

A novel finite element approximation of anisotropic curve shortening flow

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Abstract. We extend the DeTurck trick from the classical isotropic curve shortening flow to the anisotropic setting. Here, the anisotropic energy density is allowed to depend on space, which allows an interpretation in the context of Finsler metrics, giving rise to, for instance, geodesic curvature flow in Riemannian manifolds. Assuming that the density is strictly convex and smooth, we introduce a novel weak formulation for anisotropic curve shortening flow. We then derive an optimal H^1 -error bound for a continuous-in-time semidiscrete finite element approximation that uses piecewise linear elements. In addition, we consider some fully practical fully discrete schemes and prove their unconditional stability. Finally, we present several numerical simulations, including some convergence experiments that confirm the derived error bound, as well as applications to crystalline curvature flow and geodesic curvature flow.

1. Introduction

The aim of this paper is to introduce and analyze a novel approach to approximate the evolution of curves by anisotropic curve shortening flow. The evolution law that we consider arises as a natural gradient flow for the anisotropic, spatially inhomogeneous energy

$$\mathcal{E}(\Gamma) = \int_{\Gamma} a(z)\gamma(z, \nu) \, d\mathcal{H}^1(z) = \int_{\Gamma} a\gamma(\cdot, \nu) \, d\mathcal{H}^1 \quad (1.1)$$

for a closed curve Γ , with unit normal ν , that is contained in the domain $\Omega \subset \mathbb{R}^2$. In the above, $\gamma : \Omega \times \mathbb{R}^2 \rightarrow \mathbb{R}_{\geq 0}$ denotes the anisotropy function and $a : \Omega \rightarrow \mathbb{R}_{>0}$ is a positive weight function. In the spatially homogeneous case, that is,

$$\gamma(z, p) = \gamma_0(p) \quad \text{and} \quad a(z) = 1 \quad \forall z \in \Omega = \mathbb{R}^2, \quad (1.2)$$

the corresponding functional \mathcal{E} frequently occurs as an interfacial energy, for example, in models of crystal growth [31, 44]. Our more general setting is motivated by the work [14] of Bellettini and Paolini, who consider the gradient flow for a perimeter functional P_{ϕ}

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that is associated with a Finsler metric ϕ . Through [14, (2.5), (2.6)], it is shown that $\mathcal{E}(\Gamma) = P_\phi(\Gamma)$ if one chooses γ as the dual of ϕ and a in terms of the two-dimensional ϕ -volume. As an important special case, we mention that the choices

$$\gamma(z, p) = \sqrt{G^{-1}(z)p \cdot p} \quad \text{and} \quad a(z) = \sqrt{\det G(z)} \tag{1.3}$$

can be used to describe the length of a curve in a two-dimensional Riemannian manifold (\mathcal{M}, g) . In this case, $G(z)$ is the first fundamental form arising from a local parametrization of \mathcal{M} (cf. Example 2.2(3) below). Apart from being of geometric interest, the functional \mathcal{E} also has applications in image processing [17].

The natural gradient flow for the energy \mathcal{E} evolves a family of curves $\Gamma(t) \subset \Omega$ according to the law

$$\mathcal{V}_\gamma = \kappa_\gamma, \tag{1.4}$$

where \mathcal{V}_γ and κ_γ are the anisotropic normal velocity and the anisotropic curvature, respectively. The precise definitions of these quantities are based on a formula for the first variation of \mathcal{E} , and they will be given in Section 2. In the isotropic case, that is, when

$$\gamma(z, p) = |p| \quad \text{and} \quad a(z) = 1 \quad \forall z \in \Omega = \mathbb{R}^2, \tag{1.5}$$

we have that (1.4) is just the well-known curve shortening flow, $\mathcal{V} = \kappa$, with \mathcal{V} and κ denoting the normal velocity and the curvature of $\Gamma(t)$, respectively. For theoretical aspects of the anisotropic evolution law in (1.4), we refer to [13, 30]. Further information on (spatially homogeneous) anisotropic surface energies and the corresponding gradient flow can be found in [21, 30] and the references therein.

In this paper, we are interested in the numerical solution of (1.4) based on a parametric description of the evolving curves, that is, $\Gamma(t) = x(I, t)$ for some mapping $x : I \times [0, T] \rightarrow \Omega$. Here, most of the existing literature has focused on the spatially homogeneous case given by (1.2). Then, the law in (1.4) reduces to $\frac{1}{\gamma_0(v)} \mathcal{V} = \kappa_{\gamma_0}$ (see Section 2 for details), which can be viewed as a special case of the weighted anisotropic curvature flow

$$\widehat{\beta}_0(v) \mathcal{V} = \kappa_{\gamma_0}, \tag{1.6}$$

for some mobility function $\widehat{\beta}_0$; see, for example, [21, (8.20)]. In [24], a finite element scheme is proposed and analyzed for (1.6) with $\widehat{\beta}_0 = 1$. The method uses a variational formulation of the parabolic system

$$x_t = \kappa_{\gamma_0} \nu, \tag{1.7}$$

and is generalized to higher codimension in [38]. A drawback of this approach is that the above system is degenerate in the tangential direction, so that the numerical analysis requires an additional equation for the length element. A way to circumvent this difficulty consists in replacing (1.7) by a strictly parabolic system with the help of a suitable tangential motion, known as DeTurck’s trick in the literature. For the isotropic case,

corresponding schemes have been suggested and analyzed in [19, 26]. We mention that alternative parametric approaches for (1.6) allow for some benign tangential motion; see, for example, [5, 8, 34, 36].

Since the choice of (1.3) allows to describe the length of a curve in a Riemannian manifold (\mathcal{M}, g) , it is possible to use (1.4) in order to treat geodesic curvature flow in \mathcal{M} within our framework. Existing parametric approaches for this flow on a hypersurface of \mathbb{R}^3 include the work [35] for the flow on a graph, as well as [6] for the case that the hypersurface is given as a level set. Moreover, numerical schemes for the geodesic curvature flow (and other flows) of curves in a locally flat two-dimensional Riemannian manifold have been proposed in [12]. To the best of our knowledge, no error bounds have been derived for a numerical approximation of geodesic curvature flow in Riemannian manifolds in the literature so far.

In this paper, we propose and analyze a new method for solving (1.4), which is based on DeTurck’s trick and which applies to general, spatially inhomogeneous anisotropies. Let us outline the contents of the paper and describe the main results of our work. By taking advantage of the fact that the function $(z, p) \mapsto \frac{1}{2}a^2(z)\gamma^2(z, p)$ is strictly convex in p , we derive in Section 3 a strictly parabolic system whose solution satisfies (1.4). It turns out that this system can be written in a variational form, which makes it accessible to discretization by linear finite elements. In the isotropic case, the resulting numerical scheme is precisely the method proposed and analyzed in [19], while in the anisotropic case we obtain a novel scheme that can be considered as a generalization of the ideas in [19, 26]. As one of the main results of this paper, we show in Section 4 an optimal H^1 -error bound in the continuous-in-time semidiscrete case. Unlike in [24, 38], the corresponding proof does not need an equation for the length element because of the strict parabolicity of the underlying partial differential equation. In order to discretize in time, we use the backward Euler method. In particular, in Section 5, as another important contribution of our work, we introduce unconditionally stable fully discrete finite element approximations for the following scenarios:

- (a) a spatially homogeneous, smooth anisotropy function $\gamma(z, p) = \gamma_0(p)$;
- (b) a spatially homogeneous anisotropy function

$$\gamma(z, p) = \sum_{\ell=1}^L \sqrt{\Lambda_\ell p \cdot p},$$

where Λ_ℓ are symmetric and positive definite matrices;

- (c) a spatially inhomogeneous anisotropy function in the form of (1.3) to model geodesic curvature flow in a two-dimensional Riemannian manifold.

In particular, the functions in scenario (b) can be used to approximate the case of a crystal-line anisotropy (cf. [5]). Using these three fully discrete schemes, we present in Section 6

results of test calculations that confirm our error bound and show that the tangential motion that is introduced in our approach has a positive effect on the distribution of grid points along the discrete curve.

As our approach is based on a parametrization of the evolving curves, we only briefly mention numerical methods that employ an implicit description such as the level-set method or the phase field approach. The interested reader may consult [10, 37] for anisotropic curve shortening flow, as well as [15, 16, 43] for the geodesic curvature flow. These papers also provided additional references.

We end this section with a few comments about notation. Throughout the paper, we let $I = \mathbb{R}/\mathbb{Z}$ denote the periodic interval $[0, 1]$. We adopt the standard notation for Sobolev spaces, denoting the norm of $W^{\ell,p}(I)$ ($\ell \in \mathbb{N}_0, p \in [1, \infty]$) by $\|\cdot\|_{\ell,p}$ and the seminorm by $|\cdot|_{\ell,p}$. For $p = 2$, $W^{\ell,2}(I)$ will be denoted by $H^\ell(I)$ with the associated norm and seminorm written as $\|\cdot\|_\ell$ and $|\cdot|_\ell$, respectively. The above are naturally extended to vector functions, and we will write $[W^{\ell,p}(I)]^2$ for a vector function with two components. For later use, we recall the well-known Sobolev embedding $H^1(I) \hookrightarrow C^0(I)$, that is, there exists $C_I > 0$ such that

$$\|f\|_{0,\infty} \leq C_I \|f\|_1, \quad \forall f \in H^1(I). \tag{1.8}$$

Furthermore, throughout the paper C will denote a generic positive constant independent of the mesh parameter h , which will be introduced in Section 4. At times, ε will play the role of a (small) positive parameter, with $C_\varepsilon > 0$ depending on ε , but independent of h . Finally, in this paper we make use of the Einstein summation convention.

2. Anisotropy and anisotropic curve shortening flow

Let $\Omega \subset \mathbb{R}^2$ be a domain, and let $a \in C^2(\Omega, \mathbb{R}_{>0})$. Moreover, we assume that $\gamma \in C^0(\Omega \times \mathbb{R}^2, \mathbb{R}_{\geq 0}) \cap C^3(\Omega \times (\mathbb{R}^2 \setminus \{0\}), \mathbb{R}_{>0})$, as well as

$$\gamma(z, \lambda p) = \lambda \gamma(z, p), \quad \forall z \in \Omega, p \in \mathbb{R}^2, \lambda \in \mathbb{R}_{>0}, \tag{2.1}$$

which means that γ is positively one-homogeneous with respect to the second variable. It is not difficult to verify that (2.1) implies that

$$\begin{aligned} \gamma_p(z, \lambda p) &= \gamma_p(z, p), \quad \gamma_p(z, p) \cdot p = \gamma(z, p), \\ \text{and } \gamma_{pp}(z, p)p &= 0, \quad \forall z \in \Omega, p \in \mathbb{R}^2 \setminus \{0\}, \lambda \in \mathbb{R}_{>0}. \end{aligned} \tag{2.2}$$

Here, $\gamma_p = (\gamma_{p_j})_{j=1}^2$ and $\gamma_{pp} = (\gamma_{p_i p_j})_{i,j=1}^2$ denote the first and second derivatives of γ with respect to the second argument. Similarly, we let $\gamma_z = (\gamma_{z_j})_{j=1}^2$ denote the derivatives of γ with respect to the first argument. We note for later use that on differentiating (2.2) with respect to z , we immediately obtain that the functions $\gamma_{z_j}(z, \cdot)$ and $\gamma_{p z_j}(z, \cdot)$ are positively one- and zero-homogeneous, respectively, for every $z \in \Omega$. In addition, we

assume that $p \mapsto \gamma(z, p)$ is strictly convex for every $z \in \Omega$ in the sense that

$$\gamma_{pp}(z, p)q \cdot q > 0, \quad \forall z \in \Omega, \quad p, q \in \mathbb{R}^2 \text{ with } |p| = |q| = 1, \quad p \cdot q = 0. \quad (2.3)$$

We are now in a position to define anisotropic curve shortening flow. To this end, with the help of [22, Corollary 4.3], we first state the first variation of the functional \mathcal{E} in (1.1), the proof of which will be given in Appendix A.

Lemma 2.1. *Let $\Gamma \subset \Omega$ be a smooth curve with unit normal ν , unit tangent τ , and scalar curvature κ . Let V be a smooth vector field defined in an open neighborhood of Γ . Then, the first variation of \mathcal{E} at Γ in the direction V is given by*

$$d\mathcal{E}(\Gamma; V) = - \int_{\Gamma} \kappa_{\gamma} V \cdot \nu_{\gamma} a \gamma(\cdot, \nu) \, d\mathcal{H}^1, \quad (2.4)$$

where

$$\nu_{\gamma} = \frac{\nu}{\gamma(\cdot, \nu)} \quad \text{and} \quad \kappa_{\gamma} = \kappa \gamma_{pp}(\cdot, \nu) \tau \cdot \tau - \gamma_{p_i z_i}(\cdot, \nu) - \frac{\nabla a}{a} \cdot \gamma_p(\cdot, \nu) \quad \text{on } \Gamma \quad (2.5)$$

denote the anisotropic normal and the anisotropic curvature of Γ , respectively.

We remark that the definitions in (2.5) correspond to [14, (3.5) and (4.1)]. Note also that ν_{γ} is a vector that is normal to Γ , but normalized in such a way that $\gamma(z, \nu_{\gamma}(z)) = 1$, $z \in \Gamma$. We remark that although κ_{γ} clearly depends on both γ and a , we prefer to use the simpler notation that drops the dependence on a .

Following [14, (1.1)], we now consider a natural gradient flow induced by (2.4). In particular, given a family of curves $(\Gamma(t))_{t \in [0, T]}$ in Ω , we say that $\Gamma(t)$ evolves according to anisotropic curve shortening flow provided that

$$\mathcal{V}_{\gamma} = \kappa_{\gamma} \quad \text{on } \Gamma(t), \quad (2.6)$$

where $\mathcal{V}_{\gamma} = (\mathcal{V}\nu) \cdot \nu_{\gamma} = \frac{1}{\gamma(\cdot, \nu)} \mathcal{V}$, with \mathcal{V} denoting the normal velocity of $\Gamma(t)$, and where κ_{γ} is defined in (2.5). We remark that the name of the flow is inspired by the fact that solutions of (2.6) satisfy the energy relation

$$\frac{d}{dt} \int_{\Gamma(t)} a \gamma(\cdot, \nu) \, d\mathcal{H}^1 + \int_{\Gamma(t)} |\mathcal{V}_{\gamma}|^2 a \gamma(\cdot, \nu) \, d\mathcal{H}^1 = 0. \quad (2.7)$$

We note that the higher-dimensional analogue of (2.6) is usually called anisotropic mean curvature flow or anisotropic motion by mean curvature. Hence, alternative names for evolution law (2.6) in the planar case treated in this paper are anisotropic curvature flow and anisotropic motion by curvature.

Example 2.2. The following three cases fall within the framework of this paper:

- (1) *Isotropic case:* We let $\gamma(z, p) = |p|$ and $a(z) = 1$ for all $z \in \Omega = \mathbb{R}^2$ (recall (1.5)) so that $\mathcal{E}(\Gamma)$ is the length of Γ . In this case, (2.6) is just the well-known curve shortening flow

$$\mathcal{V} = \kappa \quad \text{on } \Gamma(t).$$

- (2) *Space-independent anisotropy:* We let $\gamma(z, p) = \gamma_0(p)$ and $a(z) = 1$ for all $z \in \Omega = \mathbb{R}^2$ (recall (1.2)) so that $\mathcal{E}(\Gamma) = \int_{\Gamma} \gamma_0(v) \, d\mathcal{H}^1$ is the associated anisotropic length. Then, (2.6) reduces to

$$\frac{1}{\gamma_0(v)} \mathcal{V} = \kappa_{\gamma_0} = \kappa \gamma_0''(v) \tau \cdot \tau \quad \text{on } \Gamma(t) \tag{2.8}$$

where, here and throughout, γ_0' and γ_0'' denote the gradient and Hessian of γ_0 , respectively. We observe that (2.8) corresponds to [21, (8.20)] with $\beta(v) = \frac{1}{\gamma_0(v)}$; see also [39] for a nice derivation of this law. Of course, for $\gamma_0(p) = |p|$, we obtain the isotropic case described in Example 2.2(1) above.

