Spectral Analysis of the Dirac Polaron

Dedicated to Professor Asao Arai on the occasion of his 60th birthday

by

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Abstract

A system of a Dirac particle interacting with the radiation field is considered. The Hamiltonian of the system is defined by $H = \boldsymbol{\alpha} \cdot (\hat{\mathbf{p}} - q\mathbf{A}(\hat{\mathbf{x}})) + m\beta + H_f$, where $q \in \mathbb{R}$ is a coupling constant, $\mathbf{A}(\hat{\mathbf{x}})$ the quantized vector potential and H_f the free photon Hamiltonian. Since the total momentum is conserved, H is decomposed with respect to the total momentum with fiber Hamiltonian $H(\mathbf{p})$ ($\mathbf{p} \in \mathbb{R}^3$). Since the self-adjoint operator $H(\mathbf{p})$ is bounded from below, one can define the lowest energy $E(\mathbf{p}, m) := \inf \sigma(H(\mathbf{p}))$. We prove that $E(\mathbf{p}, m)$ is an eigenvalue of $H(\mathbf{p})$ under the following conditions: (i) infrared regularization and (ii) $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$. We also discuss polarization vectors and the angular momentums.

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§1. Introduction

We consider a quantum system of a Dirac particle interacting with the radiation field. An example of a Dirac particle is the free electron. The Hilbert space for the Dirac particle is

(1.1)
$$\mathcal{H}_{\mathbf{p}} := L^2(\mathbb{R}^3_{\mathbf{x}}; \mathbb{C}^4),$$

and the free Hamiltonian for the Dirac particle is the free Dirac operator $\boldsymbol{\alpha} \cdot \hat{\mathbf{p}} + m\beta$ acting on \mathcal{H}_{p} , where $\hat{\mathbf{p}} = -i\nabla_{\mathbf{x}}$ denotes the momentum for the Dirac particle. The Hilbert space for the radiation field is the Fock space:

(1.2)
$$\mathcal{F}_{\mathrm{rad}} := \bigoplus_{n=0}^{\infty} \bigotimes_{\mathrm{sym}}^{n} L^{2}(\mathbb{R}_{\mathbf{k}}^{3} \times \{1, 2\}),$$

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where $\bigotimes_{\text{sym}}^n$ means the *n*-fold symmetric tensor product with $\bigotimes_{\text{sym}}^0 L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\}) := \mathbb{C}$. The Hilbert space for the total system is defined by

$$\mathcal{H} := \mathcal{H}_{p} \otimes \mathcal{F}_{rad}.$$

In this paper, we consider the quantum system described by the Hamiltonian

(1.4)
$$H := \boldsymbol{\alpha} \cdot (\hat{\mathbf{p}} - q\mathbf{A}(\hat{\mathbf{x}})) + m\beta + H_f,$$

where $q \in \mathbb{R}$ is a coupling constant, $\mathbf{A}(\hat{\mathbf{x}})$ denotes the quantized magnetic vector potential in the Coulomb gauge and H_f is the free photon Hamiltonian. We impose an ultraviolet cutoff in the quantized vector potential. We call the quantum system defined by (1.4) the Dirac-Maxwell model. The Hamiltonian (1.4) was introduced and discussed in the early days of quantum theory (see, e.g., [He]). By an informal perturbation theory, the Klein-Nishina formula (which gives a differential cross section for the Compton scattering) can be derived from the Dirac-Maxwell model [He]. Mathematical analysis of the Dirac-Maxwell model was initiated by A. Arai [A1, A2]. In [A3], A. Arai proved that a non-relativistic limit of the Dirac-Maxwell model converges to the Pauli-Fierz model (the non-relativistic QED). See also [A4]. The essential self-adjointness of the Hamiltonian (1.4) with an external potential was discussed by E. Stockmeyer and H. Zenk [SZ].

Since the Hamiltonian H is translation invariant, the total momentum of the system is conserved, i.e., the Hamiltonian of the system strongly commutes with the total momentum operator

(1.5)
$$\mathbf{P} := \hat{\mathbf{p}} + d\Gamma(\mathbf{k}),$$

where $d\Gamma(\mathbf{k})$ denotes the momentum operator of the radiation field. Hence the Hamiltonian can be decomposed as

(1.6)
$$H \cong \int_{\mathbb{R}^3}^{\oplus} H(\mathbf{p}) \, d\mathbf{p},$$

(1.7)
$$\mathbf{P} \cong \int_{\mathbb{P}^3}^{\oplus} \mathbf{p} \, d\mathbf{p},$$

where \cong means unitary equivalence. In this paper, we mainly study the fiber Hamiltonian $H(\mathbf{p})$ which describes the dynamics of the relativistic particle dressed in photons with total momentum \mathbf{p} . We call the quantum system described by $H(\mathbf{p})$ the *Dirac polaron*. As shown in [A2, A1], for $\mathbf{p} \in \mathbb{R}^3$, $H(\mathbf{p})$ has the form

(1.8)
$$H(\mathbf{p}) = \boldsymbol{\alpha} \cdot \mathbf{p} + m\beta + H_f - \boldsymbol{\alpha} \cdot d\Gamma(\mathbf{k}) - q\boldsymbol{\alpha} \cdot \mathbf{A},$$

which acts on $\mathbb{C}^4 \otimes \mathcal{F}_{rad}$, where **A** denotes the quantized vector potential at the origin (= **A**(**0**)). The fourth term $-\alpha \cdot d\Gamma(\mathbf{k})$ describes the reaction due to the

radiation field, and the last term $-q\boldsymbol{\alpha}\cdot\mathbf{A}$ is the electromagnetic interaction. It should be noted that $-q\boldsymbol{\alpha}\cdot\mathbf{A}$ is $not\ H(\mathbf{p})|_{q=0}$ -bounded for any non-zero q, because the reaction term $-\boldsymbol{\alpha}\cdot d\Gamma(\mathbf{k})$ is comparable to H_f , and $-q\boldsymbol{\alpha}\cdot\mathbf{A}$ is unbounded. This fact implies that $-q\boldsymbol{\alpha}\cdot\mathbf{A}$ is not a small perturbation no matter how small q is. One of the important facts about the Dirac polaron is that $H(\mathbf{p})$ is bounded from below for all values of all constants: the total momentum \mathbf{p} , the mass m and the coupling constant q (see [S1]). Hence, one can define the lowest energy by

(1.9)
$$E(\mathbf{p}, m) := \inf \sigma(H(\mathbf{p})) > -\infty,$$

where $\sigma(A)$ denotes the spectrum of A. If $H(\mathbf{p})$ has an eigenvalue E for $q \neq 0$, we say that a dressed particle state exists and the corresponding eigenvector is called a *dressed particle state*. In Section 4, we show that a dressed particle state exists under suitable conditions including (i) infrared regularization and (ii) the inequality

$$(1.10) E(\mathbf{p}, m) < E(\mathbf{p}, 0).$$

The condition (1.10) will be assumed in Theorems 4.1, 4.2 and 4.4 below. One can observe that there exist $m^* > 0$ such that (1.10) holds for all $|m| > m^*$. We expect that $m^* = 0$, but we have no proof. In Section 5, we study the angular momentum and degeneracy of eigenvalues of the Dirac polaron $H(\mathbf{p})$. We will show that the angular momentum of the \mathbf{p} -direction commutes with $H(\mathbf{p})$, and any eigenvalue of $H(\mathbf{p})$ has an even multiplicity (admitting infinity). Therefore $E(\mathbf{p}, m)$ is degenerate if it is an eigenvalue of $H(\mathbf{p})$.

This paper has three appendices. In Appendix A, we show that all spectral properties of the Dirac–Maxwell model and the Dirac polarons are independent of the choice of polarization vectors. Namely, two Hamiltonians defined by different polarization vectors are unitarily equivalent. The discussion in Appendix A is applicable to various QED models (e.g., Pauli–Fierz model). In Appendix B, we propose a general definition of angular momentum. Although the spectral properties of QED Hamiltonians are independent of the choice of polarization vectors, the definition of angular momentum depends on these vectors.

