alpha > 0$.

In the spherical coordinates the stationary solutions – that is, the harmonic maps – are given explicitly in (27) . By transforming these solutions into the coordinates g, one finds that the unique solutions to the stationary real heat flow problem are

$$
\psi_{\alpha}(r)=2\arctan(\alpha r),
$$

which may be verified by substitution into (58) . In light of later results, what will be most notable about the explicit solution is that it is independent of n .

5.1. Uniqueness of solutions to the PDE problem in the $n \geq 3$ case

PDEs of the type

$$
u_t = u_{rr} + \frac{d-1}{r}u_r - \frac{\eta(u)}{r^2},
$$
\n(59)

with

$$
\eta(0) = \eta(\pi) = \eta(2\pi), \quad \eta(x) > 0 \text{ for } x \in (0, \pi), \quad \eta(x) < 0 \text{ for } x \in (\pi, 2\pi),
$$

arise naturally in the study of the equivariant harmonic map heat flow from \mathbb{R}^d to spherically symmetric manifolds. There is a general theorem classifying when there is uniqueness of solutions and when there is not uniqueness [\[9\]](#page-39-2). It states that if

$$
\eta'(\pi) < -\frac{(d-2)^2}{4} \tag{60}
$$

then there is non-uniqueness – that is, two distinct solutions with the same initial data – while if

$$
\eta'(x) \ge -\frac{(d-2)^2}{4} \tag{61}
$$

for all x then for every initial data there is at most one solution in $L_t^{\infty} L_x^{\infty}$. We offer the following new proof of the latter case.

Proposition 5.1. Suppose that $\eta'(x) \ge -(d-2)^2/4$ for all x. There there is at most one *solution to* [\(59\)](#page-29-0) *in* $L_t^{\infty} L_x^{\infty}$ *.*

Proof. First we observe that condition [\(61\)](#page-29-1) implies the one-sided Lipschitz inequality

$$
\frac{\eta(u) - \eta(v)}{u - v} \ge \min_{x \in [0, 2\pi]} \eta'(x) \ge -\frac{(d - 2)^2}{4}.
$$

Now consider two solutions u and v of [\(59\)](#page-29-0) with the same initial data u_0 and set $\phi = u - v$. We will assume that $u_0 \in L^2 \cap L^{\infty}$; the argument to upgrade this to L^2 is standard [\[9\]](#page-39-2). Under this assumption we calculate

$$
\frac{1}{2}\frac{d}{dt}\|\phi\|_{L^2}^2 = \frac{1}{2}\frac{d}{dt}\sigma_{d-1}\int_0^\infty |\phi(r,t)|^2 r^{d-1} dr
$$

\n
$$
= \sigma_{d-1}\int_0^\infty \phi \Big[\phi_{rr} + \frac{2n-1}{r}\phi_r - \frac{\eta(u)-\eta(v)}{r^2}\Big] r^{2n-1} dr
$$

\n
$$
= -\|\phi_r\|_{L^2}^2 - \sigma_{d-1}\int_0^\infty \frac{\phi^2}{r^2} \Big[\frac{\eta(u)-\eta(v)}{u-v}\Big] r^{2n-1} dr
$$

\n
$$
\le -\|\phi_r\|_{L^2}^2 + \frac{(d-2)^2}{4}\Big\|\frac{\phi}{r}\Big\|_{L^2}^2
$$

\n
$$
\le -\|\phi_r\|_{L^2}^2 + \frac{(d-2)^2}{4}\frac{4}{(d-2)^2}\|\phi_r\|_{L^2}^2 \le 0,
$$
 (62)

where in the last line we have used Hardy's inequality with the best constant $4/(d-2)^2$. This implies that $\phi \equiv 0$, and hence that $u = v$.

In this context of the real equivariant heat flow from \mathbb{C}^n to \mathbb{CP}^n , this implies the following result (given as Theorem [1.3\(](#page-7-1)ii) in the introduction).