- (3) *Riemannian manifolds:* Suppose that (\mathcal{M}, g) is a two-dimensional Riemannian manifold. Let $F : \Omega \rightarrow \mathcal{M}$ be a local parametrization of \mathcal{M} , $\{\partial_1, \partial_2\}$ be the corresponding basis of the tangent space $T_{F(z)}\mathcal{M}$, and $g_{ij}(z) = g_{F(z)}(\partial_i, \partial_j)$, $z \in \Omega$. Also, let $G(z) = (g_{ij}(z))_{i,j=1}^2$. We set $\gamma(z, p) = \sqrt{G^{-1}(z)p \cdot p}$ and $a(z) = \sqrt{\det G(z)}$. Then, we have

$$a(z)\gamma(z, p) = \sqrt{\det G(z)G^{-1}(z)p \cdot p} = \sqrt{G(z)p^\perp \cdot p^\perp}, \tag{2.9}$$

where $p^\perp = \begin{pmatrix} p_1 \\ p_2 \end{pmatrix}^\perp = \begin{pmatrix} -p_2 \\ p_1 \end{pmatrix}$ denotes an anti-clockwise rotation of p by $\frac{\pi}{2}$. For a curve $\Gamma \subset \Omega$, the vector $\tau = -v^\perp$ is then a unit tangent and

$$\mathcal{E}(\Gamma) = \int_{\Gamma} a\gamma(\cdot, v) \, d\mathcal{H}^1 = \int_{\Gamma} \sqrt{G\tau \cdot \tau} \, d\mathcal{H}^1 \tag{2.10}$$

is the Riemannian length of the curve $\tilde{\Gamma} = F(\Gamma) \subset \mathcal{M}$. We show in Appendix B that the geodesic curvature of $\tilde{\Gamma}$ at $F(z)$ is equal to κ_γ at $z \in \Gamma$, and also that $(\Gamma(t))_{t \in [0, T]} \subset \Omega$ is a solution of (2.6) if and only if $\tilde{\Gamma}(t) = F(\Gamma(t))$ evolves according to geodesic curvature flow in \mathcal{M} .

3. DeTurck’s trick for anisotropic curve shortening flow

In what follows, we shall employ a parametric description of the evolving curves. Let $\Gamma(t) = x(I, t)$, where $x : I \times [0, T] \rightarrow \mathbb{R}^2$ and $I = \mathbb{R}/\mathbb{Z}$. In order to satisfy (2.6), we require that

$$\frac{1}{\gamma(x, v \circ x)} x_t \cdot (v \circ x) = \kappa_\gamma \circ x \quad \text{in } I \times (0, T]. \tag{3.1}$$

From now on, we fix a normal on $\Gamma(t)$ induced by the parametrization x , and since no confusion can arise, we identify $v \circ x$ with v , $\kappa_\gamma \circ x$ with κ_γ , and similarly, $x \circ x$ with x . In particular, we define the unit tangent, the unit normal, and the curvature of $\Gamma(t)$ by

$$\tau = \frac{x_\rho}{|x_\rho|}, \quad \nu = \tau^\perp, \quad \kappa = \frac{1}{|x_\rho|} \left(\frac{x_\rho}{|x_\rho|} \right)_\rho \cdot \nu = \frac{x_{\rho\rho}}{|x_\rho|^2} \cdot \nu. \tag{3.2}$$

In place of (3.1), we simply write

$$\frac{1}{\gamma(x, \nu)} x_t \cdot \nu = \kappa_\gamma \quad \text{in } I \times (0, T]. \tag{3.3}$$

Clearly, (3.3) only prescribes the normal component of the velocity vector x_t , and so there is a certain freedom in the tangential direction. Our aim is to introduce a strictly parabolic system of partial differential equations for the parametrization x , whose solution in the normal direction still satisfies (3.3).

Let us briefly review the DeTurck trick in the isotropic setting; recall (1.5) and Example 2.2(1). Then, (3.3) collapses to $x_t \cdot \nu = \kappa$, and adjoining a zero tangential velocity leads to the formulation $x_t = \kappa \nu = \frac{1}{|x_\rho|} (\frac{x_\rho}{|x_\rho|})_\rho$ as the isotropic equivalent to (1.7). We recall that optimal error bounds for a semidiscrete continuous-in-time finite element approximation of this formulation have been obtained in the seminal paper [23] by Gerd Dziuk. One difficulty of Dziuk’s original approach is that the analyzed system is degenerate in the tangential direction. DeTurck’s trick addresses this problem by removing the degeneracy through a suitable reparametrization. In fact, it is natural to consider the system

$$x_t = \frac{x_{\rho\rho}}{|x_\rho|^2} \tag{3.4}$$

(recall (3.2)) for which a semidiscretization by linear finite elements was analyzed in [19]. The appeal of this approach is that the analysis is very elegant and simple. For example, the weak formulation of (3.4) is given by

$$\int_I |x_\rho|^2 x_t \cdot \eta \, d\rho + \int_I x_\rho \cdot \eta_\rho \, d\rho = 0, \quad \forall \eta \in [H^1(I)]^2, \tag{3.5}$$

and choosing $\eta = x_t$ immediately gives rise to the estimate

$$\frac{1}{2} \frac{d}{dt} \int_I |x_\rho|^2 \, d\rho = \int_I x_\rho \cdot x_{t\rho} \, d\rho = - \int_I |x_\rho|^2 |x_t|^2 \, d\rho \leq 0, \tag{3.6}$$

which can be mimicked on the discrete level.

Our starting point for extending DeTurck’s trick to the anisotropic setting is to define the function $\Phi : \Omega \times \mathbb{R}^2 \rightarrow \mathbb{R}_{\geq 0}$ by setting

$$\Phi(z, p) = \frac{1}{2} a^2(z) \gamma^2(z, p^\perp), \quad \forall z \in \Omega, \, p \in \mathbb{R}^2. \tag{3.7}$$

We mention that the square of the anisotropy function plays an important role in the phase field approach to anisotropic mean curvature flow (cf.[1, 10, 27]).

On noting (3.2), (2.1), and (3.7) we compute, in a similar manner to the calculations in (3.6), that

$$\frac{1}{2} \frac{d}{dt} \int_I a^2(x) \gamma^2(x, \nu) |x_\rho|^2 \, d\rho = \frac{d}{dt} \int_I \Phi(x, x_\rho) \, d\rho$$

$$\begin{aligned} &= \int_I \Phi_p(x, x_\rho) \cdot x_{t\rho} + \Phi_z(x, x_\rho) \cdot x_t \, d\rho \\ &= - \int_I ([\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho)) \cdot x_t \, d\rho. \end{aligned} \tag{3.8}$$

The crucial idea is now to define positive definite matrices $H(z, p) \in \mathbb{R}^{2 \times 2}$, for $(z, p) \in \Omega \times (\mathbb{R}^2 \setminus \{0\})$, such that if a sufficiently smooth x satisfies

$$H(x, x_\rho)x_t = [\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho) \quad \text{in } I \times (0, T], \tag{3.9}$$

then x is a solution to anisotropic curve shortening flow given by (3.3). For the construction of the matrices H , it is important to relate the right-hand side in (3.9) to the right-hand side in (3.3). We begin by calculating

$$\begin{aligned} \Phi_p(z, p) &= -a^2(z)\gamma(z, p^\perp)\gamma_p^\perp(z, p^\perp), \quad \forall z \in \Omega, p \in \mathbb{R}^2 \setminus \{0\}; \\ \Phi_z(z, p) &= a^2(z)\gamma(z, p^\perp)\gamma_z(z, p^\perp) + a(z)\gamma^2(z, p^\perp)\nabla a(z), \quad \forall z \in \Omega, p \in \mathbb{R}^2, \end{aligned}$$

where we use the notation $\gamma_p^\perp(z, p) = [\gamma_p(z, p)]^\perp$. Furthermore, we obtain with the help of (2.1), (2.2), and (3.2) that

$$\begin{aligned} [\Phi_p(x, x_\rho)]_\rho &= -[a(x)\gamma(x, x_\rho^\perp)a(x)\gamma_p^\perp(x, x_\rho^\perp)]_\rho \\ &= -[a(x)\gamma(x, x_\rho^\perp)]_\rho a(x)\gamma_p^\perp(x, x_\rho^\perp) - a^2(x)\gamma(x, x_\rho^\perp)x_{j,\rho}\gamma_{pz_j}^\perp(x, x_\rho^\perp) \\ &\quad - a(x)\gamma(x, x_\rho^\perp)\nabla a(x) \cdot x_\rho\gamma_p^\perp(x, x_\rho^\perp) - a^2(x)\gamma(x, x_\rho^\perp)(\gamma_{pp}(x, x_\rho^\perp)x_{\rho\rho}^\perp)^\perp \\ &= -[a(x)\gamma(x, x_\rho^\perp)]_\rho a(x)\gamma_p^\perp(x, v) + a^2(x)\gamma(x, v)|x_\rho|^2(v_1\gamma_{pz_2}^\perp(x, v) - v_2\gamma_{pz_1}^\perp(x, v)) \\ &\quad - a(x)|x_\rho|\gamma(x, v)\nabla a(x) \cdot x_\rho\gamma_p^\perp(x, v) + a^2(x)|x_\rho|^2\gamma(x, v)\varkappa(\gamma_{pp}(x, v)\tau \cdot \tau)v. \end{aligned}$$

Similarly,

$$\Phi_z(x, x_\rho) = a^2(x)|x_\rho|^2\gamma(x, v)\gamma_z(x, v) + a(x)|x_\rho|^2\gamma^2(x, v)\nabla a(x),$$

and therefore,

$$\begin{aligned} &[\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho) \\ &= -[a(x)\gamma(x, x_\rho^\perp)]_\rho a(x)\gamma_p^\perp(x, v) + a^2(x)|x_\rho|^2\gamma(x, v)\varkappa(\gamma_{pp}(x, v)\tau \cdot \tau)v \\ &\quad + a^2(x)\gamma(x, v)|x_\rho|^2(v_1\gamma_{pz_2}^\perp(x, v) - v_2\gamma_{pz_1}^\perp(x, v) - \gamma_z(x, v)) \\ &\quad - a(x)|x_\rho|^2\gamma(x, v)(\nabla a(x) \cdot \tau\gamma_p^\perp(x, v) + \gamma(x, v)\nabla a(x)). \end{aligned} \tag{3.10}$$

Observing that $\gamma_{z_i}(x, v) = \gamma_{pz_i}(x, v) \cdot v$ (recall (2.2)), we find

$$\begin{aligned} [v_1\gamma_{pz_2}^\perp - v_2\gamma_{pz_1}^\perp - \gamma_z]1 &= -v_1\gamma_{pz_2} + v_2\gamma_{pz_1} - v_1\gamma_{pz_1} - v_2\gamma_{pz_2} \\ &= -(\gamma_{pz_1} + \gamma_{pz_2})v_1, \end{aligned}$$

and a similar argument for the second component yields

$$v_1 \gamma_{pz_2}^\perp(x, v) - v_2 \gamma_{pz_1}^\perp(x, v) - \gamma_z(x, v) = -\gamma_{p_i z_i}(x, v)v. \tag{3.11}$$

Next, since

$$\gamma_p^\perp(x, v) \cdot \tau = -\gamma_p(x, v) \cdot v = -\gamma(x, v) \quad \text{and} \quad \gamma_p^\perp(x, v) \cdot v = \gamma_p(x, v) \cdot \tau, \tag{3.12}$$

we derive

$$\begin{aligned} \nabla a \cdot \tau \gamma_p^\perp + \gamma \nabla a &= (\nabla a \cdot \tau) ((\gamma_p^\perp \cdot \tau)\tau + (\gamma_p^\perp \cdot v)v) + \gamma((\nabla a \cdot \tau)\tau + (\nabla a \cdot v)v) \\ &= (\nabla a \cdot \tau)(\gamma_p \cdot \tau)v + (\nabla a \cdot v)(\gamma_p \cdot v)v = (\nabla a \cdot \gamma_p)v. \end{aligned} \tag{3.13}$$

If we insert (3.11) and (3.13) into (3.10), recall (2.5), and use the abbreviation $\omega(x) = [a(x)\gamma(x, x_\rho^\perp)]_\rho$, we obtain

$$[\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho) = -\omega(x)a(x)\gamma_p^\perp(x, v) + a^2(x)|x_\rho|^2\gamma(x, v)x_\gamma v. \tag{3.14}$$

Let us now assume that x is a solution of (3.9), where $H(z, p)$ is an invertible matrix of the form

$$H(z, p) = \begin{pmatrix} \alpha(z, p) & -\beta(z, p) \\ \beta(z, p) & \alpha(z, p) \end{pmatrix}, \quad \forall z \in \Omega, p \in \mathbb{R}^2 \setminus \{0\}.$$

In order to determine $\alpha(z, p), \beta(z, p) \in \mathbb{R}$ such that x satisfies (3.3), we calculate

$$\begin{aligned} x_t \cdot v &= H^{-1}(x, x_\rho)([\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho)) \cdot v \\ &= \frac{1}{\alpha^2(x, x_\rho) + \beta^2(x, x_\rho)}([\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho)) \cdot H(x, x_\rho)v. \end{aligned}$$

If we multiply by $\alpha^2(x, x_\rho) + \beta^2(x, x_\rho)$ and insert (3.14) we obtain, on noting $\binom{-v_2}{v_1} = v^\perp = -\tau$ and (3.12), that

$$\begin{aligned} &(\alpha^2(x, x_\rho) + \beta^2(x, x_\rho))x_t \cdot v \\ &= (-\omega(x)a(x)\gamma_p^\perp(x, v) + a^2(x)|x_\rho|^2\gamma(x, v)x_\gamma v) \cdot (\alpha(x, x_\rho)v - \beta(x, x_\rho)\tau) \\ &= \omega(x)a(x)(-\alpha(x, x_\rho)\gamma_p(x, v) \cdot \tau - \beta(x, x_\rho)\gamma(x, v)) \\ &\quad + \alpha(x, x_\rho)a^2(x)|x_\rho|^2\gamma(x, v)x_\gamma \\ &= \alpha(x, x_\rho)a^2(x)|x_\rho|^2\gamma(x, v)x_\gamma, \end{aligned}$$

provided that $\alpha(x, x_\rho)\gamma_p(x, v) \cdot \tau + \beta(x, x_\rho)\gamma(x, v) = 0$. With this choice, we obtain that

$$\beta^2(x, x_\rho) = \alpha^2(x, x_\rho) \frac{(\gamma_p(x, v) \cdot \tau)^2}{\gamma^2(x, v)},$$

and so

$$\frac{1}{\gamma(x, v)}x_t \cdot v = \frac{\alpha(x, x_\rho)}{\alpha^2(x, x_\rho) + \beta^2(x, x_\rho)}a^2(x)|x_\rho|^2x_\gamma$$

$$\begin{aligned} &= \frac{1}{\alpha(x, x_\rho)} \frac{\gamma^2(x, v)}{\gamma^2(x, v) + (\gamma_p(x, v) \cdot \tau)^2} |x_\rho|^2 a^2(x) \kappa_\gamma \\ &= \frac{1}{\alpha(x, x_\rho)} \frac{\gamma^2(x, |x_\rho|v)}{|\gamma_p(x, v)|^2} a^2(x) \kappa_\gamma, \end{aligned}$$

where in the last step we have used the one-homogeneity of γ (recall (2.2)). Clearly, (3.3) will now hold if we choose

$$\begin{aligned} \alpha(z, p) &= \frac{a^2(z)\gamma^2(z, p^\perp)}{|\gamma_p(z, p^\perp)|^2}, \\ \beta(z, p) &= -\alpha(z, p) \frac{\gamma_p(z, p^\perp) \cdot p}{\gamma(z, p^\perp)} = -\frac{a^2(z)\gamma(z, p^\perp)\gamma_p(z, p^\perp) \cdot p}{|\gamma_p(z, p^\perp)|^2}. \end{aligned}$$

In summary, we have shown the following result:

Lemma 3.1. *Let $\Phi(z, p) = \frac{1}{2}a^2(z)\gamma^2(z, p^\perp)$ and*

$$H(z, p) = \frac{a^2(z)\gamma(z, p^\perp)}{|\gamma_p(z, p^\perp)|^2} \begin{pmatrix} \gamma(z, p^\perp) & \gamma_p(z, p^\perp) \cdot p \\ -\gamma_p(z, p^\perp) \cdot p & \gamma(z, p^\perp) \end{pmatrix}, \quad \forall z \in \Omega, p \in \mathbb{R}^2 \setminus \{0\}. \tag{3.15}$$

If $x : I \times [0, T] \rightarrow \Omega$ satisfies (3.9), then x is a solution to anisotropic curve shortening flow given in (3.3). In addition, H is positive definite in $\Omega \times (\mathbb{R}^2 \setminus \{0\})$ with

$$H(z, p)w \cdot w = \frac{a^2(z)\gamma^2(z, p^\perp)}{|\gamma_p(z, p^\perp)|^2} |w|^2, \quad \forall z \in \Omega, p \in \mathbb{R}^2 \setminus \{0\}, w \in \mathbb{R}^2. \tag{3.16}$$

Furthermore, it can be shown that system (3.9) is strictly parabolic. The proof hinges on the fact that H and Φ_{pp} are positive definite matrices in $\Omega \times (\mathbb{R}^2 \setminus \{0\})$. This property of Φ_{pp} immediately follows from our convexity assumptions on γ (recall (2.3)).