In Appendix C, we show some properties of the lowest energy $E(\mathbf{p})$ which are used in the proofs of Theorems 4.1–4.4.

§2. Definition of the model

In this paper, unless confusion may arise, we omit the symbol " \otimes " between two operators, for example, we write $A \otimes I$ as A and $I \otimes B$ as B, where I denotes the identity operator. For a closable operator T on $L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$, we denote

by $d\Gamma(T)$ and $\Gamma(T)$ the second quantization operators of T (see [RS2]), which act on $\mathcal{F}_{\rm rad}$. For $f \in L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$, we denote by a(f) and $a(f)^*$ the annihilation operator and the tre creation operator, respectively (see [RS2]), which are closed operators acting on $\mathcal{F}_{\rm rad}$. Let $\mathbf{e}^{(\lambda)} : \mathbb{R}^3 \to \mathbb{R}^3$, $\lambda = 1, 2$, be polarization vectors:

$$\mathbf{e}^{(\lambda)}(\mathbf{k}) \cdot \mathbf{e}^{(\mu)}(\mathbf{k}) = \delta_{\lambda \mu}, \quad \mathbf{e}^{(\lambda)}(\mathbf{k}) \cdot \mathbf{k} = \mathbf{0}, \quad \mathbf{k} \in \mathbb{R}^3, \ \lambda, \mu \in \{1, 2\}.$$

We write $\mathbf{e}^{(\lambda)}(\mathbf{k}) = (e_1^{(\lambda)}(\mathbf{k}), e_2^{(\lambda)}(\mathbf{k}), e_3^{(\lambda)}(\mathbf{k}))$, and we suppose that each component $e_j^{(\lambda)}(\mathbf{k})$ is a Borel measurable function of \mathbf{k} . For objects $\mathbf{a} = (a_1, a_2, a_3)$ and $\mathbf{b} = (b_1, b_2, b_3)$, we set $\mathbf{a} \cdot \mathbf{b} := \sum_{j=1}^3 a_j b_j$. For a linear $F(\cdot)$ we set $F(\mathbf{a}) := (F(a_1), F(a_2), F(a_3))$. Let ω be multiplication by the function

$$(2.1) \omega(\mathbf{k}) = |\mathbf{k}|.$$

We choose a function

$$\hat{\rho} \in L^2(\mathbb{R}^3_{\mathbf{k}}) \cap \mathrm{Dom}(\omega^{-1}),$$

where Dom means operator domain. For j = 1, 2, 3 and $\mathbf{x} \in \mathbb{R}^3$, we set

$$g_j(\mathbf{k}, \lambda; \mathbf{x}) := |\mathbf{k}|^{-1/2} \hat{\rho}(\mathbf{k}) e_i^{(\lambda)}(\mathbf{k}) e^{-i\mathbf{k}\cdot\mathbf{x}}, \quad (\mathbf{k}, \lambda) \in \mathbb{R}^3_{\mathbf{k}} \times \{1, 2\}.$$

For each fixed $\mathbf{x} \in \mathbb{R}^3$, the function $g_j(\mathbf{x})(\cdot) := g_j(\cdot; \mathbf{x})$ is in $L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1, 2\})$. The quantized magnetic vector potential at $\mathbf{x} \in \mathbb{R}^3$ is defined by

$$\mathbf{A}(\mathbf{x}) := (A_1(\mathbf{x})), A_2(\mathbf{x}), A_3(\mathbf{x})),$$

$$A_j(\mathbf{x}) := \frac{1}{\sqrt{2}} \overline{[a(g_j(\mathbf{x})) + a(g_j(\mathbf{x}))^*]}, \quad j = 1, 2, 3,$$

where, for a closable operator T, \bar{T} denotes its closure. For each $\mathbf{x} \in \mathbb{R}^3$, $A_j(\mathbf{x})$ is a self-adjoint operator on \mathcal{F}_{rad} (see [RS2]). Since $\mathbf{e}^{(\lambda)}(\mathbf{k})$'s are perpendicular to \mathbf{k} , the operators $\mathbf{A}(\mathbf{x})$ satisfy the Coulomb gauge condition

(2.3)
$$\operatorname{div} \mathbf{A}(\mathbf{x}) = \sum_{j=1}^{3} \partial_{x_{j}} A_{j}(\mathbf{x}) = 0.$$

Remark 2.1. The function $\hat{\rho}$ is called an *ultraviolet cutoff function*. A typical example of $\hat{\rho}$ is the characteristic function of the region $\{\mathbf{k} \in \mathbb{R}^3 \mid \kappa \leq |\mathbf{k}| \leq \Lambda\}$, where κ and Λ are non-negative constants. Here Λ is called an *ultraviolet cutoff*, and κ is an *infrared cutoff* if it is strictly positive.

The Hilbert space \mathcal{H} can be identified as

(2.4)
$$\mathcal{H} = L^2(\mathbb{R}^3_{\mathbf{x}}; \mathbb{C}^4 \otimes \mathcal{F}_{\text{rad}}) = \int_{\mathbb{R}^3}^{\oplus} \mathbb{C}^4 \otimes \mathcal{F}_{\text{rad}} d\mathbf{x}.$$

Under this identification, we define the quantized vector potential in the following way. Since $g_j(\mathbf{x}) \in L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$ is strongly continuous in $\mathbf{x} \in \mathbb{R}^3$, the map $\mathbf{x} \mapsto A_j(\mathbf{x})$ is a self-adjoint operator valued measurable function. Then we can define a self-adjoint operator on \mathcal{H} by

(2.5)
$$A_j(\hat{\mathbf{x}}) := \int_{\mathbb{R}^3}^{\oplus} A_j(\mathbf{x}) \, d\mathbf{x}.$$

Namely, when we identify $\Psi \in D(A_j(\hat{\mathbf{x}}))$ with an \mathcal{F}_{rad} -valued square integrable function, the action of the operator $A_j(\hat{\mathbf{x}})$ is given by $(A_j(\hat{\mathbf{x}})\Psi)(\mathbf{x}) = A_j(\mathbf{x})\Psi(\mathbf{x})$, $\mathbf{x} \in \mathbb{R}^3$. The operator valued vector

(2.6)
$$\mathbf{A}(\hat{\mathbf{x}}) := (A_1(\hat{\mathbf{x}}), A_2(\hat{\mathbf{x}}), A_3(\hat{\mathbf{x}}))$$

is also called the quantized vector potential.

The free photon Hamiltonian is the second quantization of ω :

$$(2.7) H_f := d\Gamma(\omega).$$

The Dirac-Maxwell Hamiltonian is defined by

(2.8)
$$H := \boldsymbol{\alpha} \cdot (\hat{\mathbf{p}} - q\mathbf{A}(\hat{\mathbf{x}})) + m\beta + H_f,$$

where $\hat{\mathbf{p}} = -i\nabla_{\mathbf{x}}$ and $\nabla_{\mathbf{x}}$ is the gradient operator acting in \mathcal{H}_{p} , $\boldsymbol{\alpha} = (\alpha_{1}, \alpha_{2}, \alpha_{3})$ and $\boldsymbol{\beta}$ are Dirac matrices satisfying $\alpha_{1}, \alpha_{2}, \alpha_{3}, \boldsymbol{\beta} \in M_{4}(\mathbb{C})$ and

(2.9)
$$\alpha_j \alpha_k + \alpha_k \alpha_j = 2\delta_{jk},$$

$$(2.10) \alpha_i \beta + \beta \alpha_i = 0,$$

$$\beta^2 = I_{\mathbb{C}^4},$$

the constant $m \in \mathbb{R}$ is the rest mass of the Dirac particle, and $q \in \mathbb{R}$ is a coupling constant. On the right hand side of (2.8), we omit the symbols $\otimes I$ and $I \otimes$, i.e., (2.8) is an abbreviation for

$$H = (\boldsymbol{\alpha} \cdot \hat{\mathbf{p}} + m\beta) \otimes I_{\mathcal{F}_{\text{rad}}} - q \sum_{i=1}^{3} (\alpha_{j} \otimes I_{L^{2}(\mathbb{R}_{\mathbf{x}}^{3})}) \cdot A_{j}(\hat{\mathbf{x}}) + I_{\mathcal{H}_{\mathbf{p}}} \otimes H_{f}.$$

In this paper, we use the Weyl representation for the Dirac matrices. Since all representations of the Dirac matrices are unitarily equivalent to each other, this choice does not affect the spectral properties of H (see [T, Lemma 2.25]).