Proposition 5.2. Let $n \geq 3$. For a given initial data there is at most one solution to [\(57\)](#page-28-1) *in* $L_t^{\infty} L_x^{\infty}$.

Proof. Here
$$
d = 2n
$$
 and $\eta(x) = (2n - 2) \sin(x) + \sin(2x)/2$. We calculate

$$
\eta'(x) = (2n - 2)\cos(x) + 2\cos^2(x) - 1
$$

= (2n - 6)\cos(x) + 2(\cos(x) + 1)² - 3 \ge (2n - 6)(-1) + 0 - 3 = -(2n - 3),

where the last inequality holds because $n \ge 3$ and so $(2n - 6) \ge 0$. Now using the inequality $-(2n-3) \ge -(n-1)^2$ (which is equivalent to $3 \ge -(n+1)^2$) gives condition [\(61\)](#page-29-1) and hence the result. \blacksquare

5.2. The \mathbb{CP}^2 case: Breakdown of uniqueness

The $n = 2$ case is the most interesting. From the expression $\eta'(x) = 2\cos(x) + \cos(2x)$, we see that $\eta'(\pi) = -1$, which is precisely the threshold $-(d-2)^2/4 = -1$ in the conditions (60) and (61) . The condition that would imply non-uniqueness, (60) , does not hold. However we find that

$$
\eta''(\pi) = -2\sin(\pi) - 4\sin(2\pi) = 0
$$

and

$$
\eta'''(\pi) = -2\cos(\pi) - 8\cos(2\pi) = -2(-1) - 8(+1) = -6 < 0,
$$

so in fact, by the second derivative test, π is a local *maximum* of $\eta'(x)$. This means that the condition that would imply uniqueness, [\(61\)](#page-29-1), does hold either. Hence the case of the real equivariant heat flow from \mathbb{C}^2 to \mathbb{CP}^2 is a borderline case not covered by the classification theorem of [\[10\]](#page-39-3). (Plots of η in the $n = 2$ and $n = 3$ cases are given in Figure [2,](#page-31-0) which make the difference clear.)

Then the question is does uniqueness hold or not? First, we see that the proof of uniqueness presented in the last section clearly breaks down: because the derivative goes below the threshold value $-(d - 2)^2/4$, a Lipschitz inequality of the form $\eta(u) - \eta(v)$ $(u - v) \ge -(d - 2)^2/4$ cannot hold.

On the other hand, inspecting the proof in $[10]$ of non-uniqueness in the case (60) we see that it relies critically on the following fact: if condition [\(60\)](#page-29-2) holds, then the stationary solutions (that is, the harmonic maps) of the PDE problem oscillate around the fixed point π as they converge to it. In our case, the harmonic maps are given explicitly by $\psi_{\alpha}(r) = 2 \arctan(\alpha r)$ and are clearly not oscillatory, and so that proof of non-uniqueness will not hold. In fact, what is interesting is that the harmonic maps being monotonic is

Figure 2. Plots of the function $\eta'(x)$ in the case of the real equivariant heat flow from \mathbb{C}^n to \mathbb{CP}^n in the cases $n = 2$ (left) and $n = 3$ (right). For the $n = 3$ case, we easily see that η satisfies condition [\(61\)](#page-29-1) with $d = 2n$, and hence that uniqueness in $L_t^{\infty} L_x^{\infty}$ holds. For the $n = 2$ case, we see that both [\(60\)](#page-29-2) and [\(61\)](#page-29-1) do not hold, so the case does not fit into the general classification theory.

ordinarily a sign that there is uniqueness (if the uniqueness condition (61) holds, then the harmonic maps are necessarily monotonic.) However, by using an alternative method in [\[10\]](#page-39-3) we are able to show that uniqueness for the problem from \mathbb{C}^2 to \mathbb{CP}^2 does not hold. The original theorem requires some background to state, so we state a special version adapted to our setting.