Lemma 3.2. *Let $K \subset \Omega \times (\mathbb{R}^2 \setminus \{0\})$ be compact. Then, there exists $\sigma_K > 0$ such that*

$$\Phi_{pp}(z, p)w \cdot w \geq \sigma_K |w|^2, \quad \forall (z, p) \in K, w \in \mathbb{R}^2. \tag{3.17}$$

Furthermore,

$$\begin{aligned} \Phi(z, q) - \Phi(z, p) - \Phi_p(z, p) \cdot (q - p) &\geq \frac{1}{2}\sigma_K |q - p|^2, \\ \forall (z, p), (z, q) \in K \quad \text{with} \quad \{z\} \times [p, q] &\subset K. \end{aligned} \tag{3.18}$$

Here, $[p, q] \subset \mathbb{R}^2$ denotes the line segment connecting p and q .

Proof. It is shown in [30, Remark 1.7.5] that (2.3) implies that $\Phi_{pp}(z, p)$ is positive definite for all $z \in \Omega$ and $p \neq 0$. Bound (3.17) then follows with the help of a compactness argument, while the elementary identity

$$\begin{aligned} &\Phi(z, q) - \Phi(z, p) - \Phi_p(z, p) \cdot (q - p) \\ &= \int_0^1 (\Phi_p(z, sq + (1-s)p) - \Phi_p(z, p)) \cdot (q - p) \, ds \end{aligned}$$

$$= \int_0^1 \int_0^1 s \Phi_{pp}(z, \theta sq + (1 - \theta s)p)(q - p) \cdot (q - p) \, d\theta \, ds,$$

together with (3.17), implies (3.18). ■

Lemma 3.3. *The system in (3.9) is parabolic in the sense of Petrovsky.*

Proof. On inverting the matrix $H(x, x_\rho)$, we may write (3.9) in the form

$$\begin{aligned} x_t &= H^{-1}(x, x_\rho) \Phi_{pp}(x, x_\rho) x_{\rho\rho} \\ &\quad + H^{-1}(x, x_\rho) (\Phi_{pz}(x, x_\rho) x_\rho - \Phi_z(x, x_\rho)) \quad \text{in } I \times (0, T]. \end{aligned}$$

Hence, by definition we need to show that the eigenvalues of $H^{-1}(z, p) \Phi_{pp}(z, p)$ have positive real parts for every $(z, p) \in \Omega \times (\mathbb{R}^2 \setminus \{0\})$; see, for example, [25, Definition 1.2]. Let us fix (z, p) and abbreviate $H = H(z, p)$, $A = \Phi_{pp}(z, p)$. The two eigenvalues $\lambda_1, \lambda_2 \in \mathbb{C}$ of $H^{-1}A \in \mathbb{R}^{2 \times 2}$ satisfy

$$\lambda_1 \lambda_2 = \det(H^{-1}A) = \frac{\det A}{\det H} > 0, \quad \lambda_1 + \lambda_2 = \text{tr}(H^{-1}A) = \frac{\text{tr}(H^T A)}{\det H} = \frac{H_{11} \text{tr } A}{\det H} > 0,$$

since $H_{11} > 0$ and $\det H > 0$ (recall (3.15)) and since A is symmetric positive definite in view of (3.17). Hence, either both eigenvalues are positive real numbers, or $\lambda_2 = \overline{\lambda_1}$ with $2\text{Re}\lambda_1 = \lambda_1 + \lambda_2 > 0$. ■

In view of Lemma 3.3, we expect that it is possible to prove the short-time existence of a unique smooth solution to (3.9). Moreover, existence and uniqueness of classical smooth solutions to PDEs of the form $x_t = \mathfrak{b}(x, \nu)v + a\tau$, arising from closely related curvature-driven geometric evolution equations, have been obtained in [34].

The weak formulation of (3.9) now reads as follows: Given $x_0 : I \rightarrow \Omega$, find $x : I \times [0, T] \rightarrow \Omega$ such that $x(\cdot, 0) = x_0$ and, for $t \in (0, T]$,

$$\begin{aligned} \int_I H(x, x_\rho) x_t \cdot \eta \, d\rho + \int_I \Phi_p(x, x_\rho) \cdot \eta_\rho \, d\rho \\ + \int_I \Phi_z(x, x_\rho) \cdot \eta \, d\rho = 0, \quad \forall \eta \in [H^1(I)]^2. \end{aligned} \tag{3.19}$$

Choosing $\eta = x_t$ in (3.19) yields, on recalling (3.8) and (3.16), that

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_I a^2(x) \gamma^2(x, \nu) |x_\rho|^2 \, d\rho &= \frac{d}{dt} \int_I \Phi(x, x_\rho) \, d\rho \\ &= \int_I (\Phi_p(x, x_\rho) \cdot x_{t\rho} + \Phi_z(x, x_\rho) \cdot x_t) \, d\rho \\ &= - \int_I H(x, x_\rho) x_t \cdot x_t \, d\rho \leq 0. \end{aligned} \tag{3.20}$$

Clearly, (3.20) is the desired anisotropic analogue to (3.6). So, together with the fact that x is a solution of the gradient flow given in (3.3) (recall also (2.7)), we obtain that both $\frac{1}{2} \int_I a^2(x) \gamma^2(x, \nu) |x_\rho|^2 \, d\rho$ and $\int_I a(x) \gamma(x, \nu) |x_\rho| \, d\rho$ are monotonically decreasing in time.

Example 3.4. For the following cases, we refer to the same numbering as in Example 2.2.

(1) *Isotropic case:* We have $\Phi(z, p) = \frac{1}{2}|p|^2$ and $H(z, p) = |p|^2 \text{Id}$, so that (3.19) collapses to (3.5), which is the same as [19, (12)].

(2) *Space-independent anisotropy:* We have $\Phi(z, p) = \Phi_0(p) = \frac{1}{2}\gamma_0^2(p^\perp)$, so that

$$\Phi_p(z, p) = \Phi'_0(p) = -\gamma_0(p^\perp)[\gamma'_0(p^\perp)]^\perp \tag{3.21a}$$

$$\text{and } H(z, p) = H_0(p) = \frac{\gamma_0(p^\perp)}{|\gamma'_0(p^\perp)|^2} \begin{pmatrix} \gamma_0(p^\perp) & \gamma'_0(p^\perp) \cdot p \\ -\gamma'_0(p^\perp) \cdot p & \gamma_0(p^\perp) \end{pmatrix}. \tag{3.21b}$$

Hence, the weak formulation reads

$$\int_I H_0(x_\rho) x_t \cdot \eta \, d\rho + \int_I \Phi'_0(x_\rho) \cdot \eta_\rho \, d\rho = 0, \quad \forall \eta \in [H^1(I)]^2. \tag{3.22}$$

(3) *Riemannian manifolds:* In view of (2.9), we have $\Phi(z, p) = \frac{1}{2}G(z)p \cdot p$, while

$$\begin{aligned} H(z, p) &= \frac{\det G(z)(G^{-1}(z)p^\perp \cdot p^\perp)^{\frac{3}{2}}}{|G^{-1}(z)p^\perp|^2} \begin{pmatrix} \sqrt{G^{-1}(z)p^\perp \cdot p^\perp} & \frac{G^{-1}(z)p^\perp \cdot p}{\sqrt{G^{-1}(z)p^\perp \cdot p^\perp}} \\ -\frac{G^{-1}(z)p^\perp \cdot p}{\sqrt{G^{-1}(z)p^\perp \cdot p^\perp}} & \sqrt{G^{-1}(z)p^\perp \cdot p^\perp} \end{pmatrix} \\ &= \frac{(\det G(z))G(z)p \cdot p}{|G(z)p|^2} \begin{pmatrix} G(z)p \cdot p & -G(z)p \cdot p^\perp \\ G(z)p \cdot p^\perp & G(z)p \cdot p \end{pmatrix}. \end{aligned} \tag{3.23}$$

Hence, the weak formulation reads

$$\begin{aligned} \int_I H(x, x_\rho) x_t \cdot \eta \, d\rho + \int_I G(x) x_\rho \cdot \eta_\rho \, d\rho \\ + \frac{1}{2} \int_I \eta_i G_{z_i}(x) x_\rho \cdot x_\rho \, d\rho = 0, \quad \forall \eta \in [H^1(I)]^2. \end{aligned}$$

Remark 3.5. It is a straightforward matter to extend our approach for (3.3) to the more general flow

$$\widehat{\beta}(x, v) x_t \cdot v = \varkappa_\gamma \quad \text{in } I \times (0, T]; \tag{3.24}$$

compare also with (1.6) in the space-independent case. In particular, it can be easily shown that if x is a solution to

$$\gamma(x, v) \widehat{\beta}(x, v) H(x, x_\rho) x_t = [\Phi_p(x, x_\rho)]_\rho - \Phi_z(x, x_\rho) \quad \text{in } I \times (0, T],$$

then it automatically solves (3.24). Extending our analysis in Section 4 to this more general case is straightforward, upon making the necessary smoothness assumptions on $\widehat{\beta}$.

4. Finite element approximation

In order to define our finite element approximation, let $0 = q_0 < q_1 < \dots < q_{J-1} < q_J = 1$ be a decomposition of $[0, 1]$ into intervals $I_j = [q_{j-1}, q_j]$. Let $h_j = q_j - q_{j-1}$ as well as

$h = \max_{1 \leq j \leq J} h_j$. We assume that there exists a positive constant c such that

$$h \leq ch_j, \quad 1 \leq j \leq J,$$

so that the resulting family of partitions of $[0, 1]$ is quasiuniform. Within I , we identify $q_J = 1$ with $q_0 = 0$ and define the finite element spaces

$$V^h = \{\chi \in C^0(I) \mid \chi|_{I_j} \text{ is affine, } j = 1, \dots, J\} \quad \text{and} \quad \underline{V}^h = [V^h]^2.$$

Let $\{\chi_j\}_{j=1}^J$ denote the standard basis of V^h . For later use, we let $\pi^h : C^0(I) \rightarrow V^h$ be the standard interpolation operator at the nodes $\{q_j\}_{j=1}^J$, and we use the same notation for the interpolation of vector-valued functions. It is well known that for $k \in \{0, 1\}$, $\ell \in \{1, 2\}$, and $p \in [2, \infty]$, it holds that

$$h^{\frac{1}{p}-\frac{1}{r}} \|\eta_h\|_{0,r} + h|\eta_h|_{1,p} \leq C \|\eta_h\|_{0,p}, \quad \forall \eta_h \in V^h, \quad r \in [p, \infty]; \quad (4.1a)$$

$$|\eta - \pi^h \eta|_{k,p} \leq Ch^{\ell-k} |\eta|_{\ell,p}, \quad \forall \eta \in W^{\ell,p}(I). \quad (4.1b)$$

Our semidiscrete approximation of (3.19) is now given as follows: Find $x_h : I \times [0, T] \rightarrow \Omega$ such that $x_h(\cdot, 0) = \pi^h x_0$ and, for $t \in (0, T]$, $x_h(\cdot, t) \in \underline{V}^h$ such that

$$\begin{aligned} \int_I H(x_h, x_{h,\rho}) x_{h,t} \cdot \eta_h \, d\rho + \int_I \Phi_p(x_h, x_{h,\rho}) \cdot \eta_{h,\rho} \, d\rho \\ + \int_I \Phi_z(x_h, x_{h,\rho}) \cdot \eta_h \, d\rho = 0, \quad \forall \eta_h \in \underline{V}^h. \end{aligned} \quad (4.2)$$

Expanding $x_h(\cdot, t) = \sum_{j=1}^J x_h(q_j, t) \chi_j$, we find that (4.2) gives rise to a system of ordinary differential equations (ODEs) in \mathbb{R}^{2J} which has a unique solution on some interval $[0, T_h)$. By choosing $\eta_h = x_{h,t}$, one also immediately obtains a semidiscrete analogue of (3.20).

In what follows, we assume that (3.9) has a smooth solution $x : I \times [0, T] \rightarrow \Omega$ satisfying

$$0 < c_0 \leq |x_\rho| \leq C_0 \quad \text{in } I \times [0, T] \quad \text{and} \quad \int_0^T \|x_t\|_{0,\infty} \, dt \leq C_0. \quad (4.3)$$

Let $S = x(I \times [0, T])$. Then, there exists $\delta > 0$ such that $\overline{B_\delta(S)} \subset \Omega$ and we define the compact set $K = \overline{B_\delta(S)} \times (\overline{B_{2C_0}(0)} \setminus B_{\frac{c_0}{2}}(0)) \subset \Omega \times (\mathbb{R}^2 \setminus \{0\})$. We may choose $M_K \geq 0$ and $c_1 > 0$ such that

$$\max_{|\beta| \leq 3} \max_{(z,p) \in K} |D^\beta \gamma(z, p)| \leq M_K, \quad \max_{|\beta| \leq 2} \max_{z \in \overline{B_\delta(S)}} |D^\beta a(z)| \leq M_K, \quad (4.4)$$

and

$$\gamma(z, p) \geq c_1, \quad |\gamma_p(z, p)| \geq c_1, \quad a(z) \geq c_1, \quad \forall z \in \overline{B_\delta(S)}, \quad p \in \overline{B_{2C_0}(0)} \setminus B_{\frac{c_0}{2}}(0). \quad (4.5)$$

Theorem 4.1. *Suppose that (3.9) has a smooth solution $x : I \times [0, T] \rightarrow \Omega$ satisfying (4.3). Then, there exists $h_0 > 0$ such that for $0 < h \leq h_0$, semidiscrete problem (4.2) has a unique solution $x_h : I \times [0, T] \rightarrow \Omega$, and the following error bounds hold:*

$$\max_{t \in [0, T]} \|x(\cdot, t) - x_h(\cdot, t)\|_1^2 + \int_0^T \|x_t - x_{h,t}\|_0^2 dt \leq Ch^2. \tag{4.6}$$

Proof. Let us define

$$\widehat{T}_h = \sup \left\{ t \in [0, T] \mid x_h \text{ solves (4.2) on } [0, t], \text{ with } \int_0^t \|x_{h,t}\|_{0,\infty} ds \leq 2C_0 \text{ and } \|(x - x_h)(\cdot, s)\|_{0,\infty} \leq \delta, \|(x_\rho - x_{h,\rho})(\cdot, s)\|_{0,\infty} \leq \frac{1}{2}c_0, 0 \leq s \leq t \right\}.$$