It is easy to see that H is symmetric. Although the essential self-adjointness of H was proven in [A1], we give a slightly improved result:

Proposition 2.2 (Essential self-adjointness). \bar{H} is a self-adjoint operator and essentially self-adjoint on any core for $\sqrt{-\triangle} + H_f$.

Proof. The proof is a simple application of Nelson's commutator theorem. Our choice of a comparison operator for Nelson's commutator theorem is $\sqrt{-\triangle} + H_f$. See [S2] for details.

§3. Momentum conservation and fiber Hamiltonian $H(\mathbf{p})$

The total momentum operator is defined by

(3.1)
$$\mathbf{P} := \overline{\hat{\mathbf{p}} + d\Gamma(\mathbf{k})}.$$

The Hamiltonian H strongly commutes with \mathbf{P} (see [A1]). To construct the fiber Hamiltonian, we define a self-adjoint operator

$$(3.2) Q := \overline{\mathbf{x} \cdot d\Gamma(\mathbf{k})}.$$

Let U_F be the Fourier transform from $L^2(\mathbb{R}^3_{\mathbf{x}})$ to $L^2(\mathbb{R}^3_{\mathbf{p}})$. We set

$$(3.3) U := (U_F \otimes I_{\mathbb{C}^4}) \exp(iQ).$$

Then we can identify $U\mathcal{H}$ as a constant fiber direct integral

(3.4)
$$U\mathcal{H} \cong \int_{\mathbb{R}^3}^{\oplus} \mathbb{C}^4 \otimes \mathcal{F}_{\text{rad}} d\mathbf{p}.$$

For every $\mathbf{p} \in \mathbb{R}^3$, we define

(3.5)
$$H(\mathbf{p}) := \boldsymbol{\alpha} \cdot \mathbf{p} + m\beta + H_f - \boldsymbol{\alpha} \cdot d\Gamma(\mathbf{k}) - q\boldsymbol{\alpha} \cdot \mathbf{A},$$

which acts on $\mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}}$, where $\mathbf{A} := \mathbf{A}(\mathbf{0})$.

Proposition 3.1. For all $\mathbf{p} \in \mathbb{R}^3$, $H(\mathbf{p})$ is essentially self-adjoint and

(3.6)
$$U\bar{H}U^* = \int_{\mathbb{R}^3}^{\oplus} \overline{H(\mathbf{p})} \, d\mathbf{p},$$

(3.7)
$$U\mathbf{P}U^* = \int_{\mathbb{R}^3}^{\oplus} \mathbf{p} \, d\mathbf{p},$$

where $\int_{-\infty}^{\oplus} (\cdots)$ denotes the fiber direct integral operator with respect to (3.4).

Proof. See [A2].
$$\Box$$

Remark 3.2. Physically $\overline{H(\mathbf{p})}$ is the Hamiltonian of the fixed total momentum $\mathbf{p} \in \mathbb{R}^3$. One can show that the spectral properties of $\overline{H(\mathbf{p})}$ are independent of the choice of polarization vectors, because the Hamiltonians with different polarization vectors are unitarily equivalent. See Appendix A.

Remark 3.3. We call $H(\mathbf{p})$ the *Dirac polaron Hamiltonian*; it was introduced in [A4]. It is expected that, as in the model of the H. Fröhlich polaron, electromagnetic interaction forms a quasiparticle where the bare Dirac particle is surrounded by photon clouds. Such a quasiparticle with momentum $\mathbf{p} \in \mathbb{R}^3$ is considered as a ground state of $\overline{H}(\mathbf{p})$, if it exists. The existence of a ground state of $\overline{H}(\mathbf{p})$ is the main subject of our paper.

Remark 3.4. Note that $Dom(\alpha \cdot d\Gamma(\mathbf{k})) \subset Dom(H_f)$. Hence we have $Dom(H_f) = Dom(H(\mathbf{p}))$ and $H(\mathbf{p})$ is essentially self-adjoint on $Dom(H_f)$.

One of the most important properties of $\overline{H(\mathbf{p})}$ is semi-boundedness:

Theorem 3.5 ([S1]). For any \mathbf{p} , $\overline{H(\mathbf{p})}$ is bounded from below. Moreover $H(\mathbf{p})$ is essentially self-adjoint on any core for H_f .

Proof. The first statement was shown in [S1], where it is assumed that $\hat{\rho} \in \text{Dom}(\omega^{1/2})$, but one can remove this condition by the following procedure. In [S1, ineq. (24)], it is shown that $H(\mathbf{p})$ is bounded from below, and the lower bound is a function of $\|\omega^{1/2}\mathbf{g}\|_{L^2(\mathbb{R}^3)}$ and not $\|\omega\mathbf{g}\|_{L^2(\mathbb{R}^3)}$. Therefore, firstly, we regularize $\hat{\rho}$ as $\hat{\rho}_{\lambda}(\mathbf{k}) := \hat{\rho}(\mathbf{k})\chi_{|\mathbf{k}| \leq \lambda}$, and then we obtain the lower bound of the regularized Hamiltonian $H_{\lambda}(\mathbf{p}) \geq C_{\epsilon}$. Since C_{λ} converges as $\lambda \to \infty$ and $H_{\mathbf{p}}$ converges to $H(\mathbf{p})$ on a finite particle subspace, we get $H(\mathbf{p}) \geq \lim_{\epsilon \to +0} C_{\epsilon} > -\infty$. The second statement follows from Wüst's Theorem [RS2] and the bound

(3.8)
$$\|\boldsymbol{\alpha} \cdot (d\Gamma(\mathbf{k}) - q\mathbf{A})\Psi\|^2 \le \|(H_f + E)\Psi\|^2, \quad \Psi \in \text{Dom}(H_f),$$

for some E > 0. The bound (3.8) was given in [S1].

Thus we can define the lowest energy of the Dirac polaron with total momentum ${\bf p}$ by

(3.9)
$$E(\mathbf{p}, m) := \inf \sigma(\overline{H(\mathbf{p})}).$$

The energy $E(\mathbf{p}, m)$ depends on all parameters $(\mathbf{p}, m, q) \in \mathbb{R}^3 \times \mathbb{R} \times \mathbb{R}$. When the m-dependence in $E(\mathbf{p}, m)$ is not important, we write $E(\mathbf{p}, m)$ as $E(\mathbf{p})$.

§4. Existence of a ground state

For a self-adjoint operator bounded below, T, we say that T has a ground state if $\inf \sigma(T)$ is an eigenvalue of T. In this section, we give criteria for $\overline{H(\mathbf{p})}$ to have a ground state.

Theorem 4.1. Suppose that $\hat{\rho}$ is spherically symmetric and

(4.1)
$$\int_{\mathbb{R}^3} \frac{q^2}{(E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}|)^2} \frac{|\hat{\rho}(\mathbf{k})|^2}{|\mathbf{k}|} d\mathbf{k} < 1.$$

Assume that $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$. Then the Dirac polaron Hamiltonian $\overline{H(\mathbf{p})}$ has a ground state.

Using the lower bound on $E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}|$ which is proved in Theorem C.10 of Appendix C, we obtain the following result:

Theorem 4.2. Assume that $\hat{\rho}$ be spherically symmetric and $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$. Assume the infrared regularity condition $\hat{\rho} \in \text{Dom}(\omega^{-3/2})$ holds. Then there exists a constant $q_0 > 0$ such that for all q with $|q| < q_0$, $\overline{H(\mathbf{p})}$ has a ground state.