Theorem ([\[10,](#page-39-3) Theorem 2.2]). *Suppose that the "equator map"* $u(r, t) \equiv \pi$ (which is a *time-independent solution of the PDE) does not minimize the energy*

$$
E(f) = \int_0^1 \left[|f'|^2 + \frac{\gamma(f)}{r^2} \right] r^{d-1} dr,
$$

where $\gamma'(x) = \eta(x)$. Then there exists a self-similar weak solution of the initial value *problem* [\(59\)](#page-29-0) *that is not constant in time and that has the same initial data as the equator map*, $u_0(r) \equiv \pi$.

Using this, we prove part (i) of Theorem [1.3.](#page-7-1)

Proposition 5.3. For the case $n = 2$ there is non-uniqueness of problem [\(57\)](#page-28-1): there are *two distinct solutions with initial data* $u_0(r) \equiv \pi$.

Proof. The key aspect of the proof is capturing the fact that in the $n = 2$ case, the condition $\eta'(x) \ge -(d-2)^2/4 = -1$ in [\(61\)](#page-29-1) is violated. If the non-uniqueness condition $\eta'(\pi) < -1$ in [\(60\)](#page-29-2) held, this would be easy. However because $\eta'(\pi) = -1$, we need to do a higherorder expansion of $\eta'(x)$ around π to show this. Once we establish that the condition $\eta'(x) \ge -1$ is violated, we follow [\[20\]](#page-39-14) and construct h based on a function which almost saturates that Hardy inequality.

Let $u = \pi$ denote the equator map and h be any function. We have

$$
E(h) - E(u) = \int_0^1 \left[|h'|^2 + \frac{\gamma(h) - \gamma(\pi)}{r^2} \right] r^{d-1} dr,\tag{63}
$$

where $\gamma'(x) = \eta(x)$. One calculates

$$
\gamma'(\pi) = \eta(\pi) = 0, \qquad \gamma''(\pi) = \eta'(\pi) = -1,
$$

$$
\gamma'''(\pi) = \eta''(\pi) = 0, \quad \gamma''''(\pi) = \eta'''(\pi) = -6.
$$

Therefore by a Taylor expansion, if we choose δ small then there exists a constant $C > 0$ such that

$$
\gamma(x) - \gamma(\pi) \le -(x - \pi)^2 - C(x - \pi)^4 \tag{64}
$$

for all $x \in [\pi - \delta, \pi + \delta]$. The constant C is positive because $\gamma^{(4)}(\pi) < 0$.

To use inequality (64) in the energy expression (63) , we need to choose h valued in $[\pi - \delta, \pi + \delta].$

Following [\[20\]](#page-39-14), we define, for any $\varepsilon > 0$, the function f_{ε} : [0, 1] $\to \mathbb{R}$ by

$$
f_{\varepsilon}(r) = \begin{cases} \varepsilon^{-1} & \text{for } 0 \le r \le \varepsilon, \\ r^{-1} & \text{for } \varepsilon \le r \le 1/2, \\ 4(1-r) & \text{for } 1/2 \le r \le 1. \end{cases} \tag{65}
$$

One verifies that $f(r)$ satisfies

$$
\int_0^1 \left| \frac{f}{r} \right|^2 r^3 \, dr \le \int_0^1 |f'|^2 r^3 \, dr \le \left(1 + \frac{B}{|\log(\varepsilon)|} \right) \int_0^1 \left| \frac{f}{r} \right|^2 r^3 \, dr \tag{66}
$$

for some $B > 0$ independent of ε . That is, f is close to saturating the Hardy inequality, which in this case has best constant $4/(d - 2)^2 = 1$. Then set

$$
h(r) = \pi - \delta \frac{f_{\varepsilon}(r)}{\|f\|_{L^{\infty}}} = \pi - \delta \frac{f_{\varepsilon}(r)}{2}.
$$

We observe that $h(r) \in [\pi - \delta, \pi + \delta]$ for all r.