Let $(\rho, t) \in I \times [0, \widehat{T}_h]$. Since $|(1 - \lambda)x(\rho, t) + \lambda x_h(\rho, t) - x(\rho, t)| \leq \|(x - x_h)(\cdot, t)\|_{0,\infty} \leq \delta$ for all $\lambda \in [0, 1]$, we find that $[x(\rho, t), x_h(\rho, t)] \subset B_\delta(\overline{S})$. Arguing in a similar way for the first derivative, we deduce that

$$[x(\rho, t), x_h(\rho, t)] \times [x_\rho(\rho, t), x_{h,\rho}(\rho, t)] \subset K, \quad \forall (\rho, t) \in I \times [0, \widehat{T}_h]. \tag{4.7}$$

Comparing (3.19) and (4.2), we see that the error $e = x - x_h$ satisfies

$$\begin{aligned} & \int_I H(x_h, x_{h,\rho}) e_t \cdot \eta_h \, d\rho + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot \eta_{h,\rho} \, d\rho \\ &= \int_I (H(x_h, x_{h,\rho}) - H(x, x_\rho)) x_t \cdot \eta_h \, d\rho + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot \eta_{h,\rho} \, d\rho \\ & \quad + \int_I (\Phi_z(x_h, x_{h,\rho}) - \Phi_z(x, x_\rho)) \cdot \eta_h \, d\rho, \quad \forall \eta_h \in \underline{V}^h. \end{aligned} \tag{4.8}$$

Using the identity

$$e = x - \pi^h x + \pi^h e \tag{4.9}$$

and choosing $\eta_h = \pi^h e_t$ in (4.8), we obtain

$$\begin{aligned} & \int_I H(x_h, x_{h,\rho}) e_t \cdot e_t \, d\rho + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot e_{t\rho} \, d\rho \\ &= \int_I H(x_h, x_{h,\rho}) e_t \cdot (x_t - \pi^h x_t) \, d\rho \\ & \quad + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot (x_t - \pi^h x_t)_\rho \, d\rho \\ & \quad + \int_I (H(x_h, x_{h,\rho}) - H(x, x_\rho)) x_t \cdot \pi^h e_t \, d\rho \\ & \quad + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e_t)_\rho \, d\rho \\ & \quad + \int_I (\Phi_z(x_h, x_{h,\rho}) - \Phi_z(x, x_\rho)) \cdot \pi^h e_t \, d\rho =: \sum_{i=1}^5 S_i. \end{aligned} \tag{4.10}$$

Let us begin with the two terms on the left-hand side of (4.10). Clearly, (3.16), (4.4), (4.5), and (4.7) imply that

$$\int_I H(x_h, x_{h,\rho}) e_t \cdot e_t \, d\rho = \int_I \frac{a^2(x_h) \gamma^2(x_h, x_{h,\rho}^\perp)}{|\gamma_p(x_h, x_{h,\rho}^\perp)|^2} |e_t|^2 \, d\rho \geq \tilde{c}_0 \|e_t\|_0^2, \tag{4.11}$$

where $\tilde{c}_0 = M_K^{-2} c_1^4$. Next, we write

$$\begin{aligned} & (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot e_{t\rho} \\ &= \Phi_p(x_h, x_\rho) \cdot x_{t\rho} + \Phi_p(x_h, x_{h,\rho}) \cdot x_{h,t\rho} - \Phi_p(x_h, x_\rho) \cdot x_{h,t\rho} - \Phi_p(x_h, x_{h,\rho}) \cdot x_{t\rho} \\ &= [\Phi(x_h, x_{h,\rho}) - \Phi_p(x_h, x_\rho) \cdot x_{h,\rho}]_t + \Phi_p(x_h, x_\rho) \cdot x_{t\rho} - \Phi_p(x_h, x_{h,\rho}) \cdot x_{t\rho} \\ &\quad - \Phi_z(x_h, x_{h,\rho}) \cdot x_{h,t} + [\Phi_p(x_h, x_\rho)]_t \cdot x_{h,\rho}. \end{aligned}$$

Since $p \mapsto \Phi(z, p)$ and $p \mapsto \Phi_{z_j}(z, p)$ are positively homogeneous of degree 2, we have

$$\Phi_p(x_h, x_\rho) \cdot x_\rho = 2\Phi(x_h, x_\rho) \quad \text{and} \quad \Phi_{pz_j}(x_h, x_\rho) \cdot x_\rho = 2\Phi_{z_j}(x_h, x_\rho),$$

and therefore,

$$\begin{aligned} & (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot e_{t\rho} \\ &= [\Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)]_t - [\Phi(x_h, x_\rho)]_t \\ &\quad + \Phi_p(x_h, x_\rho) \cdot x_{t\rho} - \Phi_p(x_h, x_{h,\rho}) \cdot x_{t\rho} - \Phi_z(x_h, x_{h,\rho}) \cdot x_{h,t} \\ &\quad + [\Phi_p(x_h, x_\rho)]_t \cdot x_{h,\rho} \\ &= [\Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)]_t \\ &\quad - (\Phi_z(x_h, x_{h,\rho}) + \Phi_z(x_h, x_\rho)) \cdot x_{h,t} \\ &\quad - \Phi_p(x_h, x_{h,\rho}) \cdot x_{t\rho} + x_{h,j,t} \Phi_{pz_j}(x_h, x_\rho) \cdot x_{h,\rho} + \Phi_{pp}(x_h, x_\rho) x_{t\rho} \cdot x_{h,\rho} \\ &= [\Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)]_t \\ &\quad - (\Phi_p(x_h, x_{h,\rho}) - \Phi_{pp}(x_h, x_\rho) x_{h,\rho}) \cdot x_{t\rho} \\ &\quad - (\Phi_{z_j}(x_h, x_{h,\rho}) - \Phi_{z_j}(x_h, x_\rho) - \Phi_{pz_j}(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)) x_{h,j,t} \\ &= [\Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)]_t \\ &\quad - (\Phi_p(x_h, x_{h,\rho}) - \Phi_p(x_h, x_\rho) - \Phi_{pp}(x_h, x_\rho) (x_{h,\rho} - x_\rho)) \cdot x_{t\rho} \\ &\quad - (\Phi_{z_j}(x_h, x_{h,\rho}) - \Phi_{z_j}(x_h, x_\rho) - \Phi_{pz_j}(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho)) x_{h,j,t}, \tag{4.12} \end{aligned}$$

where the last equality follows from the relation $\Phi_p(z, p) = \Phi_{pp}(z, p)p$ (recall (2.2)). If we combine (4.12) with (4.11) and use a Taylor expansion together with (4.4) and (4.7), we obtain for the left-hand side of (4.10) that

$$\begin{aligned} & \int_I H(x_h, x_{h,\rho}) e_t \cdot e_t \, d\rho + \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x_h, x_{h,\rho})) \cdot e_{t\rho} \, d\rho \\ & \geq \tilde{c}_0 \|e_t\|_0^2 + \frac{d}{dt} \int_I \Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho) \, d\rho \\ & \quad - C(\|x_{t\rho}\|_{0,\infty} + \|x_{h,t}\|_{0,\infty}) \|e_\rho\|_0^2. \tag{4.13} \end{aligned}$$

Let us next estimate the terms on the right-hand side of (4.10). To begin, we obtain from (3.15), (4.4), (4.5), and (4.1b) that

$$S_1 \leq C \|e_t\|_0 \|x_t - \pi^h x_t\|_0 \leq Ch \|x_{t\rho}\|_0 \|e_t\|_0 \leq \varepsilon \|e_t\|_0^2 + C_\varepsilon h^2 \|x_{t\rho}\|_0^2. \tag{4.14a}$$

The remaining terms involve differences between Φ_p , H , and Φ_z , which will be estimated with the help of (4.4) and (4.7). Using (4.1b), we have

$$S_2 \leq C \|e_\rho\|_0 \|(x_t - \pi^h x_t)_\rho\|_0 \leq Ch \|x_{t\rho\rho}\|_0 \|e_\rho\|_0 \leq \|e_\rho\|_0^2 + Ch^2 \|x_{t\rho\rho}\|_0^2, \tag{4.14b}$$

as well as

$$\begin{aligned} S_3 &\leq C \|e\|_1 \|x_t\|_{0,\infty} \|\pi^h e_t\|_0 \leq C \|e\|_1 (\|e_t\|_0 + \|x_t - \pi^h x_t\|_0) \\ &\leq \varepsilon \|e_t\|_0^2 + C_\varepsilon \|e\|_1^2 + Ch^2 \|x_{t\rho}\|_0^2, \end{aligned} \tag{4.14c}$$

and similarly,

$$S_5 \leq C \|e\|_1 \|\pi^h e_t\|_0 \leq \varepsilon \|e_t\|_0^2 + C_\varepsilon \|e\|_1^2 + Ch^2 \|x_{t\rho}\|_0^2. \tag{4.14d}$$

Finally, noting once again identity (4.9) and estimate (4.1b), we have

$$\begin{aligned} S_4 &= \frac{d}{dt} \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e)_\rho \, d\rho \\ &\quad - \int_I [\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)]_t \cdot (\pi^h e)_\rho \, d\rho \\ &= \frac{d}{dt} \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e)_\rho \, d\rho \\ &\quad + \int_I [\Phi_{pz}(x_h, x_\rho) e_t + (\Phi_{pz}(x, x_\rho) - \Phi_{pz}(x_h, x_\rho)) x_t] \cdot (\pi^h e)_\rho \, d\rho \\ &\quad + \int_I [(\Phi_{pp}(x, x_\rho) - \Phi_{pp}(x_h, x_\rho)) x_{t\rho}] \cdot (\pi^h e)_\rho \, d\rho \\ &\leq \frac{d}{dt} \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e)_\rho \, d\rho \\ &\quad + C(\|e_t\|_0 + \|e\|_{0,\infty} \|x_t\|_1) (\|e_\rho\|_0 + \|x_\rho - (\pi^h x)_\rho\|_0) \\ &\leq \frac{d}{dt} \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e)_\rho \, d\rho + \varepsilon \|e_t\|_0^2 + C_\varepsilon \|e\|_1^2 \\ &\quad + C_\varepsilon h^2 \|x_{\rho\rho}\|_0^2, \end{aligned} \tag{4.14e}$$

where in the last inequality we have also used embedding result (1.8). If we insert (4.13) and (4.14) into (4.10), and choose ε sufficiently small, we obtain

$$\frac{1}{2} \tilde{c}_0 \|e_t\|_0^2 + \mu'(t) \leq C(1 + \|x_{h,t}\|_{0,\infty}) \|e\|_1^2 + Ch^2, \tag{4.15}$$

where

$$\begin{aligned} \mu(t) &= \int_I \Phi(x_h, x_{h,\rho}) - \Phi(x_h, x_\rho) - \Phi_p(x_h, x_\rho) \cdot (x_{h,\rho} - x_\rho) \, d\rho \\ &\quad - \int_I (\Phi_p(x_h, x_\rho) - \Phi_p(x, x_\rho)) \cdot (\pi^h e)_\rho \, d\rho. \end{aligned}$$

Clearly, we have from (3.18), on noting (4.9) and (4.1b), that

$$\begin{aligned} \mu(t) &\geq \frac{1}{2} \sigma_K \|e_\rho\|_0^2 - C \|e\|_0 \|(\pi^h e)_\rho\|_0 \\ &\geq \frac{1}{2} \sigma_K \|e_\rho\|_0^2 - C \|e\|_0 (\|e_\rho\|_0 + \|x_\rho - (\pi^h x)_\rho\|_0) \\ &\geq \frac{1}{4} \sigma_K \|e_\rho\|_0^2 - C_1 \|e\|_0^2 - Ch^2, \end{aligned}$$

and hence,

$$\begin{aligned} \|e\|_1^2 &= \|e\|_0^2 + \|e_\rho\|_0^2 \\ &\leq \|e\|_0^2 + \frac{4}{\sigma_K} [\mu(t) + C_1 \|e\|_0^2 + Ch^2] \\ &\leq \frac{4}{\sigma_K} (C_2 \|e\|_0^2 + \mu(t)) + Ch^2, \end{aligned} \tag{4.16}$$

where $C_2 = C_1 + \frac{\sigma_K}{4}$. Integrating (4.15) with respect to time, and observing (4.16) as well as $\mu(0) \leq Ch^2$, we derive

$$\begin{aligned} &\frac{1}{2} \tilde{c}_0 \int_0^t \|e_\tau\|_0^2 \, ds + (\mu(t) + C_2 \|e(t)\|_0^2) \\ &\leq C \int_0^t (1 + \|x_{h,t}\|_{0,\infty}) \|e\|_1^2 \, ds + 2C_2 \int_0^t \|e\|_0 \|e_\tau\|_0 \, ds + Ch^2 \\ &\leq C \int_0^t (1 + \|x_{h,t}\|_{0,\infty}) (\mu(s) + C_2 \|e\|_0^2) \, ds + \frac{1}{4} \tilde{c}_0 \int_0^t \|e_\tau\|_0^2 \, ds \\ &\quad + C \int_0^t \|e\|_0^2 \, ds + Ch^2, \end{aligned}$$

and hence,

$$\mu(t) + C_2 \|e(t)\|_0^2 \leq Ch^2 + C \int_0^t (1 + \|x_{h,t}\|_{0,\infty}) (\mu(s) + C_2 \|e\|_0^2) \, ds, \quad 0 \leq t < \widehat{T}_h.$$

Since $\int_0^{\widehat{T}_h} \|x_{h,t}\|_{0,\infty} \, ds \leq 2C_0$ by definition, we deduce with the help of Gronwall's

inequality and (4.16) that

$$\int_0^{\widehat{T}_h} \|e_t\|_0^2 \, ds + \sup_{0 \leq s \leq \widehat{T}_h} \|e(s)\|_1^2 \leq Ch^2. \tag{4.17}$$

In particular, we have on recalling (1.8) that

$$\|x(\cdot, \widehat{T}_h) - x_h(\cdot, \widehat{T}_h)\|_{0,\infty} = \|e(\cdot, \widehat{T}_h)\|_{0,\infty} \leq C \|e(\cdot, \widehat{T}_h)\|_1 \leq Ch \leq \frac{1}{2}\delta,$$

provided that $0 < h \leq h_0$. Next, we have from (4.1b), (4.1a), (4.9), and (4.17) that

$$\begin{aligned} \|(x_\rho - x_{h,\rho})(\cdot, \widehat{T}_h)\|_{0,\infty} &\leq \|(x_\rho - (\pi^h x)_\rho)(\cdot, \widehat{T}_h)\|_{0,\infty} + \|((\pi^h x)_\rho - x_{h,\rho})(\cdot, \widehat{T}_h)\|_{0,\infty} \\ &\leq Ch + Ch^{-\frac{1}{2}} \|(x_{h,\rho} - (\pi^h x)_\rho)(\cdot, \widehat{T}_h)\|_0 \\ &\leq Ch^{\frac{1}{2}} + Ch^{-\frac{1}{2}} \|e_\rho(\cdot, \widehat{T}_h)\|_0 \leq Ch^{\frac{1}{2}}, \end{aligned}$$

and similarly, by (4.3), that

$$\begin{aligned} \int_0^{\widehat{T}_h} \|x_{h,t}\|_{0,\infty} \, ds &\leq \int_0^{\widehat{T}_h} (\|x_t\|_{0,\infty} \, ds + \|e_t\|_{0,\infty}) \, ds \\ &\leq C_0 + Ch^{-\frac{1}{2}} \int_0^{\widehat{T}_h} \|e_t\|_0 \, ds \leq C_0 + Ch^{\frac{1}{2}}. \end{aligned}$$

By choosing h_0 to be smaller if necessary, we may therefore assume that the inequalities $\|(x_\rho - x_{h,\rho})(\cdot, \widehat{T}_h)\|_{0,\infty} \leq \frac{1}{4}c_0$ and $\int_0^{\widehat{T}_h} \|x_{h,t}\|_{0,\infty} \, ds \leq \frac{3}{2}C_0$ hold. Suppose that $\widehat{T}_h < T$. Then, there exists an $\varepsilon > 0$ such that x_h exists on $[0, \widehat{T}_h + \varepsilon]$ with $\|(x - x_h)(\cdot, t)\|_{0,\infty} \leq \delta$, $\|(x_\rho - x_{h,\rho})(\cdot, t)\|_{0,\infty} \leq \frac{1}{2}c_0$ for $0 \leq t \leq \widehat{T}_h + \varepsilon$, and $\int_0^{\widehat{T}_h + \varepsilon} \|x_{h,t}\|_{0,\infty} \, ds \leq 2C_0$, which contradicts the definition of \widehat{T}_h . Thus, $\widehat{T}_h = T$ and the theorem is proved. ■

5. Fully discrete schemes

From now on, let the L^2 -inner product on I be denoted by (\cdot, \cdot) . Due to the nonlinearities present in (4.2), for a fully practical scheme we need to introduce numerical quadrature. For our purposes, it is sufficient to consider classical mass lumping. Hence, for two piecewise continuous functions with possible jumps at the nodes $\{q_j\}_{j=1}^J$, we define the mass lumped L^2 -inner product $(u, v)^h$ via

$$(u, v)^h = \frac{1}{2} \sum_{j=1}^J h_j [(uv)(q_j^-) + (uv)(q_j^+)], \tag{5.1}$$

where $u(q_j^\pm) = \lim_{\delta \searrow 0} u(q_j \pm \delta)$. The definition in (5.1) naturally extends to vector-valued functions.