Remark 4.3. Since $E(\mathbf{p}, m)$ is concave in m (Proposition C.1) and since we have $\lim_{m\to\infty} E(\mathbf{p}, m) = -\infty$, there exists $m^* \geq 0$ such that $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$ for all $|m| > m^*$.

The proof of Theorem 4.1 is based on estimates of a photon number bound. The condition (4.1) can be considered as a restriction on the coupling constant q. There are two ways to remove this restriction. The first one is the method discovered by C. Gérard [Ge], and the other is the photon derivative bound developed in [GLL]. In this paper, we use the photon derivative bound. We need some additional assumptions:

(A) (i) $\hat{\rho}$ is a spherically symmetric function. (ii) There is an open set $S \subset \mathbb{R}^3$ such that $\bar{S} = \operatorname{supp} \hat{\rho}$ and $\hat{\rho}$ is continuously differentiable on S. (iii) For all R > 0, the bounded region $S_R := \{\mathbf{k} \in S \mid |\mathbf{k}| < R\}$ has the cone property (see [LL] for the definition).

The theorem below proves the existence of a ground state of the Dirac polaron for all values of the coupling constant q:

Theorem 4.4. Assume that condition (Λ) holds. Moreover assume that

$$(4.2) \quad \hat{\rho} \in \text{Dom}(\omega^{-3/2}), \quad |\mathbf{k}|^{-5/2} \hat{\rho}(\mathbf{k}) \in L^p(S_R), \quad |\mathbf{k}|^{-3/2} |\nabla \hat{\rho}(\mathbf{k})| \in L^p(S_R),$$

for all $p \in [1,2)$ and R > 0. Suppose that $E(\mathbf{p},m) < E(\mathbf{p},0)$. Then $\overline{H(\mathbf{p})}$ has a ground state.

Remark 4.5. We now give an example. Let $\chi_{\kappa,\Lambda}(\mathbf{k})$ be the characteristic function of the region $\{\mathbf{k} \in \mathbb{R}^3 \mid \kappa < |\mathbf{k}| < \Lambda\}$. For all $\kappa > 0$ and $\Lambda < \infty$, the cutoff function $\hat{\rho} = \chi_{\kappa,\Lambda}$ satisfies (Λ) and (4.2). The function $\hat{\rho}(\mathbf{k}) = |\mathbf{k}| \exp(-\lambda |\mathbf{k}|)$ $(\lambda > 0)$ also satisfies condition (Λ) and (4.2).

Remark 4.6. It is known that, in non-relativistic QED, the existence of a dressed particle requires the restriction $|\mathbf{p}|/m \leq 1$ (see [C]). On the other hand, Theorems 4.1–4.4 do not require a restriction on $|\mathbf{p}|/m$. This fact is a crucial difference between relativistic and non-relativistic dynamics. This result can be interpreted as follows. In general, the velocity operator is defined by $i = \sqrt{-1}$ times the commutator of the energy Hamiltonian with the position. Hence, the velocity operators of the non-relativistic particle and Dirac particle are defined by

$$\hat{\mathbf{p}}/m = i[\hat{\mathbf{p}}^2/2m, \mathbf{x}],$$

(4.4)
$$\boldsymbol{\alpha} = [\boldsymbol{\alpha} \cdot \hat{\mathbf{p}} + m\beta, \mathbf{x}],$$

respectively. Hence the non-relativistic particle can move faster than light, and the particle with velocity $|\mathbf{p}|/m > 1$ makes a shock wave of light and loses its kinetic energy. Therefore such a non-relativistic particle is unstable in the presence of electromagnetic interaction. On the other hand, since the speed of the Dirac particle is smaller than that of light, $\|\boldsymbol{\alpha}\| \leq 1$, this kind of catastrophe does not occur, and the dressed electron state is stable for all $|\mathbf{p}|$.

Remark 4.7. It is easy to see that the Hermitian matrix $\alpha \cdot \mathbf{p} + m\beta$ has two eigenvalues $\pm \sqrt{\mathbf{p}^2 + m^2}$, each of which is two-fold degenerate. Let $u_i^{(\pm)} \in \mathbb{C}^4$, i = 1, 2, be the corresponding normalized eigenvectors:

$$(\boldsymbol{\alpha} \cdot \mathbf{p} + m\beta)u_i^{(\pm)} = \pm \sqrt{\mathbf{p}^2 + m^2} u_i^{(\pm)}, \quad i = 1, 2.$$

Let $\Omega := (1,0,0,\ldots) \in \mathcal{F}_{\mathrm{rad}}$ be the vacuum. It is the unique eigenvector of both H_f and $d\Gamma(k_j), j=1,2,3$. We set $\Phi_i^{(\pm)} := u_i^{(\pm)} \otimes \Omega, j=1,2$. Clearly,

$$H(\mathbf{p})|_{q=0}\Phi_i^{(\pm)} = \pm \sqrt{\mathbf{p}^2 + m^2} \,\Phi_i^{(\pm)}, \quad i = 1, 2.$$

Thus, in the case q=0, $H(\mathbf{p})|_{q=0}$ has two eigenvalues $\pm\sqrt{\mathbf{p}^2+m^2}$. These eigenvectors $\Phi_i^{(+)}$, i=1,2 (resp. $\Phi_i^{(-)}$, i=1,2) describe states of a freely moving positive (resp. negative) energy particle with momentum \mathbf{p} . Hence, if photons and the Dirac particle are decoupled, a Dirac particle associated with a positive eigenvalue exists and the positive eigenvalue is embedded. We are interested in the fate of those eigenvalues when interaction is switched on. As is shown in Fig. 1, the lowest energy $E(\mathbf{p},m)$ converges to $-\sqrt{\mathbf{p}^2+m^2}$ as $q\to 0$. According to textbooks of physics (e.g. [B, He]), it is expected that any positive energy electron falls down to a negative energy state by a spontaneous emission of photons. Hence it is expected that the eigenvalue $+\sqrt{\mathbf{p}^2+m^2}$ is unstable under the perturbation $q\mathbf{\alpha}\cdot\mathbf{A}$. Theorems 4.1–4.4 ensure that a negative energy dressed electron exists under some conditions. But the instability of $\sqrt{\mathbf{p}^2+m^2}$ has not been proved yet.

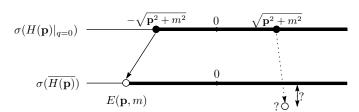


Figure 1. Spectrum of $H(\mathbf{p})|_{q=0}$ and $H(\mathbf{p})$.

§5. Angular momentum and degeneracy of eigenvalues

In this section we show that the angular momentum around the **j**-axis (where $\mathbf{j} \in \mathbb{R}^3 \setminus \{0\}$) of the Dirac polaron is conserved if **p** is parallel to **j** and $\hat{\rho}(\mathbf{k})$ has axial symmetry around **j**. Let $(\overline{H}(\mathbf{p}), \mathbf{e})$ be a Dirac polaron model with an arbitrarily given polarization vectors $\mathbf{e} = (\mathbf{e}^{(1)}, \mathbf{e}^{(2)})$. The total angular momentum around the **j**-axis in the system $(\overline{H}(\mathbf{p}), \mathbf{e})$ is defined by

$$J_{\mathbf{j}}(\mathbf{e}) := S_{\mathbf{j}} + L_{\mathbf{j}}(\mathbf{e}),$$

where $S_{\mathbf{j}} := \bigoplus^2 (\mathbf{j} \cdot \vec{\sigma})/2$, $\vec{\sigma} = (\sigma_1, \sigma_2, \sigma_3)$ are the Pauli matrices, and $L_{\mathbf{j}}(\mathbf{e})$ is an angular momentum for the radiation field, which is defined in Appendix B.