We then have

$$
E(h) - E(u) = \int_0^1 \left[|h'|^2 + \frac{\gamma(h) - \gamma(\pi)}{r^2} \right] r^{d-1} dr
$$

\n
$$
\leq \int_0^1 \left[|h'|^2 + \frac{-(h - \pi)^2 - C(h - \pi)^4}{r^2} \right] r^{d-1} dr
$$

\n
$$
= \int_0^1 \left[\frac{\delta^2}{4} |f'_\varepsilon|^2 - \frac{\delta^2}{4} \left| \frac{f_\varepsilon}{r} \right|^2 - C \frac{\delta^4}{16} \left| \frac{f_\varepsilon}{r} \right|^2 |f_\varepsilon|^2 \right] r^{d-1} dr.
$$

Now using bound [\(66\)](#page-32-2) we determine that

$$
E(h) - E(u) \leq \frac{\delta^2}{4} \int_0^1 \left| \frac{f_{\varepsilon}}{r} \right|^2 \left(\frac{B}{|\log \varepsilon|} - C \frac{\delta^2}{4} |f_{\varepsilon}|^2 \right) r^3 \, dr,
$$

 \blacksquare

and by choosing ε sufficiently small we may make the right-hand side negative.

We thus determine that $E(h) < E(u)$, and hence there are two solutions.

5.3. The $n \geq 3$ case: Precise dynamics of the self-similar solutions

We finally present some results on the dynamics of the self-similar solutions in the real heat flow case when $n > 3$. The methods of analysis here are not original, and our results are based on analogous results elsewhere. Our motivation in presenting them here is to show how, in this special case, one can determine precise dynamics of the self-similar solutions; it would be very satisfactory to extend these results to the general case of the GLL equation.

We first recall the self-similar problem.

Definition 5.3. The self-similar real heat flow problem is the ODE

$$
0 = \phi''_{\beta}(r) + \left(\frac{2n-1}{r} + \frac{r}{2}\right)\phi'_{\beta}(r) - \frac{1}{r^2}\eta(\phi_{\beta})
$$

subject the initial conditions $\phi_{\beta}(0) = 0$ and ϕ_{β}' $'_{\beta}(0) = \beta > 0.$

From Proposition [5.2](#page-30-0) we know that for every $\beta > 0$ there is a unique global solution to this problem and that there exists $\phi_B(\infty) \in \mathbb{R}$ such that $\lim_{r \to \infty} \phi_B(r) = \phi_B(\infty)$.

Proposition 5.4. Let ϕ_{β} be the solution of the self-similar problem and ψ_{β} the solution *of the stationary problem.*

- (i) *We have the bound* $\phi_{\beta}(r) \leq \psi_{\beta}(r)$ *.*
- (ii) *The function* ϕ_B *is monotonically increasing and* $\phi_B(r) < \pi$ *.*
- (iii) *For fixed* $r > 0$ *, the function* $\beta \mapsto \phi_{\beta}(r)$ *is strictly increasing,* $\phi_0(r) = 0$ *, and* $\lim_{\beta \to \infty} \phi_{\beta}(r) = \pi.$
- (iv) *The function* $\beta \mapsto \phi_{\beta}(\infty)$ *is strictly increasing,* $\phi_0(\infty) = 0$ *, and* $\lim_{\beta \to \infty} \phi_{\beta}(\infty)$ $=$ π .

Figure 3. Plots of $\phi_B(r)$ for $r \in [0, 2.5]$ and $\beta = 0.25, 0.5, 1, 2, 4.5, 10, 30,$ and 100.

The content of this proposition may be seen at a glance in Figure [3.](#page-33-0) Note that in light of the non-uniqueness theorem for $n = 2$, we do not expect the same dynamics in the $n = 2$ case: in fact, we expect a self-similar profile whose asymptotic limit is π .

Lemma 5.5. Suppose that for all $r \in [0, R]$, we have $\phi_B(r) < \pi$. Then ϕ_B is increasing *on* [0, R].