In particular, we will consider fully discrete approximations of

$$\begin{aligned} & (H(x_h, x_{h,\rho})x_{h,t}, \eta_h)^h + (\Phi_p(x_h, x_{h,\rho}), \eta_{h,\rho})^h \\ & + (\Phi_z(x_h, x_{h,\rho}), \eta_h)^h = 0, \quad \forall \eta_h \in \underline{V}^h \end{aligned} \tag{5.2}$$

in place of (4.2). Using this quadrature does not affect our derived error estimate in (4.6), as can be shown with standard techniques.

In order to discretize (5.2) in time, let $t_m = m\Delta t$, $m = 0, \dots, M$, with the uniform time step size $\Delta t = \frac{T}{M} > 0$. In the following, we let $x_h^0 = x_h(\cdot, 0) = \pi^h x_0 \in \underline{V}^h$ and for $m = 0, \dots, M - 1$, let $x_h^{m+1} \in \underline{V}^h$ be the solution of a system of algebraic equations, which we will specify. In general, we will attempt to define fully discrete approximations that are unconditionally stable, in the sense that they satisfy the following discrete analogue of (3.20):

$$\begin{aligned} & (\Phi(x_h^k, x_{h,\rho}^k) - \Phi(x_h^0, x_{h,\rho}^0), 1)^h \\ & \leq -\Delta t \sum_{m=0}^{k-1} \left(H(x_h^m, x_{h,\rho}^m) \frac{x_h^{m+1} - x_h^m}{\Delta t}, \frac{x_h^{m+1} - x_h^m}{\Delta t} \right)^h \leq 0, \end{aligned} \tag{5.3}$$

for $k = 1, \dots, M$. Here, the second inequality is a consequence of (3.16).

5.1. Space-independent anisotropic curve shortening flow

Let $\gamma(z, p) = \gamma_0(p)$ be an anisotropy function and let $\Phi_0(p) = \frac{1}{2}\gamma_0^2(p)$. Using Example 3.4(2), we propose the scheme

$$\frac{1}{\Delta t} (H_0(x_{h,\rho}^m)(x_h^{m+1} - x_h^m), \eta_h)^h + (\Phi'_0(x_{h,\rho}^{m+1}), \eta_{h,\rho}) = 0, \quad \forall \eta_h \in \underline{V}^h, \tag{5.4}$$

where H_0 is defined in (3.21b). We remark that in the isotropic case, scheme (5.4) is linear and collapses to the fully discrete approximation in [19, p. 108].

Lemma 5.1. *A solution $(x_h^m)_{m=0}^M$ to (5.4) satisfies stability bound (5.3).*

Proof. Choose $\eta_h = x_h^{m+1} - x_h^m$ in (5.4), use (3.18), and sum from $m = 0$ to $k - 1$. ■

A disadvantage of the scheme in (5.4) is that at each time level, a nonlinear system of equations needs to be solved. Following the approach in [20] (see also [38]), one could alternatively consider a linear scheme by introducing a suitable stabilization term and treating the elliptic term in (5.4) fully explicitly.

If we restrict our attention to a special class of anisotropies, then a linear and unconditionally stable approximation can be introduced that does not rely on a stabilization term. This idea goes back to [5], and was extended to the phase field context in [10]. In fact, a wide class of anisotropies can either be modeled or at least very well approximated by

$$\gamma_0(p) = \sum_{\ell=1}^L \sqrt{\Lambda_\ell p \cdot p}, \tag{5.5}$$

where $\Lambda_\ell \in \mathbb{R}^{2 \times 2}$, $\ell = 1, \dots, L$, are symmetric and positive definite; see [4, 5]. Hence, assumption (2.3) is satisfied. In order to be able to apply the ideas in [10, 11], we define the auxiliary function $\phi_0(p) = \gamma_0(p^\perp)$, so that $\Phi_0(p) = \frac{1}{2}\gamma_0^2(p^\perp) = \frac{1}{2}\phi_0^2(p)$. Observe that ϕ_0 also falls within the class of densities of the form (5.5), that is,

$$\phi_0(p) = \gamma_0(p^\perp) = \sum_{\ell=1}^L \sqrt{\tilde{\Lambda}_\ell p \cdot p}, \quad \text{where} \quad \tilde{\Lambda}_\ell = \det(\Lambda_\ell)\Lambda_\ell^{-1}.$$

Moreover, we recall from [10, 11] that $\Phi'_0(p) = B(p)p$, if we introduce the matrices

$$B(p) = \begin{cases} \gamma_0(p^\perp) \sum_{\ell=1}^L \frac{\tilde{\Lambda}_\ell}{\sqrt{\tilde{\Lambda}_\ell p \cdot p}} & p \neq 0, \\ L \sum_{\ell=1}^L \tilde{\Lambda}_\ell & p = 0. \end{cases} \tag{5.6}$$

On recalling once again the definition of H_0 from (3.21b), we then consider the scheme

$$\frac{1}{\Delta t} (H_0(x_{h,\rho}^m)(x_h^{m+1} - x_h^m), \eta_h)^h + (B(x_{h,\rho}^m)x_h^{m+1}, \eta_{h,\rho}) = 0, \quad \forall \eta_h \in \underline{V}^h, \tag{5.7}$$

which is inspired by the treatment of the anisotropy in [10, 11] and leads to a system of linear equations. We note that for the case $L = 1$, the two schemes in (5.7) and (5.4) are identical.

Lemma 5.2. *Suppose that $x_h^m \in \underline{V}^h$ with $x_{h,\rho}^m \neq 0$ in I . Then, there exists a unique solution $x_h^{m+1} \in \underline{V}^h$ to (5.7). Moreover, a solution $(x_h^m)_{m=0}^M$ to (5.7) satisfies the stability bound in (5.3).*

Proof. Existence follows from uniqueness, and so we consider the following homogeneous system: Find $X_h \in \underline{V}^h$ such that

$$\frac{1}{\Delta t} (H_0(x_{h,\rho}^m)X_h, \eta_h)^h + (B(x_{h,\rho}^m)X_{h,\rho}, \eta_{h,\rho}) = 0, \quad \forall \eta_h \in \underline{V}^h.$$

Choosing $\eta_h = X_h$, and observing that the matrices $B(p)$ are positive definite, we obtain

$$0 = (H_0(x_{h,\rho}^m)X_h, X_h)^h + \Delta t (B(x_{h,\rho}^m)X_{h,\rho}, X_{h,\rho}) \geq (H_0(x_{h,\rho}^m)X_h, X_h)^h,$$

which implies that $X_h = 0$ in view of (3.16). This proves the existence of a unique solution. In order to show the stability bound, we recall from [10, Corollary 2.3] that with B as defined in (5.6), it holds that

$$B(q)p \cdot (p - q) \geq \Phi_0(p) - \Phi_0(q), \quad \forall p, q \in \mathbb{R}^2.$$

Hence, choosing $\eta_h = x_h^{m+1} - x_h^m$ in (5.7) and summing from $m = 0$ to $k - 1$ yields (5.3). ■

5.2. Curve shortening flow in Riemannian manifolds

Let us denote by $\text{Sym}(2, \mathbb{R})$ the set of symmetric 2×2 matrices over \mathbb{R} . For $A, B \in \text{Sym}(2, \mathbb{R})$, we define $A \succcurlyeq B$ if and only if $A - B$ is positive semidefinite. We say that a differentiable function $f : \Omega \rightarrow \text{Sym}(2, \mathbb{R})$ is convex if

$$f(w) \succcurlyeq f(z) + (w_i - z_i)f_{z_i}(z) \quad \text{for all } z, w \in \Omega \text{ such that } [z, w] \subset \Omega.$$

We now consider the situation in Example 3.4(3). Let $G : \Omega \rightarrow \text{Sym}(2, \mathbb{R})$, so that $\Phi(z, p) = \frac{1}{2}G(z)p \cdot p$, where $G(z)$ is positive definite for $z \in \Omega$. In order to obtain an unconditionally stable scheme, we adapt an idea from [12] and assume that we can split G into

$$G = G_+ + G_- \quad \text{such that } \pm G_{\pm} : \Omega \rightarrow \text{Sym}(2, \mathbb{R}) \text{ are convex.} \quad (5.8)$$

Such a splitting exists if there exists a constant $c_G \in \mathbb{R}_{\geq 0}$ such that

$$\frac{1}{2}\lambda_i\lambda_j G_{z_i z_j}(z) + c_G |\lambda|^2 \text{Id} \succcurlyeq 0, \quad \forall z \in \Omega, \lambda \in \mathbb{R}^2.$$

In that case, one may choose $G_+(z) = G(z) + c_G |z|^2 \text{Id}$ and $G_-(z) = -c_G |z|^2 \text{Id}$. It follows from (5.8) that

$$\begin{aligned} (w_i - z_i)(G_{+,z_i}(w) + G_{-,z_i}(z)) &\succcurlyeq G(w) - G(z), \\ \forall w, z \in \Omega \text{ such that } [z, w] &\subset \Omega. \end{aligned} \quad (5.9)$$

We now consider the scheme

$$\begin{aligned} \frac{1}{\Delta t} (H(x_h^m, x_{h,\rho}^m)(x_h^{m+1} - x_h^m), \eta_h)^h + (G(x_h^m)x_{h,\rho}^{m+1}, \eta_{h,\rho})^h \\ + \frac{1}{2}(\eta_{h,i}(G_{+,z_i}(x_h^{m+1}) + G_{-,z_i}(x_h^m))x_{h,\rho}^{m+1}, x_{h,\rho}^{m+1})^h = 0, \quad \forall \eta_h \in \underline{V}^h, \end{aligned} \quad (5.10)$$

where H is as defined in (3.23). We note that in general (5.10) is a nonlinear scheme.

Lemma 5.3. *Let $(x_h^m)_{m=0}^M$ be a solution to equation (5.10), with $[x_h^m(q_j), x_h^{m+1}(q_j)] \subset \Omega$ for $j = 1, \dots, J$ and $m = 0, \dots, M - 1$. Then, stability bound (5.3) is satisfied.*

Proof. Let us again choose $\eta_h = x_h^{m+1} - x_h^m$ in (5.10) and calculate with the help of (5.9)

$$\begin{aligned} (G(x_h^m)x_{h,\rho}^{m+1}, x_{h,\rho}^{m+1} - x_{h,\rho}^m)^h \\ + \frac{1}{2}((x_{h,i}^{m+1} - x_{h,i}^m)(G_{+,z_i}(x_h^{m+1}) + G_{-,z_i}(x_h^m))x_{h,\rho}^{m+1}, x_{h,\rho}^{m+1})^h \\ \geq (G(x_h^m)x_{h,\rho}^{m+1}, x_{h,\rho}^{m+1})^h - (G(x_h^m)x_{h,\rho}^{m+1}, x_{h,\rho}^m)^h \\ + \frac{1}{2}((G(x_h^{m+1}) - G(x_h^m))x_{h,\rho}^{m+1}, x_{h,\rho}^{m+1})^h \end{aligned}$$

$$\begin{aligned}
 &= \frac{1}{2}(G(x_h^m)(x_{h,\rho}^{m+1} - x_{h,\rho}^m), x_{h,\rho}^{m+1} - x_{h,\rho}^m)^h \\
 &\quad + (\Phi(x_h^{m+1}, x_{h,\rho}^{m+1}), 1)^h - (\Phi(x_h^m, x_{h,\rho}^m), 1)^h \\
 &\geq (\Phi(x_h^{m+1}, x_{h,\rho}^{m+1}), 1)^h - (\Phi(x_h^m, x_{h,\rho}^m), 1)^h,
 \end{aligned}$$

since $G(x_h^m)$ is symmetric and positive definite. This yields the desired result in a similar manner to the proof of Lemma 5.2. ■

In the special case that the Riemannian manifold is conformally equivalent to the Euclidean plane, that is, when $G(z) = g(z)\text{Id}$ for $g : \Omega \rightarrow \mathbb{R}_{>0}$, several numerical schemes have been proposed in [12, §3.1]. In this situation, our fully discrete approximation (5.10) collapses to the new scheme

$$\begin{aligned}
 &\frac{1}{\Delta t}(g^2(x_h^m)(x_h^{m+1} - x_h^m), \eta_h |x_{h,\rho}^m|^2)^h + (g(x_h^m)x_{h,\rho}^{m+1}, \eta_{h,\rho})^h \\
 &\quad + \frac{1}{2}((\nabla g_+(x_h^{m+1}) + \nabla g_-(x_h^m)), \eta_h |x_{h,\rho}^{m+1}|^2)^h = 0, \quad \forall \eta_h \in \underline{V}^h, \tag{5.11}
 \end{aligned}$$

where $g = g_+ + g_-$ and $\pm g_\pm : \Omega \rightarrow \mathbb{R}$ are convex. We observe that for the special case $g(z) = (z_1)^2$, scheme (5.11) is in fact very close to approximation [3, (4.4)], modulo the different time scaling factor that arises in the context of mean curvature flow for axisymmetric hypersurfaces in \mathbb{R}^3 considered there.

6. Numerical results

We implemented our fully discrete schemes within the finite element toolbox Alberta [41]. Where the systems of equations arising at each time level are nonlinear, they are solved using a Newton method or a Picard-type iteration, while all linear (sub)problems are solved with the help of the sparse factorization package UMFPACK; see [18]. For example, for the solution of (5.4) we employ a Newton iteration, while the Picard iteration for the solution of (5.10) is defined through $x_h^{m+1,0} = x_h^m$ and, for $i \geq 0$, by $x_h^{m+1,i+1} \in \underline{V}^h$ such that

$$\begin{aligned}
 &\frac{1}{\Delta t}(H(x_h^m, x_{h,\rho}^m)(x_h^{m+1,i+1} - x_h^m), \eta_h)^h \\
 &\quad + (G(x_h^m)x_{h,\rho}^{m+1,i+1}, \eta_{h,\rho})^h + \frac{1}{2}(\eta_{h,i}(G_{+,z_i}(x_h^{m+1,i}) \\
 &\quad + G_{-,z_i}(x_h^m))x_{h,\rho}^{m+1,i}, x_{h,\rho}^{m+1,i})^h = 0, \quad \forall \eta_h \in \underline{V}^h. \tag{6.1}
 \end{aligned}$$

In all our simulations, the Newton solver for (5.4) converged in at most one iteration, while the Picard iteration in (6.1) always converged in at most three iterations.