Proposition 5.1. The spectrum of $J_{\mathbf{j}}(\mathbf{e})$ is the set of half-integers:

$$\sigma(J_{\mathbf{i}}(\mathbf{e})) = \mathbb{Z}_{1/2} := \{\pm 1/2, \pm 3/2, \pm 5/2, \dots\}.$$

In particular, $J_{\mathbf{i}}(\mathbf{e})$ decomposes as

(5.1)
$$J_{\mathbf{j}}(\mathbf{e}) \cong \bigoplus_{z \in \mathbb{Z}_{1/2}} z$$

with respect to the identification

$$\mathbb{C}^4\otimes\mathcal{F}_{\mathrm{rad}}\congigoplus_{z\in\mathbb{Z}_{1/2}}\mathcal{F}(z).$$

We conclude this section with the following:

Theorem 5.2. Let \mathbf{j} be a unit vector parallel to \mathbf{p} . Assume that $\hat{\rho}(\mathbf{k}) = \hat{\rho}(R\mathbf{k})$, $\mathbf{k} \in \mathbb{R}^3$, for all $R \in O(3)$ with $R\mathbf{j} = \mathbf{j}$. Then $\overline{H(\mathbf{p})}$ strongly commutes with $J_{\mathbf{j}}(\mathbf{e})$. In particular, $\overline{H(\mathbf{p})}$ decomposes as

$$\overline{H(\mathbf{p})} \cong \bigoplus_{z \in \mathbb{Z}_{1/2}} H(\mathbf{p}:z),$$

corresponding to the decomposition (5.1). Moreover, for all $z \in \mathbb{Z}_{1/2}$, $H(\mathbf{p} : z)$ is unitarily equivalent to $H(\mathbf{p} : -z)$, and the multiplicity of any eigenvalue of $\overline{H(\mathbf{p})}$ is even.

Remark 5.3. In [Hi], F. Hiroshima defines an angular momentum in QED which differs from our definition.

§6. Proof of Theorems 4.1–4.4

For a constant $\nu \geq 0$, we define a regularized Hamiltonian to avoid the risk of infrared divergence:

(6.1)
$$H_{\nu}(\mathbf{p}) := \boldsymbol{\alpha} \cdot \mathbf{p} + m\beta + H_{f}(\nu) - \boldsymbol{\alpha} \cdot d\Gamma(\mathbf{k}) - q\boldsymbol{\alpha} \cdot \mathbf{A},$$

where

(6.2)
$$H_f(\nu) := d\Gamma(\omega_{\nu}), \quad \omega_{\nu}(\mathbf{k}) = (1+\nu)|\mathbf{k}| + \nu.$$

Let $N_f := d\Gamma(1)$ be the photon number operator. Note that we have $H_f(\nu) = H_f + \nu(H_f + N_f)$ and $H_0(\mathbf{p}) = H(\mathbf{p})$. By the Kato-Rellich theorem, one can easily show that, for all $\nu > 0$, $H_{\nu}(\mathbf{p})$ is self-adjoint on $\text{Dom}(H_f(\nu))$, and essentially self-adjoint on any core for $H_f(\nu)$. Since $H_{\nu}(\mathbf{p}) \geq H(\mathbf{p}) H_{\nu}(\mathbf{p})$ is also bounded from below. We set $\mathcal{D} := \text{Dom}(H_f) \cap \text{Dom}(N_f)$. Then \mathcal{D} is a common core for $\overline{H_{\nu}(\mathbf{p})}$ ($\nu \geq 0$). We set

(6.3)
$$E_{\nu}(\mathbf{p}) := \inf \sigma(\overline{H_{\nu}(\mathbf{p})}).$$

For $\nu > 0$, the massive Hamiltonian $H_{\nu}(\mathbf{p})$ was studied in [A1, A2], where A. Arai showed that $H_{\nu}(\mathbf{p})$ has a ground state for all $\nu > 0$.

Lemma 6.1 (Existence of a ground state for $\nu > 0$). Assume that $\nu > 0$. Then

(6.4)
$$\inf \sigma_{\text{ess}}(H_{\nu}(\mathbf{p})) - E_{\nu}(\mathbf{p}) \ge \nu.$$

In particular, $H_{\nu}(\mathbf{p})$ has a ground state.

Proof. See [A2].
$$\Box$$

By Lemma 6.1, for all $\nu > 0$, $H_{\nu}(\mathbf{p})$ has a normalized ground state $\Phi_{\nu}(\mathbf{p}) \in \text{Dom}(H_f(\nu))$. In the following, we construct a ground state of $H_0(\mathbf{p})$ as a suitable limit of $\Phi_{\nu}(\mathbf{p})$. Since $\Phi_{\nu}(\mathbf{p})$ is normalized, there exists a sequence $\{\Phi_{\nu_j}(\mathbf{p})\}_{j=1}^{\infty}$ with $\lim_{j\to\infty} \nu_j = 0$ such that $\{\Phi_{\nu_j}\}_j$ has a weak limit.

Lemma 6.2. Let $\{\nu_j\}_{j=1}^{\infty}$ be a sequence such that Φ_{ν_j} has a weak limit $\Phi_0(\mathbf{p}) := \text{w-lim}_{j\to\infty} \Phi_{\nu_j}$. Assume $\Phi_0 \neq 0$. Then $\Phi_0 \in \text{Dom}(\overline{H(\mathbf{p})})$ and Φ_0 is a ground state of $\overline{H(\mathbf{p})}$.

Proof. For all $\Psi \in \mathcal{D}$, one has

(6.5)
$$\langle H(\mathbf{p})\Psi, \Phi_0 \rangle = \lim_{j \to \infty} \langle \Psi, H(\mathbf{p})\Phi_{\nu_j} \rangle = \lim_{j \to \infty} \langle \Psi, \{E_{\nu_j}(\mathbf{p}) - \nu_j(H_f + N_f)\}\Phi_{\nu_j} \rangle.$$

By Proposition C.9, we have $E_{\nu_j}(\mathbf{p}) \to E_0(\mathbf{p})$ as $j \to \infty$. By assumption (2), we have

(6.6)
$$\lim_{j \to \infty} \nu_j |\langle \Psi, (H_f + N_f) \Phi_{\nu_j} \rangle| \le \lim_{j \to \infty} \nu_j ||(H_f + N_f) \Psi|| \cdot ||\Phi_{\nu_j}|| = 0.$$

Hence $\langle H(\mathbf{p})\Psi, \Phi_0 \rangle = \langle \Psi, E(\mathbf{p})\Phi_0 \rangle$ for all $\Psi \in \mathcal{D}$. Since \mathcal{D} is a core for $\overline{H(\mathbf{p})}$, we have $\Phi_0 \in \text{Dom}(\overline{H(\mathbf{p})})$ and $\overline{H(\mathbf{p})}\Phi_0 = E(\mathbf{p})\Phi_0$.

 $E_{\nu}(\mathbf{p})$ and $H_{\nu}(\mathbf{p})$ depend on \mathbf{p}, m, ν , etc. When we need to indicate such dependence, we write $E_{\nu}(\mathbf{p}, m, \dots)$ and $H_{\nu}(\mathbf{p}, m, q, \dots)$.