Proof. Because $\beta > 0$, the solution is initially increasing. For a contradiction, let r_0 be the first critical point in [0, R]. Because ϕ_β is initially increasing, r_0 must be a local maximum. However, from the ODE we have

$$
\phi''_{\beta}(r_0) = -\Big(\frac{2n-1}{r} + \frac{r}{2}\Big)\phi'_{\beta}(r_0) + \eta(\phi_{\beta}(r_0)) = \eta(\phi_{\beta}(r_0)) > 0,
$$

where $\eta(\phi_{\beta}(r_0)) > 0$ because $\phi_{\beta}(r_0) \in (0, \pi)$. The condition ϕ_{β}'' $\int_{\beta}^{\prime\prime}(r_0) > 0$ contradicts r_0 being a maximum. Hence ϕ_B is increasing on [0, R].

Proof of Proposition [5.4\(](#page-33-1)i). Let $\varepsilon > 0$ and consider the functions ϕ_B and $\psi_{B+\varepsilon}(r)$. Define

$$
f(r) = r^2(\psi_{\beta+\varepsilon}(r) - \phi_{\beta}(r)).
$$

We will show that $f(r) \ge 0$ for all r. Letting $\varepsilon \to 0$ will then give the result.

By continuity of derivatives given by the well-posedness theory, there is an initial interval [0, δ) on which $\psi_{\beta+\varepsilon}(r) - \phi_{\beta}(r)$ is increasing, and hence, as r^2 is also increasing, the function f is increasing on this interval.

Now suppose that f has a critical point. Let r_0 be the first critical point. Because f is initially increasing, this critical point must be a local maximum. Because f is increasing on $(0, r_0)$, we have $f(r_0) > 0$.

We then calculate

$$
f''(r) = r^2(\psi''_{\beta+\varepsilon}(r) - \phi''_{\beta}(r)) + 4r(\psi'_{\beta+\varepsilon}(r) - \phi'_{\beta}(r)) + 2(\psi_{\beta+\varepsilon}(r) - \phi_{\beta}(r))
$$

=
$$
\frac{4 - (2n - 1)}{r}f' + \frac{2(2n - 1) - 6}{r^2}f + \frac{r^3}{2}\phi'_{\beta} + \eta(\psi_{\beta+\varepsilon}) - \eta(\phi_{\beta}).
$$
 (67)

First, we have the Lipschitz bound

$$
\eta(\psi_{\beta+\varepsilon}(r_0))-\eta(\phi_{\beta}(r_0))\geq -(2n-3)(\psi_{\beta+\varepsilon}(r_0)-\phi_{\beta}(r_0)),
$$

where we have used the fact that $f(r_0) = \psi_{\beta+\varepsilon}(r_0) - \phi_{\beta}(r_0) > 0$ to multiply across by $\psi_{\beta+\varepsilon}(r_0) - \phi_{\beta}(r_0).$

Second, because $f(r_0) > 0$, $\phi_\beta(r_0) < \psi_{\beta+\varepsilon}(r_0) < \pi$, and hence by the lemma, ϕ_β is increasing on [0, r_0]. Therefore ϕ' $\frac{1}{\beta}(r_0) \geq 0.$

Using both of these inequalities, as well as $f'(r_0) = 0$, in [\(67\)](#page-34-0) yields

$$
f''(r_0) \ge +\frac{2(2n-1)-6}{r_0^2} f(r_0) + 0 - \frac{2n-3}{r_0^2} f(r_0)
$$

$$
= \frac{2n-5}{r_0^2} f(r_0) > 0,
$$

which contradicts r_0 being a local maximum. Hence f has no critical points; it is increasing for all r. In particular, it is always positive, so $\phi_B(r) < \psi_{B+\varepsilon}(r)$ for all r. Taking the limit $\varepsilon \to 0$ then gives $\phi_B(r) < \psi_B(r)$.

Proof of Proposition [5.4\(](#page-33-1)ii). The previous bound gives $\phi_{\beta}(r) \leq \psi_{\beta}(r) < \pi$ for all r. Hence by the lemma, $\phi_B(r)$ is always increasing.