For all our numerical simulations, we use a uniform partitioning of $[0, 1]$, so that $q_j = jh, j = 0, \dots, J$, with $h = \frac{1}{J}$. Unless otherwise stated, we use $J = 128$ and $\Delta t = 10^{-4}$. On recalling (1.1), for $\chi_h \in \underline{V}^h$ we define the discrete energy

$$\mathcal{E}^h(\chi_h) = (a(\chi_h), \gamma(\chi_h, \chi_{h,\rho}^\perp))^h.$$

We also consider the ratio

$$r^m = \frac{\max_{j=1,\dots,J} |x_h^m(q_j) - x_h^m(q_{j-1})|}{\min_{j=1,\dots,J} |x_h^m(q_j) - x_h^m(q_{j-1})|} \tag{6.2}$$

between the longest and shortest element of $\Gamma_h^m = x_h^m(I)$; we are often interested in the evolution of this ratio over time. We stress that no redistribution of vertices was necessary during any of our numerical simulations. In the isotropic case, this can be explained by the diffusive character of the tangential motion induced by (3.4), since the flow can be rewritten as $x_t = \kappa\nu - (\frac{1}{|x_\rho|})\rho\tau$, as has been pointed out in, for example, [34, p. 1477]. Our numerical experiments indicate that while the induced tangential motion from (3.9) may not be diffusive in general, it is sufficiently well-behaved to avoid coalescence of vertices in practice.

6.1. Space-independent anisotropic curve shortening flow

In this subsection, we consider the situation from Example 2.2(2); in addition, see Example 3.4(2). Anisotropies of the form $\gamma(z, p) = \gamma_0(p)$ can be visualized by their Frank diagram $\mathcal{F} = \{p \in \mathbb{R}^2 : \gamma_0(p) \leq 1\}$ and their Wulff shape $\mathcal{W} = \{q \in \mathbb{R}^2 : \gamma_0^*(q) \leq 1\}$, where γ_0^* is the dual of γ_0 defined by

$$\gamma_0^*(q) = \sup_{p \in \mathbb{R}^2 \setminus \{0\}} \frac{p \cdot q}{\gamma_0(p)}.$$

We recall from [28] that the boundary of the Wulff shape, $\partial\mathcal{W}$, is the solution of the isoperimetric problem for $\mathcal{E}(\Gamma) = \int_\Gamma \gamma_0(\nu) \, d\mathcal{H}^1$. Moreover, it was shown in [42] that self-similarly shrinking boundaries of Wulff shapes are a solution to anisotropic curve shortening flow. In particular,

$$\Gamma(t) = \{q \in \mathbb{R}^2 \mid \gamma_0^*(q) = \sqrt{1 - 2t}\} \tag{6.3}$$

solves (2.8). We demonstrate this behavior in Figure 1 for the ‘‘elliptic’’ anisotropy

$$\gamma_0(p) = \sqrt{p_1^2 + \delta^2 p_2^2}, \quad \delta = 0.5. \tag{6.4}$$

Observe that (6.4) is a special case of (5.5) with $L = 1$, so that the scheme in (5.4) collapses to the linear scheme in (5.7). The evolution in Figure 1 nicely shows how the curve shrinks self-similarly to a point. We can also see that the scheme in (5.4) induces a tangential motion that moves the initially equidistributed vertices along the curve, so that

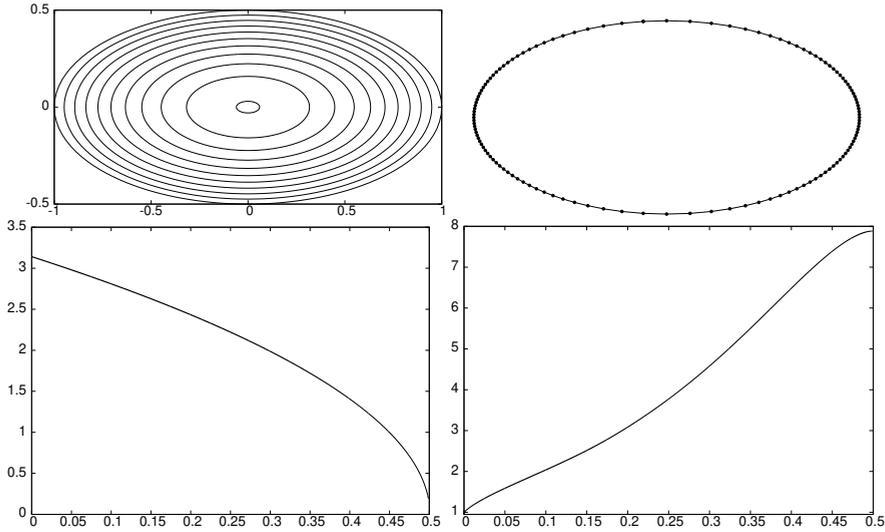


Figure 1. Anisotropic curvature flow for (6.4) using scheme (5.4). The solution at times $t = 0, 0.05, \dots, 0.45, 0.499$, as well as the distribution of vertices on $x_h^M(I)$, are shown. Below we show a plot of the discrete energy $\mathcal{E}^h(x_h^m)$ and of ratio (6.2) over time.

eventually a higher density of vertices can be observed in regions of larger curvature. We note that this behavior is not dissimilar to the behavior observed in the numerical experiments in [5].

We now use the exact solution given by (6.3) to perform a convergence experiment for our proposed finite element approximation (see (4.2)). To this end, we choose the particular parametrization

$$x(\rho, t) = (1 - 2t)^{\frac{1}{2}} \begin{pmatrix} \cos(2\pi\rho) \\ \delta \sin(2\pi\rho) \end{pmatrix} \tag{6.5}$$

and define

$$f = H_0(x_\rho)x_t - [\Phi'_0(x_\rho)]_\rho, \tag{6.6}$$

so that (6.5) is the exact solution of (3.22) with the additional right-hand side (f, η) . Upon adding the right-hand side $(f(t_{m+1}), \eta_h)^h$ to (5.4), we can thus use (6.5) as a reference solution for a convergence experiment of our proposed finite element approximation. We report on the observed H^1 - and L^2 -errors for the scheme in (5.4) for a sequence of mesh sizes in Table 1. Here we partition the time interval $[0, T]$, with $T = 0.45$, into uniform time steps of size $\Delta t = h^2$ for $h = J^{-1} = 2^{-k}$, $k = 5, \dots, 9$. The observed numerical results confirm the optimal convergence rate for the H^1 -error from Theorem 4.1.

J	$\max_{m=0,\dots,M} \ x(\cdot, t_m) - x_h^m\ _0$	EOC	$\max_{m=0,\dots,M} \ x(\cdot, t_m) - x_h^m\ _1$	EOC
32	1.2337e-02	—	2.8140e-01	—
64	3.1870e-03	1.95	1.4076e-01	1.00
128	8.0360e-04	1.99	7.0386e-02	1.00
256	2.0133e-04	2.00	3.5194e-02	1.00
512	5.0361e-05	2.00	1.7597e-02	1.00

Table 1. Errors for the convergence test for (6.5) over the time interval $[0, 0.45]$ for scheme (5.4) with the additional right-hand side $(f(t_{m+1}), \eta_h)^h$ from (6.6). We also display the experimental orders of convergence (EOC).

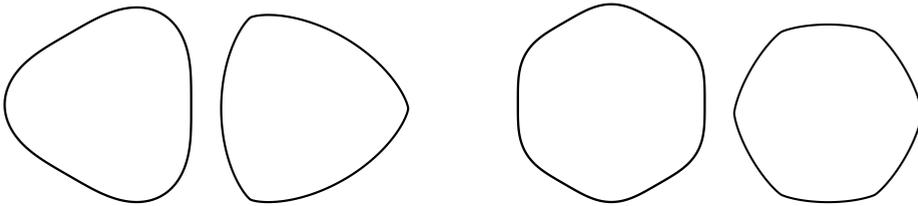


Figure 2. Frank diagram (left) and Wulff shape (right) for (6.7) with $(k, \delta) = (3, 0.124)$ and $(6, 0.028)$.

Next we consider smooth anisotropies as in [24, (7.1)] and [8, (4.4a)]. To this end, let

$$\gamma_0(p) = |p|(1 + \delta \cos(k\theta(p))), \quad p = |p| \begin{pmatrix} \cos \theta(p) \\ \sin \theta(p) \end{pmatrix}, \quad k \in \mathbb{N}, \delta \in \mathbb{R}_{\geq 0}. \quad (6.7)$$

It is not difficult to verify that this anisotropy satisfies (2.3) if and only if $\delta < \frac{1}{k^2-1}$; see also [8, p. 27]. In order to visualize this family of anisotropies, we show a Frank diagram and a Wulff shape for the cases $k = 3$ and $k = 6$, respectively, in Figure 2. We show the evolutions for anisotropic curve shortening flow induced by these two anisotropies, starting in each case from an equidistributed approximation of a unit circle, in Figure 3. Here we use the fully discrete scheme in (5.4). We observe from the evolutions in Figure 3 that the shape of the curve quickly approaches the Wulff shape, while it continuously shrinks towards a point. It is interesting to note that the ratio in (6.2) increases only slightly and then appears to remain nearly constant for the remainder of the evolution.

One motivation for choosing a sixfold anisotropy, as in (6.7) with $k = 6$, is its relevance for modeling ice crystal growth; see, for example, [7, 9]. Here it is desirable to choose a (nearly) crystalline anisotropy, which means that the Wulff shape exhibits flat sides and sharp corners. With the help of the class of anisotropies in (5.5), this is possible. We immediately demonstrate how this can be achieved for a general k -fold symmetry, for even $k \in \mathbb{N}$. On choosing $L = k/2$, we define the rotation matrix $Q(\theta) = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$

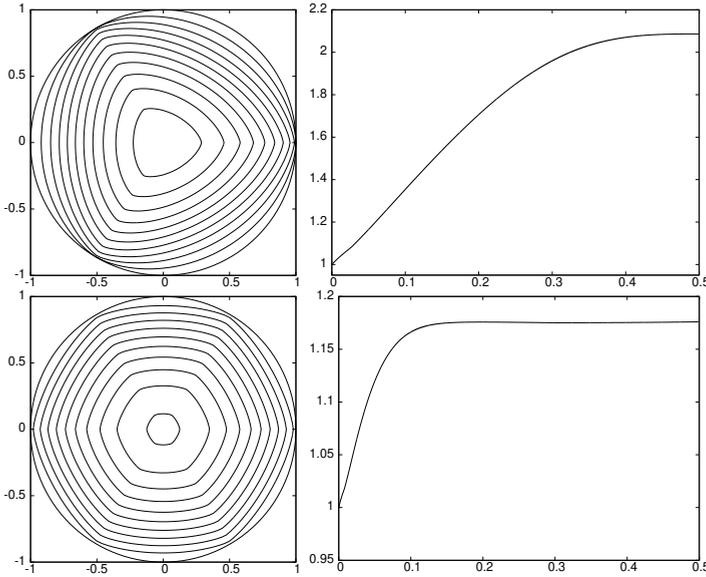


Figure 3. Anisotropic curvature flow for (6.7), with $(k, \delta) = (3, 0.124)$ (top) and $(6, 0.028)$ (bottom), using scheme (5.4). The solution at times $t = 0, 0.05, \dots, 0.5$ is shown. We also show plots of ratio (6.2) over time.

and the diagonal matrix $D(\delta) = \text{diag}(1, \delta^2)$, and then let

$$\gamma_0(p) = \sum_{\ell=1}^L \sqrt{\left[\left(Q\left(\frac{\pi}{L}\right)^\ell \right)^T D(\delta) \left(Q\left(\frac{\pi}{L}\right)^\ell \right) \right]} p \cdot p, \quad \delta \in \mathbb{R}_{>0}. \tag{6.8}$$

We visualize some Wulff shapes of (6.8) for $L = 2, 3, 4$ in Figure 4 and observe that these Wulff shapes, for $\delta \rightarrow 0$, will approach a square, a regular hexagon and a regular octagon, respectively. Of course, the associated crystalline anisotropic energy densities, when $\delta = 0$, are no longer differentiable and so, the theory developed in this paper no longer applies. Yet, for a fixed $\delta > 0$, all the assumptions in this paper are satisfied and our scheme (see (5.7)) works extremely well. As an example, we repeat the simulations in Figure 3 for anisotropy (6.8) with $L = 2$ and $\delta = 10^{-2}$, now using the scheme in (5.7). From the evolution shown in Figure 5, it can be seen that the initial curve assumes the shape of a smoothed square that then shrinks to a point. We also observe that after an initial increase, the ratio in (6.2) decreases and eventually reaches a steady state. The final distribution of mesh points is such that there is a slightly lower density of vertices on the nearly flat parts of the curve.

Inspired by the computations in [37, Fig. 6.1], we now consider evolutions for an initial curve that consists of a $\frac{3}{2}\pi$ -segment of the unit circle merged with parts of a square of side length 2. For our computations we employ the scheme in (5.7), with the discretization

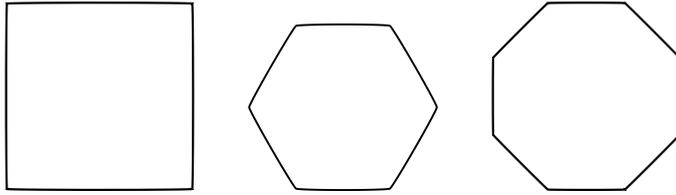


Figure 4. Wulff shapes (scaled) for (6.8) with $L = 2, 3, 4$ and $\delta = 10^{-2}$.

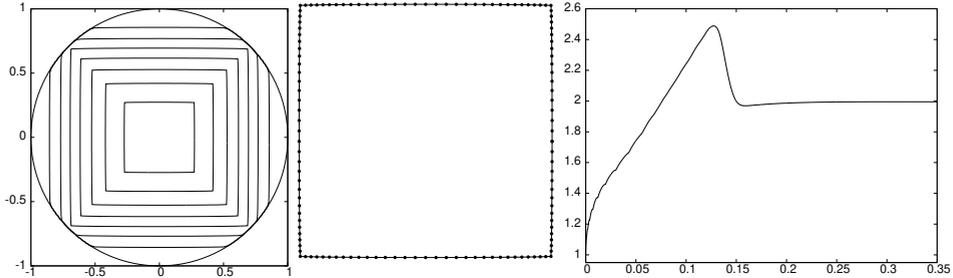


Figure 5. Anisotropic curvature flow for (6.8) with $L = 2$ and $\delta = 10^{-2}$, using scheme (5.7). We show the solution at times $t = 0, 0.05, \dots, 0.35$ and the distribution of vertices on $x_h^M(I)$, as well as the evolution of ratio (6.2) over time.

parameters $J = 256$ and $\Delta t = 10^{-4}$. The evolutions for the three anisotropies visualized in Figure 4 can be seen in Figure 6. We observe that the smooth part of the initial curve transitions into a crystalline shape, while the initial facets of the curve that are aligned with the Wulff shape simply shrink. The other facets disappear, some immediately and some over time, as they are replaced by facets aligned with the Wulff shape. The evolution of the left nonconvex corner in the initial curve is particularly interesting, as it shows three qualitatively very different behaviors for the three chosen anisotropies.