In this section, we use the identification

$$\mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}} = \bigoplus_{n=0}^{\infty} \mathbb{C}^4 \otimes \mathcal{F}^{(n)}, \quad \mathcal{F}^{(n)} := \bigotimes_{\mathrm{sym}}^n L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1, 2\}),$$

and each vector $\Psi^{(n)} \in \mathbb{C}^4 \otimes \mathcal{F}^{(n)}$ is identified with a Hilbert space valued function $\Psi^{(n)}(\mathbf{k}, \lambda; \cdot) : \mathbb{R}^3_{\mathbf{k}} \times \{1, 2\} \to \mathbb{C}^4 \otimes \mathcal{F}^{(n-1)}$. For all $(\mathbf{k}, \lambda) \in \mathbb{R}^3 \times \{1, 2\}$, we define a map

$$(6.7) a_{\lambda}(\mathbf{k}) : \mathbb{C}^{4} \otimes \mathcal{F}_{\mathrm{rad}} \to \prod_{n=0}^{\infty} \mathbb{C}^{4} \otimes \mathcal{F}^{(n)} := \{ (\Phi^{(n)})_{n=0}^{\infty} \mid \Phi^{(n)} \in \mathbb{C}^{4} \otimes \mathcal{F}^{(n)} \}$$

by

$$a_{\lambda}(\mathbf{k})\Psi := (\Psi^{(1)}(\mathbf{k},\lambda), \sqrt{2}\,\Psi^{(2)}(\mathbf{k},\lambda;\cdot), \dots, \sqrt{n}\,\Psi^{(n)}(\mathbf{k},\lambda;\cdot), \dots) \in \prod_{n=0}^{\infty} \mathbb{C}^4 \otimes \mathcal{F}^{(n)}.$$

For almost every (\mathbf{k}, λ) , $a_{\lambda}(\mathbf{k})$ is well-defined as a linear map. The smeared annihilation operator a(f) formally satisfies

(6.9)
$$a(f)\Psi = \sum_{\lambda=1,2} \int_{\mathbb{R}^3} d\mathbf{k} f(\mathbf{k}, \lambda)^* a_{\lambda}(\mathbf{k}) \Psi.$$

It is not necessary to consider $a_{\lambda}(\mathbf{k})$ as an operator valued distribution. This definition of $a_{\lambda}(\mathbf{k})$ is useful for our purpose below (Proposition 6.3). In general, $a_{\lambda}(\mathbf{k})\Psi \notin \mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}}$, but one can show that $a_{\lambda}(\mathbf{k})\Psi \in \mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}}$ for a class of vectors $\Psi \in \mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}}$. Let $\mathbf{w} : \mathbb{R}^3 \to [0, \infty)$ be an almost positive Borel measurable function. Then, for any $\Psi \in \mathrm{Dom}(d\Gamma(\mathbf{w})^{1/2})$ and for almost every $(\mathbf{k}, \lambda) \in \mathbb{R}^3 \times \{1, 2\}$, the vector $a_{\lambda}(\mathbf{k})\Psi$ is a $\mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}}$ -valued function, because, for any $\Psi \in \mathrm{Dom}(d\Gamma(\mathbf{w})^{1/2})$, one has

$$(6.10) \qquad \|d\Gamma(\mathbf{w})^{1/2}\Psi\|^2 = \sum_{n=1}^{\infty} \sum_{\lambda=1,2} \int_{\mathbb{R}^3} d\mathbf{k} \, \mathbf{w}(\mathbf{k}) n \|\Psi^{(n)}(\mathbf{k},\lambda;\cdot)\|_{\mathbb{C}^4 \otimes \mathcal{F}^{(n-1)}}^2 < \infty,$$

and hence $\sum_{n=1}^{\infty} n \|\Psi^{(n)}(\mathbf{k}, \lambda; \cdot)\|_{\mathbb{C}^4 \otimes \mathcal{F}^{(n-1)}}^2 < \infty$ for almost every (\mathbf{k}, λ) .

We set $\mathbf{g}(\mathbf{k}, \lambda) := \mathbf{g}(\mathbf{k}, \lambda; 0)$.

Proposition 6.3. Let $\nu > 0$. Then $a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p}) \in \text{Dom}(H_{\nu}(\mathbf{p}))$ and

(6.11)
$$a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p}) = \frac{q}{\sqrt{2}}(H_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1}\boldsymbol{\alpha} \cdot \mathbf{g}(\mathbf{k}, \lambda)\Phi_{\nu}(\mathbf{p})$$

for almost every $(\mathbf{k}, \lambda) \in \mathbb{R}^3 \times \{1, 2\}$.

Proof. For all $f \in \text{Dom}(\omega_{\nu})$ and $\Psi \in \mathcal{D}$, we have

$$\begin{split} \langle (H_{\nu}(\mathbf{p}) - E_{\nu}(\mathbf{p}))\Psi, a(f)\Phi_{\nu}(\mathbf{p}) \rangle \\ &= \bigg\langle \Psi, \big\{ -a(\omega_{\nu}f) + \boldsymbol{\alpha} \cdot a(\mathbf{k}f) + \frac{q}{\sqrt{2}}\boldsymbol{\alpha} \cdot \langle f, \mathbf{g} \rangle \big\} \Phi_{\nu}(\mathbf{p}) \bigg\rangle. \end{split}$$

Hence

$$\sum_{\lambda=1,2} \int_{\mathbb{R}^3} d\mathbf{k} f(\mathbf{k}, \lambda)^* \langle (H_{\nu}(\mathbf{p}) - E_{\nu}(\mathbf{p})) \Psi, a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p}) \rangle
= \sum_{\lambda=1,2} \int_{\mathbb{R}^3} d\mathbf{k} f(\mathbf{k}, \lambda)^* \langle \Psi, -\omega_{\nu}(\mathbf{k}) a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p})
+ \alpha \cdot \mathbf{k} a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p}) + \frac{q}{\sqrt{2}} \alpha \cdot \mathbf{g}(\mathbf{k}, \lambda) \Phi_{\nu}(\mathbf{p}) \rangle.$$

Since $\text{Dom}(\omega_{\nu})$ is dense in $L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$, we have

$$\begin{split} \langle (H_{\nu}(\mathbf{p}) - E_{\nu}(\mathbf{p}))\Psi, a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p}) \rangle \\ &= \left\langle \Psi, (-\omega_{\nu}(\mathbf{k})a_{\lambda}(\mathbf{k}) + \boldsymbol{\alpha} \cdot \mathbf{k} \, a_{\lambda}(\mathbf{k}) + \frac{q}{\sqrt{2}} \boldsymbol{\alpha} \cdot \mathbf{g}(\mathbf{k}, \lambda)) \Phi_{\nu}(\mathbf{p}) \right\rangle \end{split}$$

for almost every $(\mathbf{k}, \lambda) \in \mathbb{R}^3 \times \{1, 2\}$ and all $\Psi \in \mathcal{D}$. This means that $a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p}) \in D(H_{\nu}(\mathbf{p}))$ and

$$(H_{\nu}(\mathbf{p}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}) - \boldsymbol{\alpha} \cdot \mathbf{k})a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p}) = \frac{q}{\sqrt{2}}\boldsymbol{\alpha} \cdot \mathbf{g}(\mathbf{k}, \lambda)\Phi_{\nu}(\mathbf{p}).$$

Hence (6.11) follows.

Lemma 6.4. Suppose that $\hat{\rho}$ is spherically symmetric and $\hat{\rho} \in \text{Dom}(\omega^{-3/2})$. Assume that $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$. Then

(6.12)
$$\limsup_{\nu \to 0} \|N_f^{1/2} \Phi_{\nu}(\mathbf{p})\|^2 \le \int_{\mathbb{R}^3} d\mathbf{k} \, \frac{q^2}{(E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}|)^2} \, \frac{|\hat{\rho}(\mathbf{k})|^2}{|\mathbf{k}|} < \infty,$$

(6.13)
$$\limsup_{\nu \to 0} \|H_f^{1/2} \Phi_{\nu}(\mathbf{p})\|^2 \le \int_{\mathbb{R}^3} d\mathbf{k} \, \frac{q^2}{(E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}|)^2} |\hat{\rho}(\mathbf{k})|^2 < \infty.$$

Proof. By Proposition 6.3 and (6.10) with w = 1, we have

$$||N_f^{1/2} \Phi_{\nu}(\mathbf{p})||^2 \leq \sum_{\lambda=1}^2 \int_{\mathbb{R}^3} \frac{q^2}{2} \frac{||\boldsymbol{\alpha} \cdot \mathbf{g}(\mathbf{k}, \lambda) \Phi_{\nu}(\mathbf{p})||^2}{(E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| + \nu)^2} d\mathbf{k}$$
$$= \int_{\mathbb{R}^3} \frac{q^2}{(E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| + \nu)^2} \frac{|\hat{\rho}(\mathbf{k})|^2}{|\mathbf{k}|} d\mathbf{k}.$$

By Theorem C.10 and $\hat{\rho} \in \text{Dom}(\omega^{-3/2})$, the right hand side of (6.12) is finite. Hence, by Proposition C.9 and the Lebesgue convergence theorem, one has (6.12). The proof of (6.13) is similar. The only thing we have to do is set $w(\mathbf{k}) = \omega(\mathbf{k})$.