Proof of Proposition [5.4\(](#page-33-1)iii). Set $\alpha < \beta$. We wish to show that $\phi_{\alpha}(r) < \phi_{\beta}(r)$, which follows from a maximum principle analysis of $g(r) = r^2(\phi_\beta(r) - \phi_\alpha(r))$. The analysis is similar to the proof of item (ii). The function g is initially increasing. If r_0 denotes the first critical point, which must be a maximum, one calculates

$$
g''(r_0) = \left[\frac{4n-8}{r^2} + 1\right]g(r_0) + \frac{\phi_\beta(r_0) - \phi_\alpha(r_0)}{r^2}
$$

$$
\geq \left[\frac{4n-8}{r^2} + 1\right]g(r_0) - \frac{2n-3}{r^2}g(r_0) = \left[\frac{2n-5}{r^2} + 1\right]g(r_0) \geq 0,
$$

a contradiction. Therefore g is increasing for all r , and in particular is positive, and hence $\psi_{\beta}(r) > \psi_{\alpha}(r)$.

Proof of Proposition [5.4\(](#page-33-1)iv)*.* The proof follows from a similar maximum principle argument as in the previous proof to show that the function $h(r) = (r/(2+r))^2(\psi_{\beta}(r) \psi_{\alpha}(r)$ is increasing. One then has, for $r > 1$,

$$
\left(\frac{r}{2+r}\right)^2 (\psi_{\beta}(r) - \psi_{\alpha}(r)) \ge \frac{1}{9} (\psi_{\beta}(1) - \psi_{\alpha}(1)) > 0,
$$

and hence on taking limits $(\psi_{\beta}(\infty) - \psi_{\alpha}(\infty)) \ge (1/9)(\psi_{\beta}(1) - \psi_{\alpha}(1)) > 0$, which is what we wanted to prove.

A. Some standard results

A.1. Hardy inequalities

Theorem A.1 (Generalized radial Hardy inequality). *Suppose that* $f : \mathbb{R}^d \to \mathbb{R}$ *is radial. Then for all* $p \ge 1$ *and* $k \ge 0$ *such that* $p < d/(k + 1)$ *there holds*

$$
\left\| \frac{f}{r^{k+1}} \right\|_{L^p} \le \frac{p}{d - p(k+1)} \left\| \frac{f_r}{r^k} \right\|_{L^p}.
$$
 (68)

Proof. We suppose that f is smooth and compactly supported. The result for arbitrary f then follows from a standard density argument.

We have

$$
\frac{d}{dr}\left(\frac{f}{r^k}\right) = -k\frac{f}{r^{k+1}} + \frac{f_r}{r^k}.
$$

Multiplying this equation by $(f/r^{k+1})^{p-1}r^{d-1}$ and integrating over $[0, \infty)$ yields

$$
\int_0^\infty \frac{d}{dr} \Big(\frac{f(r)}{r^k}\Big) \Big(\frac{f}{r^{k+1}}\Big)^{p-1} r^{d-1} = -\frac{k}{s(d)} \left\| \frac{f}{r^{k+1}} \right\|_{L^p}^p + \int_0^\infty \frac{f_r}{r^k} \Big(\frac{f}{r^{k+1}}\Big)^{p-1} r^{d-1},
$$

where $s(d)$ is the measure of the unit sphere in \mathbb{R}^d . Now performing integration by parts on the term on the left we find

$$
\int_0^\infty \frac{d}{dr} \left(\frac{f}{r^k}\right) \left(\frac{f}{r^{k+1}}\right)^{p-1} r^{d-1} = -\int_0^\infty \left(\frac{f}{r^k}\right) \frac{d}{dr} \left[\left(\frac{f}{r^{k+1}}\right)^{p-1} r^{d-1}\right] dr
$$

$$
+ \left[\left(\frac{f}{r^{k+1}}\right)^p r^d\right]_{r=0}^{r=\infty} . \tag{69}
$$