For the final simulations in this subsection, we choose as initial data a polygon that is very similar to the initial curve from [2, Fig. 0]. In their seminal work, Almgren and Taylor consider motion by crystalline curvature, which is the natural generalization of anisotropic curve shortening flow to purely crystalline anisotropies—that is, when the Wulff shape is a polygon. For motion by crystalline curvature, a system of ODEs for the sizes and positions of all the facets of an evolving polygonal curve has to be solved. Here the initial curve needs to be admissible in the sense that it only exhibits facets that also appear in the Wulff shape, and any two of its neighboring facets are also neighbors in the Wulff shape. Hence, the initial curve for the computations shown in Figure 7, for which we employed the scheme in (5.7) with $J = 512$ and $\Delta t = 10^{-4}$, is admissible for an eightfold anisotropy, with a regular octagon as Wulff shape. Our simulation for the smoothed anisotropy in (6.8) with $L = 4$ and $\delta = 10^{-4}$ agrees remarkably well with the evolution shown in [2, Fig. 0]. In fact, it is natural to conjecture that in the limit $\delta \rightarrow 0$, anisotropic curve shortening flow

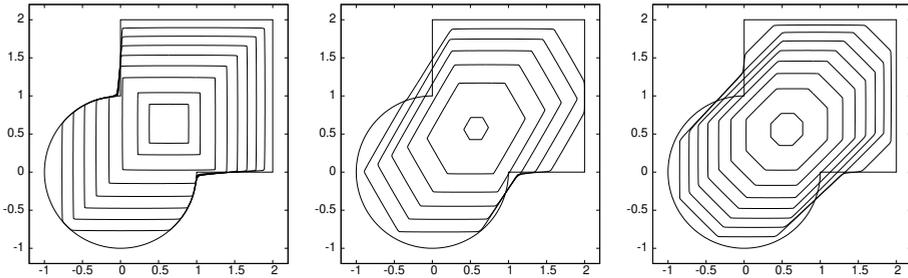


Figure 6. Anisotropic curvature flow for (6.8) with $L = 2, 3, 4$ and $\delta = 10^{-2}$, using scheme (5.7). We show the solution at times $t = 0, 0.1, \dots, 0.7, 0.75$ (left), $t = 0, 0.05, \dots, 0.3$ (middle), and $t = 0, 0.02, \dots, 0.16$ (right).

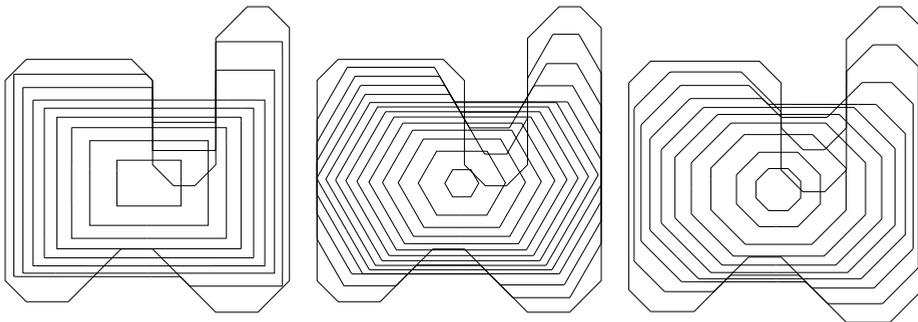


Figure 7. Anisotropic curvature flow for (6.8) with $L = 2, 3, 4$ and $\delta = 10^{-4}$, using scheme (5.7). We show the solution at times $t = 0, 2, \dots, 16$ (left), $t = 0, 0.5, \dots, 6, 6.4$ (middle) and $t = 0, 0.4, \dots, 3.2, 3.4$ (right).

for the anisotropies in (6.8) converges to flow by crystalline curvature with respect to the crystalline surface energies in (6.8) with $\delta = 0$. We stress that for the cases $L = 2$ and $L = 3$, when the Wulff shape is a square and a regular hexagon, respectively, the initial curve in Figure 7 is no longer admissible in the sense described above. As we only deal with the case $\delta > 0$, our scheme (see (5.7)) has no difficulties in computing the evolutions shown in Figure 7 for $L = 2$ and $L = 3$. We observe once again that new facets appear where the initial polygon is not aligned with the Wulff shape, while the admissible facets simply shrink.

6.2. Curve shortening flow in Riemannian manifolds

In this subsection, we consider the setup from Example 2.2(3); see also Example 3.4(3). At first we look at the simpler case of a manifold that is conformally flat, so that we can employ the scheme in (5.11). As an example, we take $G(z) = g(z)\text{Id}$ with $g(z) = (z_1)^{-2}$ and note that with $\Omega = \{z \in \mathbb{R}^2 : z_1 > 0\}$, we obtain a model for the hyperbolic plane,

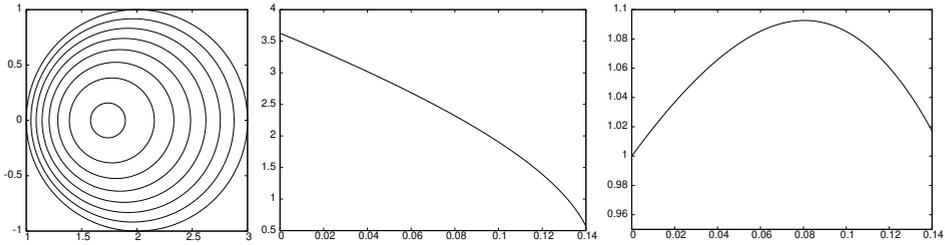


Figure 8. Curvature flow in the hyperbolic plane, using scheme (5.11). The solution at times $t = 0, 0.02, \dots, 0.14$ is shown. We also show plots of the discrete energy $\mathcal{E}^h(x_h^m)$ and of ratio (6.2) over time.

which is a two-dimensional manifold that cannot be embedded into \mathbb{R}^3 , as was proved by Hilbert [32]; see also [40, §11.1]. From [12, Appendix A], and on noting Lemma B.2, we recall that a true solution for (2.6)—that is, geodesic curvature flow in the hyperbolic plane—is given by a family of translating and shrinking circles in Ω :

$$\Gamma(t) = \begin{pmatrix} a(t) \\ 0 \end{pmatrix} + r(t)\mathbb{S}^1, \quad a(t) = e^{-t}a(0), \quad r(t) = (r^2(0) - a^2(0)[1 - e^{-2t}])^{\frac{1}{2}}, \quad (6.9)$$

with $a(0) > r(0) > 0$ and $\mathbb{S}^1 = \{z \in \mathbb{R}^2 : |z| = 1\}$. In Figure 8 we show such an evolution, starting from a unit circle centered at $\begin{pmatrix} 2 \\ 0 \end{pmatrix}$, computed with scheme (5.11) where, since g is convex in Ω , we choose $g_+ = g$. We observe that during the evolution the discrete geodesic length is decreasing, while the approximation to the shrinking circle remains nearly equidistributed throughout. At the final time $T = 0.14$, the maximum difference between $r(T)$ and $|x_h^M(q_j) - \begin{pmatrix} a(T) \\ 0 \end{pmatrix}|$, for $1 \leq j \leq J$, is less than $6 \cdot 10^{-3}$, indicating that the polygonal curve $\Gamma_h^M = x_h^M(I)$ is a very good approximation of the true solution $\Gamma(T)$ from (6.9).

For the remainder of this subsection, we consider general Riemannian manifolds that are not necessarily conformally flat. An example application is the modeling of geodesic curvature flow on a hypersurface in \mathbb{R}^3 that is given by a graph. In particular, we assume that

$$F(z) = (z_1, z_2, \varphi(z))^T, \quad \varphi \in C^3(\Omega). \quad (6.10)$$

The induced matrix G is then given by $G(z) = \text{Id} + \nabla\varphi(z) \otimes \nabla\varphi(z)$, and the splitting given by (5.8) for scheme (5.10) can be defined by $G_+(z) = G(z) + c_\varphi|z|^2\text{Id}$ and $G_-(z) = -c_\varphi|z|^2\text{Id}$, with $c_\varphi \in \mathbb{R}_{\geq 0}$ chosen sufficiently large. In all our computations, we observed a monotonically decreasing discrete energy when choosing $c_\varphi = 0$, and so we always let $G_+ = G$.

We begin with a convergence experiment on the right circular cone defined by $\varphi(z) = b|z|$ and $\Omega = \mathbb{R}^2 \setminus \{0\}$ in (6.10), for some $b \in \mathbb{R}_{\geq 0}$. A simple calculation verifies that the family of curves $\tilde{\Gamma}(t) = r(t)(\mathbb{S}^1 \times \{b\})$, with $r(t) = [r^2(0) - \frac{2t}{1+b^2}]^{\frac{1}{2}}$ and $r(0) > 0$, evolves under geodesic curvature flow on $\mathcal{M} = F(\Omega)$. In fact, it is not difficult

J	$\max_{m=0,\dots,M} \ x(\cdot, t_m) - x_h^m\ _0$	EOC	$\max_{m=0,\dots,M} \ x(\cdot, t_m) - x_h^m\ _1$	EOC
32	1.6096e-02	—	3.5595e-01	—
64	4.2080e-03	1.94	1.7805e-01	1.00
128	1.0635e-03	1.98	8.9032e-02	1.00
256	2.6656e-04	2.00	4.4517e-02	1.00
512	6.6685e-05	2.00	2.2259e-02	1.00
1024	1.6674e-05	2.00	1.1129e-02	1.00

Table 2. Errors for the convergence test for (6.11), with $b = \sqrt{3}$ and $r(0) = 1$, over the time interval $[0, \frac{1}{2}]$ for scheme (5.10) with $G_+ = G$.

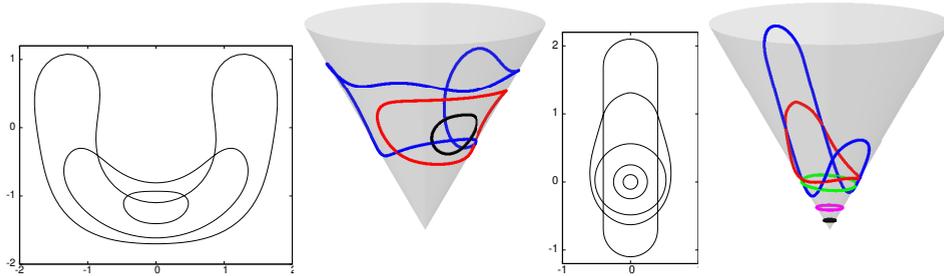


Figure 9. Geodesic curvature flow on the cone defined by (6.10) with $\varphi(z) = \sqrt{3}|z|$. We show the evolution of x_h^m in Ω , as well as of $F(x_h^m)$ on \mathcal{M} , at times $t = 0, 1, 1.8$ (left) and $t = 0, 0.2, 0.6, 1, 1.1$ (right).

to show that the particular parametrization

$$x(\rho, t) = [r^2(0) - \frac{2t}{1 + b^2}]^{\frac{1}{2}} \begin{pmatrix} \cos(2\pi\rho) \\ \sin(2\pi\rho) \end{pmatrix}, \tag{6.11}$$

so that $\tilde{\Gamma}(t) = F(x(I, t))$, solves (3.9). In a similar manner to Table 1, we report on the H^1 - and L^2 -errors between (6.11), for $b = \sqrt{3}$ and $r(0) = 1$, and the discrete solutions for scheme (5.10) in Table 2. Here, for a sequence of mesh sizes we use uniform time steps of size $\Delta t = h^2$, for $h = J^{-1} = 2^{-k}$, $k = 5, \dots, 9$. Once again, the observed numerical results confirm the optimal convergence rate from Theorem 4.1.

On the same cone \mathcal{M} , we perform two computations for a curve evolving by geodesic curvature flow. For the simulation on the left of Figure 9, it can be observed that as the initial curve $F(x_h^0(I))$ is homotopic to a point on \mathcal{M} , it shrinks to a point away from the apex. On recalling [29, Conjecture 5.1], due to Charles M. Elliott, on the right of Figure 9 we also show a numerical experiment for a curve that is not homotopic to a point on \mathcal{M} . According to the conjecture, any such curve should shrink to a point at the apex in finite time, and this is indeed what we observe.

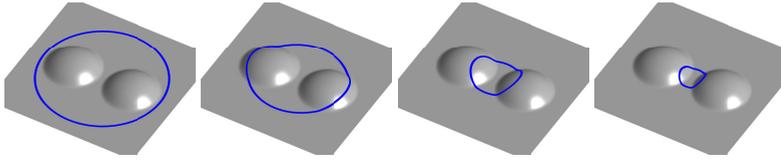


Figure 10. Geodesic curvature flow on the graph defined by (6.12) with $\lambda_1 = \lambda_2 = 1$. We show the evolution of $F(x_h^m)$ on \mathcal{M} at times $t = 0, 1, 2, 2.2$.

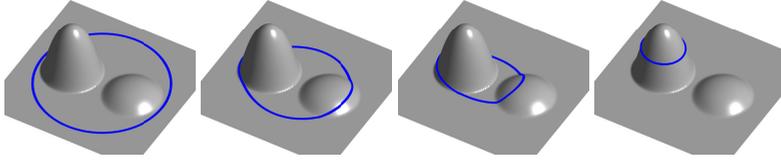


Figure 11. Geodesic curvature flow on the graph defined by (6.12) with $(\lambda_1, \lambda_2) = (5, 1)$. We show the evolution of $F(x_h^m)$ on \mathcal{M} at times $t = 0, 1, 2, 4$.

For the final set of numerical simulations, we model a surface with two mountains. Following [45], we define

$$\varphi(z) = \lambda_1 \psi(|z|^2) + \lambda_2 \psi(|z - \binom{2}{0}|^2), \quad \lambda_1, \lambda_2 \in \mathbb{R}_{\geq 0},$$

$$\text{where } \psi(s) = \begin{cases} e^{-\frac{1}{1-s}} & s < 1, \\ 0 & s \geq 1, \end{cases} \tag{6.12}$$

and let $\Omega = \mathbb{R}^2$. We show three evolutions for geodesic curvature flow on such surfaces in Figures 10, 11, and 12. In each case we start the evolution from an equidistributed approximation of a circle of radius 2 in Ω , centered at the origin. In the first two simulations, the curve manages to continuously decrease its length in \mathbb{R}^3 , until it shrinks to a point. To achieve this in the second example, the curve needs to “climb up” the higher mountain. However, in the final example the two mountains are too steep, and so the curve can no longer decrease its length by climbing higher. In fact, the curve approaches a steady state for the flow, that is, a geodesic on \mathcal{M} : a curve with vanishing geodesic curvature. The plot of the discrete energy in Figure 12 confirms that the evolution is approaching a geodesic.

A. First variation of the anisotropic energy

Proof of Lemma 2.1. Abbreviating $\tilde{\gamma}(z, p) = a(z)\gamma(z, p)$, $(z, p) \in \Omega \times \mathbb{R}^2$, we temporarily write \mathcal{E} in (1.1) as

$$\mathcal{E}(\Gamma) = \int_{\Gamma} \tilde{\gamma}(\cdot, \nu) \, d\mathcal{H}^1.$$

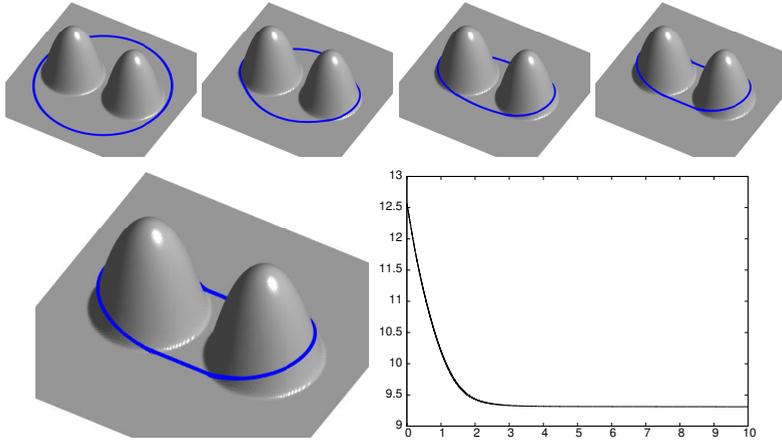


Figure 12. Geodesic curvature flow on the graph defined by (6.12) with $\lambda_1 = \lambda_2 = 5$. We show the evolution of $F(x_h^m)$ on \mathcal{M} at times $t = 0, 1, 2, 4$. Below we show a plot of $F(x_h^M)$ on \mathcal{M} , as well as a plot of the discrete energy $\mathcal{E}^h(x_h^m)$ over time.