Proof of Theorem 4.1. By Proposition C.2, we have

$$0 \le E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}| \le 2|\mathbf{k}|.$$

Hence, by (4.1),

$$\frac{q^2}{4} \int_{\mathbb{R}^3} \frac{|\hat{\rho}(\mathbf{k})|^2}{|\mathbf{k}|^3} d\mathbf{k} \le \int_{\mathbb{R}^3} \frac{q^2}{(E(\mathbf{p} - \mathbf{k}) - E(\mathbf{p}) + |\mathbf{k}|)^2} \frac{|\hat{\rho}(\mathbf{k})|^2}{|\mathbf{k}|} d\mathbf{k} < 1,$$

which implies $\hat{\rho} \in \text{Dom}(\omega^{-3/2})$. Hence (6.12) and (6.13) hold.

Since $\Phi_{\nu}(\mathbf{p})$ is a unit vector, there exists a subsequence ν_{j} such that $\nu_{j} \to 0$ as $j \to \infty$ and $\Phi_{0}(\mathbf{p}) := \text{w-lim}_{j \to \infty} \Phi_{\nu_{j}}(\mathbf{p})$ exists. Then, by (6.12) and (6.13),

$$\lim_{j \to \infty} \|N_f^{1/2} \Phi_{\nu_j}\| < 1, \qquad \lim_{j \to \infty} \|H_f^{1/2} \Phi_{\nu_j}\| < \infty,$$

which implies that $\Phi_0(\mathbf{p}) \in \text{Dom}(N_f^{1/2}) \cap \text{Dom}(H_f^{1/2})$. Hence $\Phi_0(\mathbf{p}) \in Q(\overline{H(\mathbf{p})})$, where Q denotes the form domain. For any $\varphi \in \text{Dom}(H(\mathbf{p}))$, we have

$$\langle (H(\mathbf{p}) - E(\mathbf{p}))\varphi, \Phi_0(\mathbf{p}) \rangle = \lim_{j \to \infty} \langle (H(\mathbf{p}) - E(\mathbf{p}))\varphi, \Phi_{\nu_j}(\mathbf{p}) \rangle$$
$$= \lim_{j \to \infty} \langle \varphi, (E_{\nu_j}(\mathbf{p}) - E(\mathbf{p}) - \nu_j (H_f + N_f)) \Phi_{\nu_j}(\mathbf{p}) \rangle = 0.$$

Thus $\Phi_0(\mathbf{p}) \in \text{Dom}(\overline{H(\mathbf{p})})$ and $(\overline{H(\mathbf{p})} - E(\mathbf{p}))\Phi_0(\mathbf{p}) = 0$. Therefore, if $\Phi_0(\mathbf{p}) \neq 0$, then $\Phi_0(\mathbf{p})$ is a ground state of $\overline{H(\mathbf{p})}$. Since \mathbb{C}^4 is a finite-dimensional space, the vacuum component $\Phi_{\nu_i}(\mathbf{p})^{(0)}$ strongly converges to $\Phi_0(\mathbf{p})^{(0)}$. Hence

(6.14)
$$\|\Phi_0(\mathbf{p})\|^2 \ge \|\Phi_0(\mathbf{p})^{(0)}\|^2 = \lim_{j \to \infty} \|\Phi_{\nu_j}(\mathbf{p})^{(0)}\|^2 = \lim_{j \to \infty} \langle \Phi_{\nu_j}(\mathbf{p}), P_{\Omega}\Phi_{\nu_j}(\mathbf{p}) \rangle,$$

where P_{Ω} is the orthogonal projection on the vacuum $(1, 0, 0, ...) \in \mathcal{F}_{rad}$. Thus, using (6.14) and $N_f \geq 1 - P_{\Omega}$, we have

$$\|\Phi_0(\mathbf{p})\|^2 \ge 1 - \lim_{j \to \infty} \|N_f^{1/2} \Phi_{\nu_j}(\mathbf{p})\|^2 > 0.$$

This means that $\Phi_0(\mathbf{p}) \neq 0$ and $\Phi_0(\mathbf{p})$ is a ground state of $\overline{H(\mathbf{p})}$.

Proof of Theorem 4.2. Theorem 4.2 follows immediately from Theorems 4.1 and C.10.

Next, we prepare some lemmas for the proof of Theorem 4.4. For a Hilbert space \mathcal{K} , we denote by $\mathsf{B}(\mathcal{K})$ the set of all bounded operators on \mathcal{K} . The next lemma follows from the second resolvent equation.

Lemma 6.5. Let $\nu > 0$. For each $\mathbf{j} \in \mathbb{R}^3$ with $|\mathbf{j}| = 1$, the operator valued function $\mathbb{R}^3 \setminus \{\mathbf{0}\} : \mathbf{k} \to (H_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \in \mathsf{B}(\mathbb{C}^4 \otimes \mathcal{F}_{\mathrm{rad}})$ is differentiable in the sense of operator norm, and

$$\begin{split} &\partial_{\mathbf{j}}(H_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}\\ &=(H_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}\bigg(\boldsymbol{\alpha}\cdot\mathbf{j}-(1+\nu)\frac{\mathbf{k}\cdot\mathbf{j}}{|\mathbf{k}|}\bigg)(H_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}, \end{split}$$

where $\partial_{\mathbf{j}}$ means the \mathbf{j} -direction derivative.

We fix the following polarization vectors in the rest of this section:

(6.15)
$$\mathbf{e}^{(1)}(\mathbf{k}) = \frac{(k_2, -k_1, 0)}{\sqrt{k_1^2 + k_2^2}}, \quad \mathbf{e}^{(2)}(\mathbf{k}) := \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}^{(1)}(\mathbf{k}).$$

Now, recall the definition of the set S (defined in condition (Λ)). We set

$$\mathsf{X} := S \setminus \{\mathbf{k} \in \mathbb{R}^3 \mid k_1 = k_2 = 0\}, \quad \mathsf{X}_R := S_R \cap \mathsf{X}.$$

By Lemma 6.5 and (6.15), we obtain the following result:

Lemma 6.6. Under the assumptions of Theorem 4.4, $a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p})$ is strongly continuously differentiable in X and

$$\partial_{j} a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p}) = \frac{q}{\sqrt{2}} (H_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \left(\alpha_{j} - (1 + \nu) \frac{k_{j}}{|\mathbf{k}|} \right)$$

$$\times (H_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \boldsymbol{\alpha} \cdot \mathbf{g}(\mathbf{k}, \lambda) \Phi_{\nu}(\mathbf{p})$$

$$+ \frac{q}{\sqrt{2}} (H_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \boldsymbol{\alpha} \cdot (\partial_{j} \mathbf{g}(\mathbf{k}, \lambda)) \Phi_{\nu}(\mathbf{p}),$$

where ∂_j denotes the strong derivative in k_j (j = 1, 2, 3).