The boundary term corresponding to $r = \infty$ is 0 because f is compactly supported. For the $r = 0$ term we find

$$
\lim_{r \to 0} \left(\frac{f}{r^{k+1}} \right)^p r^d = \lim_{r \to 0} f(r)^p r^{d-p(k+1)} = 0
$$

if $d - p(k + 1) > 0$. We therefore have

$$
\int_0^\infty \frac{d}{dr} \left(\frac{f}{rk}\right) \left(\frac{f}{rk+1}\right)^{p-1} r^{d-1}
$$
\n
$$
= -\int_0^\infty \left(\frac{f}{rk}\right) \frac{d}{dr} \left[f(r)^{p-1} r^{d-1-(p-1)(k+1)}\right] dr
$$
\n
$$
= -\int_0^\infty \left(\frac{f}{rk}\right) \left[(d-1-(p-1)(k+1))f(r)^{p-1} r^{d-2-(p-1)(k+1)}\right] dr
$$
\n
$$
+ (p-1)f(r)^{p-2} f_r(r) r^{d-1-(p-1)(k+1)} \left] dr
$$
\n
$$
= -\frac{(d-p(k+1)+k)}{s(d)} \left\| \frac{f}{rk+1} \right\|_{L^p}^p - (p-1) \int_0^\infty \frac{f_r}{rk} \left(\frac{f}{rk+1}\right)^{p-1} r^{d-1}.
$$

Substituting this into [\(69\)](#page-36-1) and combining terms we get

$$
(d - p(k+1)) \left\| \frac{f}{r^{k+1}} \right\|_{L^p}^p = -s(d) p \int_{\mathbb{R}^d} \frac{f_r}{r^k} \left(\frac{f}{r^{k+1}} \right)^{p-1} dx \le p \left\| \frac{f_r}{r^{k+1}} \right\|_{L^p} \left\| \frac{f}{r^{k+1}} \right\|_{L^p}^{p-1},
$$

which upon dividing through by the norm of f/r^{k+1} gives the result.

Corollary A.2. *Suppose that* $f: \mathbb{R}^d \to X$ *is radial with* $X = \mathbb{C}$ *or* $X = \mathbb{R}^m$ *. Then for all* $p \ge 1$ *and* $k \ge 0$ *such that* $p < d/(k + 1)$ *, there is a constant* $C(d, p, X)$ *such that*

$$
\left\| \frac{f}{r^{k+1}} \right\|_{L^p} \le C(d, p, X) \left\| \frac{f_r}{r^k} \right\|_{L^p}.
$$
 (70)

:

Proof. Take $X = \mathbb{C}$ and write f as $f(r) = a(r) + ib(r)$ for real-valued functions a and b. Using that $||u||_{LP} \sim ||\text{Re } u||_{LP} + ||\text{Im } u||_{LP}$, we have

$$
\left\| \frac{f}{r^{k+1}} \right\|_{L^p} \lesssim \left\| \frac{a}{r^{k+1}} \right\|_{L^p} + \left\| \frac{b}{r^{k+1}} \right\|_{L^p} \lesssim \left\| \frac{a_r}{r^k} \right\|_{L^p} + \left\| \frac{b_r}{r^k} \right\|_{L^p} \lesssim \left\| \frac{f_r}{r^k} \right\|_{L^p}
$$

A similar argument holds for $X = \mathbb{R}^m$ writing f in terms of its real-valued coordinate functions.