Let us fix a curve $\Gamma \subset \Omega$ and a smooth vector field V defined in an open neighborhood of Γ . We infer from [22, Corollary 4.3] and (2.2) that the first variation of $\mathcal{E}(\Gamma)$ in the direction V is given by

$$\begin{aligned}
 d\mathcal{E}(\Gamma; V) &= \int_{\Gamma} ((\tilde{\gamma}(\cdot, \nu) - \tilde{\gamma}_p(\cdot, \nu) \cdot \nu)\kappa + \partial_\nu \tilde{\gamma}(\cdot, \nu) + \operatorname{div}_{\Gamma} \tilde{\gamma}_p(\cdot, \nu) \\
 &\quad + \tilde{\gamma}_{pp}(\cdot, \nu) : \nabla_{\Gamma} \nu) V \cdot \nu \, d\mathcal{H}^1 \\
 &= \int_{\Gamma} (\partial_\nu \tilde{\gamma}(\cdot, \nu) + \operatorname{div}_{\Gamma} \tilde{\gamma}_p(\cdot, \nu) + \tilde{\gamma}_{pp}(\cdot, \nu) : \nabla_{\Gamma} \nu) V \cdot \nu \, d\mathcal{H}^1. \tag{A.1}
 \end{aligned}$$

Here we note that the differential operators $\partial_\nu f = f_{z_i} \nu_i$ and $\operatorname{div}_{\Gamma} f = f_{i,z_i} - f_{i,z_j} \nu_j \nu_i$ on Γ only act on the first variable of functions defined in $\Omega \times \mathbb{R}^2$. In addition, we observe that the Weingarten map $\nabla_{\Gamma} \nu$ is given by $\nabla_{\Gamma} \nu = -\kappa \tau \otimes \tau$. We then calculate, on noting (2.2), that

$$\begin{aligned}
 \partial_\nu \tilde{\gamma}(\cdot, \nu) &= \partial_\nu a \gamma(\cdot, \nu) + a \gamma_{z_i}(\cdot, \nu) \nu_i, \\
 \operatorname{div}_{\Gamma} \tilde{\gamma}_p(\cdot, \nu) &= \operatorname{div}_{\Gamma} (a \gamma_p(\cdot, \nu)) \\
 &= a (\gamma_{p_i z_i}(\cdot, \nu) - \gamma_{p_i z_j}(\cdot, \nu) \nu_i \nu_j) + (a_{z_i} - \partial_\nu a \nu_i) \gamma_{p_i}(\cdot, \nu) \\
 &= a (\gamma_{p_i z_i}(\cdot, \nu) - \gamma_{z_j}(\cdot, \nu) \nu_j) + \nabla a \cdot \gamma_p(\cdot, \nu) - \partial_\nu a \gamma(\cdot, \nu), \\
 \tilde{\gamma}_{pp}(\cdot, \nu) : \nabla_{\Gamma} \nu &= -a \gamma_{pp}(\cdot, \nu) : \kappa \tau \otimes \tau = -a \kappa \gamma_{pp}(\cdot, \nu) \tau \cdot \tau.
 \end{aligned}$$

If we insert the above relations into (A.1) and recall (2.5), we obtain

$$d\mathcal{E}(\Gamma; V) = - \int_{\Gamma} \kappa_{\gamma} V \cdot \nu \, a \, d\mathcal{H}^1 = - \int_{\Gamma} \kappa_{\gamma} V \cdot \nu_{\gamma} \gamma(\cdot, \nu) a \, d\mathcal{H}^1,$$

which is (2.4). ■

B. Geodesic curve shortening flow in Riemannian manifolds

In this appendix we prove the claims formulated at the end of Example 2.2(3). Here we will make use of standard concepts in Riemannian geometry, and we refer the reader to the textbook [33] for further details.

Let $F : \Omega \rightarrow \mathcal{M}$ be a local parametrization of a two-dimensional Riemannian manifold (\mathcal{M}, g) and denote by $\{\partial_1, \partial_2\}$ the corresponding basis of the tangent space $T_{F(z)}\mathcal{M}$, for $z \in \Omega$. We also let $g_{ij}(z) = g_{F(z)}(\partial_i, \partial_j)$, $G(z) = (g_{ij}(z))_{i,j=1}^2$, $(g^{ij}(z))_{i,j=1}^2 = G^{-1}(z)$, $\gamma(z, p) = \sqrt{G^{-1}(z)p \cdot p}$, and $a(z) = \sqrt{\det G(z)}$, for $z \in \Omega$ and $p \in \mathbb{R}^2$, which induces energy equivalence (2.10). Let $\tilde{\Gamma}$ be a smooth curve in \mathcal{M} with unit tangent τ_g and a unit normal ν_g such that $\{\tau_g, \nu_g\}$ is an orthonormal basis of the tangent space $T\mathcal{M}$, that is, $g(\tau_g, \tau_g) = g(\nu_g, \nu_g) = 1$ and $g(\tau_g, \nu_g) = 0$. Then, the geodesic curvature κ_g of $\tilde{\Gamma}$ is defined by

$$\kappa_g = g\left(\frac{D}{ds}\tau_g, \nu_g\right) \quad \text{on } \tilde{\Gamma}, \tag{B.1}$$

where $\frac{D}{ds}\tau_g$ is the covariant derivative of τ_g .

Lemma B.1. *Let $\Gamma \subset \Omega$ be a smooth curve. Then, the anisotropic curvature of Γ and the geodesic curvature of $\tilde{\Gamma} = F(\Gamma)$ coincide in the sense that $\kappa_g \circ F = \kappa_{\gamma}$ on Γ .*

Proof. Let $\Gamma = x(I)$ for a parametrization $x : I \rightarrow \Omega$, so that $\tilde{\Gamma} = \tilde{x}(I)$ for $\tilde{x} = F \circ x$. Denoting by \tilde{s} the arc length of \tilde{x} , we see that $\tau_g = \tilde{x}_{\tilde{s}}$ and $\nu_g = \frac{1}{\gamma(x, \nu)} g^{ij}(x) \nu_j \partial_i$ form an orthonormal basis of $T_{\tilde{x}}\mathcal{M}$. Using the formula in [33, Lemma 5.1.2], we may write

$$\frac{D}{ds}\tau_g = \frac{D}{d\tilde{s}}\tilde{x}_{\tilde{s}} = (x_{k, \tilde{s}\tilde{s}} + \Gamma_{ij}^k(x) x_{i, \tilde{s}} x_{j, \tilde{s}}) \partial_k, \tag{B.2}$$

where $(\Gamma_{ij}^k(x))_{i,j,k=1}^2$ are the Christoffel symbols of \mathcal{M} at $F(x)$. Since $\partial_{\tilde{s}} = [G(x)x_{\rho} \cdot x_{\rho}]^{-\frac{1}{2}} \partial_{\rho}$, (B.1), (B.2), and (3.2) imply

$$\begin{aligned} \kappa_g \circ \tilde{x} &= g_{\tilde{x}}\left(\frac{D}{ds}\tau_g, \nu_g\right) = g_{kr}(x)(x_{k, \tilde{s}\tilde{s}} + \Gamma_{ij}^k(x) x_{i, \tilde{s}} x_{j, \tilde{s}}) \frac{1}{\gamma(x, \nu)} g^{lr}(x) \nu_l \\ &= \frac{1}{\gamma(x, \nu)} (x_{\tilde{s}\tilde{s}} \cdot \nu + \Gamma_{ij}^k(x) x_{i, \tilde{s}} x_{j, \tilde{s}} \nu_k) = \frac{1}{\gamma(x, \nu)} \frac{x_{\rho\rho} \cdot \nu + \Gamma_{ij}^k(x) x_{i, \rho} x_{j, \rho} \nu_k}{G(x)x_{\rho} \cdot x_{\rho}} \\ &= \frac{1}{\gamma(x, \nu)} \frac{\kappa + \Gamma_{ij}^k(x) \tau_i \tau_j \nu_k}{G(x)\tau \cdot \tau}. \end{aligned} \tag{B.3}$$

On the other hand, on recalling $\gamma(z, p) = \sqrt{G^{-1}(z)p \cdot p}$ and $a(z) = \sqrt{\det G(z)}$, we observe that

$$\gamma_p(z, p) = \frac{G^{-1}(z)p}{\gamma(z, p)}, \quad \gamma_{pp}(z, p) = \frac{G^{-1}(z)}{\gamma(z, p)} - \frac{G^{-1}(z)p \otimes G^{-1}(z)p}{\gamma^3(z, p)}, \quad (\text{B.4a})$$

$$\gamma_{pz_j}(z, p) = \frac{(G^{-1})_{z_j}(z)p}{\gamma(z, p)} - \frac{1}{2} \frac{(G^{-1})_{z_j}(z)p \cdot p}{\gamma^3(z, p)} G^{-1}(z)p, \quad (\text{B.4b})$$

$$a_{z_j}(z) = \frac{1}{2} \text{tr}(G^{-1}(z)G_{z_j}(z))a(z). \quad (\text{B.4c})$$

We infer from (B.4a) that

$$\begin{aligned} \gamma_{pp}(x, v)\tau \cdot \tau &= \frac{1}{\gamma^3(x, v)} [(G^{-1}(x)v \cdot v)(G^{-1}(x)\tau \cdot \tau) - (G^{-1}(x)v \cdot \tau)^2] \\ &= \frac{\det G^{-1}(x)}{\gamma^3(x, v)}. \end{aligned} \quad (\text{B.5})$$

For ease of notation, we drop the dependencies on x from now on. It is well known that

$$g_{kl,z_i} = g_{kr}\Gamma_{il}^r + g_{lr}\Gamma_{ik}^r, \quad i, k, l = 1, 2. \quad (\text{B.6})$$

Combining (B.6) with the relation $(G^{-1})_{z_i} = -G^{-1}G_{z_i}G^{-1}$, we find that

$$\begin{aligned} [(G^{-1})_{z_i}v]_j &= -g^{jk}g_{kl,z_i}g^{lm}v_m = -g^{jk}g^{lm}(g_{kr}\Gamma_{il}^r + g_{lr}\Gamma_{ik}^r)v_m \\ &= -g^{lm}\Gamma_{il}^j v_m - g^{jk}\Gamma_{ik}^m v_m, \end{aligned}$$

as well as

$$(G^{-1})_{z_i}v \cdot v = [(G^{-1})_{z_i}v]_j v_j = -(g^{lm}\Gamma_{il}^j v_m + g^{jk}\Gamma_{ik}^m v_m)v_j = -2g^{jk}\Gamma_{ik}^m v_m v_j.$$

If we insert the above relations into (B.4b), we obtain

$$\gamma_{pz_j}(\cdot, v) = -\frac{g^{lm}\Gamma_{jl}^j v_m + g^{jk}\Gamma_{jk}^m v_m}{\gamma(\cdot, v)} + \frac{g^{lk}\Gamma_{jk}^m v_m v_l}{\gamma^3(\cdot, v)} g^{jr} v_r.$$

Next we infer with the help of (B.4c) that

$$\frac{a_{z_j}}{a} \gamma_{pz_j}(\cdot, v) = \frac{1}{2} \frac{g^{kl}g_{kl,z_j}g^{jr} v_r}{\gamma(\cdot, v)} = \frac{1}{2} \frac{g^{kl}g^{jr}(g_{kr}\Gamma_{jl}^r + g_{lr}\Gamma_{jk}^r)v_r}{\gamma(\cdot, v)} = \frac{g^{jr}\Gamma_{jk}^k v_r}{\gamma(\cdot, v)}.$$

As a result,

$$\begin{aligned} \gamma_{pz_j}(\cdot, v) + \frac{\nabla a}{a} \cdot \gamma_p(\cdot, v) &= -\frac{g^{jk}\Gamma_{jk}^m v_m}{\gamma(\cdot, v)} + \frac{g^{lk}g^{jr}\Gamma_{jk}^m v_m v_l v_r}{\gamma^3(\cdot, v)} \\ &= \frac{1}{\gamma^3(\cdot, v)} (g^{lk}g^{jr} - g^{jk}g^{lr})\Gamma_{jk}^m v_m v_l v_r. \end{aligned}$$

Clearly,

$$g^{lk}g^{jr} - g^{jk}g^{lr} = \begin{cases} (\det G)^{-1} & (l, k, j, r) = (1, 1, 2, 2), (2, 2, 1, 1), \\ -(\det G)^{-1} & (l, k, j, r) = (1, 2, 2, 1), (2, 1, 1, 2), \\ 0 & \text{otherwise,} \end{cases}$$

so that

$$\begin{aligned} \gamma_{p_j z_j}(\cdot, \nu) + \frac{\nabla a}{a} \cdot \gamma_p(\cdot, \nu) &= -\frac{\det G^{-1}}{\gamma^3(\cdot, \nu)} (\Gamma_{22}^m \nu_m \nu_1^2 + \Gamma_{11}^m \nu_m \nu_2^2 - 2\Gamma_{12}^m \nu_m \nu_1 \nu_2) \\ &= -\frac{\det G^{-1} \Gamma_{kl}^m \nu_m \tau_k \tau_l}{\gamma^3(\cdot, \nu)}, \end{aligned}$$

since $\tau = -\nu^\perp$. Combining this relation with (2.5), (B.5), (B.3), and the fact that $\gamma^2(\cdot, \nu) = G^{-1} \nu \cdot \nu = (\det G^{-1}) G \tau \cdot \tau$, we finally obtain that

$$\begin{aligned} \kappa_\gamma \circ x &= \frac{\det G^{-1}}{\gamma^3(\cdot, \nu)} (\kappa + \Gamma_{kl}^m \tau_k \tau_l \nu_m) = \frac{1}{\gamma(\cdot, \nu)} \frac{\kappa + \Gamma_{kl}^m \tau_k \tau_l \nu_m}{G \tau \cdot \tau} \\ &= \kappa_g \circ \tilde{x} = (\kappa_g \circ F) \circ x \quad \text{in } I, \end{aligned}$$

as claimed. ■

A family of curves $(\tilde{\Gamma}(t))_{t \in [0, T]}$ in \mathcal{M} is said to evolve by geodesic curvature flow if

$$\mathcal{V}_g = \kappa_g \quad \text{on } \tilde{\Gamma}(t), \tag{B.7}$$

where \mathcal{V}_g is the normal velocity in the direction of the unit normal ν_g from the definition given in (B.1), that is, $\mathcal{V}_g = g(\tilde{x}_t \circ \tilde{x}^{-1}, \nu_g)$ with $\tilde{x} : I \times [0, T] \rightarrow \Omega$ being a parametrization of $(\tilde{\Gamma}(t))_{t \in [0, T]}$.

Lemma B.2. *Let $(\Gamma(t))_{t \in [0, T]}$ be a smooth family of curves in Ω . Then, anisotropic curve shortening flow for $(\Gamma(t))_{t \in [0, T]}$ in Ω (see (2.6)) is equivalent to geodesic curvature flow for $(F(\Gamma(t)))_{t \in [0, T]}$ in \mathcal{M} , (B.7).*

Proof. In a similar manner to the proof of Lemma B.1, we assume that $(\Gamma(t))_{t \in [0, T]}$ is parametrized by $x : I \times [0, T] \rightarrow \Omega$, so that $\tilde{x} = F \circ x$ parametrizes $(F(\Gamma(t)))_{t \in [0, T]}$. Let $\nu_g = \frac{1}{\gamma(x, \nu)} g^{ij}(x) \nu_j \partial_i$. Then, it follows from $\tilde{x}_t = x_{k,t} \partial_k$ that

$$\begin{aligned} (\mathcal{V}_g \circ F) \circ x &= \mathcal{V}_g \circ \tilde{x} = g_{\tilde{x}}(\tilde{x}_t, \tilde{\nu}) = \frac{1}{\gamma(x, \nu)} g^{ij}(x) \nu_j x_{k,t} g_{\tilde{x}}(\partial_k, \partial_i) \\ &= \frac{1}{\gamma(x, \nu)} g^{ij}(x) g_{ki}(x) \nu_j x_{k,t} = \frac{1}{\gamma(x, \nu)} x_t \cdot \nu \\ &= \mathcal{V}_\gamma \circ x \quad \text{in } I \times (0, T]. \end{aligned} \tag{B.8}$$

Combining (B.8) and Lemma B.1 yields the desired result. ■

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