A.2. Local well-posedness for a class of singular ODEs

Theorem A.3. *Consider the Cauchy problem*

$$
f''(r) = A(f'(r), f(r), r) - k\left(\frac{f'(r)}{r} - \frac{f(r)}{r^2}\right) + \frac{1}{r^2}B(f(r)),
$$

(71)

$$
f(0) = 0,
$$

$$
f'(0) = \alpha \in \mathbb{C},
$$

where

- $k > 0$.
- $A(z_1, z_2, r)$ *is a smooth function with* $A(\alpha, 0, 0) = 0$ *,*
- $B(z)$ *is a smooth function such that* $|B(z)| \leq C |z|^3$ *in a neighborhood of* 0*, and* $(\partial B/\partial z)(0) = (\partial B/\partial \overline{z})(0) = 0.$

There exists $r_0 > 0$ *such that there is a unique solution among all functions* $f:[0, r_0] \to \mathbb{C}$ *satisfying*

$$
|f(r)|_{L^{\infty}([0,r_0])} + \left| \frac{f'(r) - f'(0)}{r} \right|_{L^{\infty}([0,r_0])} < \infty.
$$
 (72)

The unique solution in this space is second differentiable at $r = 0$ and satisfies $f''(0) = 0$.

Let us make two remarks on the conditions in the theorem.

- Condition [\(72\)](#page-37-2) on f is equivalent to both f and f' belonging to L^{∞} and f' satisfying a Lipschitz condition at $r = 0$.
- The assumptions on B ensure that its behavior as $r \to 0$ is non-singular; indeed, one readily verifies that, for smooth f, $B(f(r))/r^2 \rightarrow 0$ as $r \rightarrow 0$. With this formulation of the Cauchy problem the singular behavior occurs only in the term $\kappa (f'(r)/r$ $f(r)/r^2$).

The proof of the theorem involves a standard, if delicate, fixed point argument; details may be found in [\[8\]](#page-39-15).

A.3. An integration inequality

Proposition A.4. Suppose that $A'(r) + c_1 rA(r) \le c_2 r^{-k}$ for $c_1 > 0$. Then for any $r_0 > 0$, $A(r) \leq C(c_1, r_0) (A(r_0)e^{-c_1r^2/4} + c_2r^{-k+1}).$

Proof. We may write the equation as

$$
\frac{d}{dr}(e^{c_1r^2/2}A(r)) \leq \frac{c_2}{r^k}e^{c_1r^2/2},
$$

which on integration gives

$$
A(r) \le e^{c_1(r_0^2 - r^2)/2} A(1) + c_2 e^{-c_1 r^2/2} \int_{r_0}^r \frac{1}{s^k} e^{c_1 s^2/2} ds
$$

 \blacksquare

$$
= e^{c_1(r_0^2 - r^2)/2} A(1) + c_2 \frac{1}{r^{k-1}} \left(\frac{1}{r^{-k+1} e^{c_1 r^2/2}} \int_{r_0}^r \frac{1}{s^k} e^{c_1 s^2/2} ds \right).
$$

To prove the result we show that the term in the brackets is bounded independently of r . This term is clearly a continuous function of r . Moreover, we have from the condition $c_1 > 0$,

$$
\lim_{r \to \infty} r^{-k+1} e^{c_1 r^2/2} = \infty \quad \text{and} \quad \lim_{r \to \infty} \int_{r_0}^r \frac{1}{s^k} e^{c_1 s^2/2} = \infty,
$$

which means, by l'Hopital's rule, that

$$
\lim_{r \to \infty} \left(\frac{1}{r^{-k+1} e^{c_1 r^{2}/2}} \int_{r_0}^r \frac{1}{s^k} e^{c_1 s^{2}/2} ds \right)
$$
\n
$$
= \lim_{r \to \infty} \left(\frac{1}{(-k+1)r^{-k} e^{c_1 r^{2}/2} + c_1 r^{-k+2} e^{c_1 r^{2}/2}} \cdot \frac{1}{r^k} e^{c_1 r^{2}/2} \right)
$$
\n
$$
= \lim_{r \to \infty} \frac{1}{-k+1 + c_1 r^2} = 0.
$$

We thus have for all $r \in [r_0,\infty)$,

$$
\left(\frac{1}{r^{-k+1}e^{c_1r^2/2}}\int_{r_0}^r\frac{1}{s^k}e^{c_1s^2/2}\,ds\right)\leq C(r_0,c_1),
$$

which completes the proof.

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