We set

$$\Psi_j(\mathbf{k},\lambda) = (\Psi_j^{(n)}(\mathbf{k},\lambda;\cdot))_{n=0}^{\infty} := \partial_j a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p}).$$

Lemma 6.7. Under the assumptions of Theorem 4.4,

$$\partial_j \Phi_{\nu}^{(n)}(\mathbf{p})(\mathbf{k}, \lambda; X; k_2, \dots, k_n) = \frac{1}{\sqrt{n}} \Psi_j^{(n-1)}(\mathbf{k}, \lambda; X; k_2, \dots, k_n), \quad k_{\ell} = (\mathbf{k}_{\ell}, \lambda_{\ell}),$$

for all $X \in \{1, 2, 3, 4\}$, $\mathbf{k}, \mathbf{k}_{\ell} \in X$, $n \in \mathbb{N}$, $\lambda, \lambda_{\ell} = 1, 2$ and j = 1, 2, 3, where ∂_{j} is the distributional derivative with respect to k_{j} .

Note that the ∂_j on the left hand side is a distributional derivative and the one in Ψ_i is a strong derivative.

Proof. In this proof, for simplicity, we do not indicate $X, \lambda, \lambda_{\ell}$ and \mathbf{p} . The operator δ_h is defined by $\delta_h f(\mathbf{k}) := f(\mathbf{k} + h\mathbf{j}) - f(\mathbf{k})$ for all functions $f(\mathbf{k})$. Let $\psi(\mathbf{k}, \mathbf{k}_2, \dots, \mathbf{k}_n) \in C_0^{\infty}(X^{n+1})$ be arbitrary. Clearly, we have $(\partial_j \psi)(\mathbf{k}, K) = \lim_{h \to 0} h^{-1}(\psi(\mathbf{k} + h\mathbf{j}, K) - \psi(\mathbf{k}, K))$ uniformly, where $K = (\mathbf{k}_2, \dots, \mathbf{k}_n)$ and \mathbf{j} is the unit vector of the j-th axis. By the definition of the distributional derivative, we have

$$\begin{split} \int_{\mathbb{R}^{3n}} d\mathbf{k} \, dK \, \psi(\mathbf{k}, K) \partial_j \Phi_{\nu}^{(n)}(\mathbf{k}, K) &= -\int_{\mathbb{R}^{3n}} d\mathbf{k} \, dK \, (\partial_j \psi)(\mathbf{k}, K) \Phi_{\nu}^{(n)}(\mathbf{k}, K) \\ &= -\lim_{h \to 0} \int_{\mathbb{R}^{3n}} d\mathbf{k} \, dK \, \frac{1}{-h} (\delta_{-h} \psi)(\mathbf{k}, K) \Phi_{\nu}^{(n)}(\mathbf{k}, K) \\ &= \lim_{h \to 0} \int_{\mathbb{R}^{3n}} d\mathbf{k} \, dK \, \psi(\mathbf{k}, K) \frac{1}{h} (\delta_h \Phi_{\nu}^{(n)})(\mathbf{k}, K). \end{split}$$

By Schwarz' inequality, we have

$$\left| \int_{\mathbb{R}^{3}} d\mathbf{k} \left[\int_{\mathbb{R}^{3(n-1)}} dK \, \psi(\mathbf{k}, K) \left\{ \frac{1}{h} [\Phi_{\nu}^{(n)}(\mathbf{k} + h\mathbf{j}, K) - \Phi_{\nu}^{(n)}(\mathbf{k}, K)] - \frac{1}{\sqrt{n}} \Psi^{(n-1)}(\mathbf{k}, K) \right\} \right] \right| \\
\leq \int_{\mathbb{R}^{3}} d\mathbf{k} \, \|\psi(\mathbf{k}, \cdot)\|_{L^{2}(\mathbb{R}^{3(n-1)})} \left\| \frac{\delta_{h}}{h} \Phi_{\nu}^{(n)}(\mathbf{k}, \cdot) - \frac{1}{\sqrt{n}} \Psi^{(n-1)}(\mathbf{k}, \cdot) \right\|_{L^{2}(\mathbb{R}^{3(n-1)})}.$$

Note that, for all $\mathbf{k} \in \mathsf{X}$, $h^{-1}\delta_h\Phi_{\nu}^{(n)}(\mathbf{k},\cdot)$ strongly converges to $\frac{1}{\sqrt{n}}\Psi^{(n-1)}(\mathbf{k},\cdot)$ in $L^2(\mathsf{X}^{3(n-1)})$ by Lemma 6.6. Moreover, by Lemma 6.6 and the assumption that $\hat{\rho}$ is continuously differentiable, the function $\mathbf{k}\mapsto\Psi^{(n-1)}(\mathbf{k},\cdot)$ is strongly continuous in X . Let D be the closure of $\{\mathbf{k}\in\mathbb{R}^3\mid \|\psi(\mathbf{k},\cdot)\|_{L^2(\mathbb{R}^{3(n-1)})}\neq 0\}$. Note that $D\subset\mathsf{X}$ is a compact set and $d:=\mathrm{dist}(D,X^c)>0$.

For every $\mathbf{k} \in D$ and h with |h| < d, we have

$$\frac{\delta_h}{h} \Phi_{\nu}^{(n)}(\mathbf{k}, \cdot) = s - \int_0^1 \frac{1}{\sqrt{n}} \Psi^{(n-1)}(\mathbf{k} + th\mathbf{j}, \cdot) dt,$$

where s- \int means the strong integral in $L^2(\mathsf{X}^{3(n-1)})$. Since $\|\Psi^{(n-1)}(\mathbf{k},\cdot)\|_{L^2(\mathbb{R}^{3(n-1)})}$ is continuous in $\mathbf{k} \in \mathsf{X}$, it is bounded on the compact set D. For any $\mathbf{k} \in D$ and

|h| < d, we have

$$\begin{split} & \left\| \frac{\delta_h}{|h|} \Phi_{\nu}^{(n)}(\mathbf{k}, \cdot) - \frac{1}{\sqrt{n}} \Psi^{(n-1)}(\mathbf{k}, \cdot) \right\|_{L^2(\mathbb{R}^{3(n-1)})} \\ & \leq \sup_{|t| \leq 1} \frac{1}{\sqrt{n}} \| \Psi^{(n-1)}(\mathbf{k} + th\mathbf{j}, \cdot) \|_{L^2(\mathbb{R}^{3(n-1)})} + \frac{1}{\sqrt{n}} \| \Psi^{(n-1)}(\mathbf{k}, \cdot) \|_{L^2(\mathbb{R}^{3(n-1)})} \leq \mathrm{const}, \end{split}$$

where "const" means a constant independent of **k** and h. Applying the Lebesgue dominated convergence theorem, we can see that the right hand side of (6.16) converges to zero as $|h| \to 0$.

By Lemmas 6.5–6.6 and direct calculations, we obtain the following inequality.

Lemma 6.8. Under the assumptions of Theorem 4.4,

$$\begin{split} \|\partial_{j}a_{\lambda}(\mathbf{k})\Phi_{\nu}(\mathbf{p})\| &\leq \frac{|q|}{\sqrt{2}}(2+\nu)(E_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-2}\frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}} \\ &+ \frac{|q|}{\sqrt{2}}(E_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}\frac{|\partial_{j}\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}} \\ &+ \frac{|q|}{\sqrt{2}}(E_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}\frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{3/2}} \\ &+ \frac{|q|}{\sqrt{2}}(E_{\nu}(\mathbf{p}-\mathbf{k})-E_{\nu}(\mathbf{p})+\omega_{\nu}(\mathbf{k}))^{-1}\frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}}|\partial_{j}\mathbf{e}^{(\lambda)}(\mathbf{k})| \end{split}$$

for all $\mathbf{k} \in X$, $\lambda = 1, 2, j = 1, 2, 3$.

Our polarization vectors (6.15) satisfy

(6.17)
$$|\partial_j \mathbf{e}^{(\lambda)}(\mathbf{k})| \le \frac{2}{\sqrt{k_1^2 + k_2^2}} \quad \text{for } \mathbf{k} \in \mathbb{R}^3 \setminus \{\mathbf{k}' \in \mathbb{R}^3 \mid k_1' = k_2' = 0\}.$$

We set

$$f_{\nu}^{(1)}(\mathbf{k}) := (E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-2} \frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}},$$

$$f_{\nu}^{(2)}(\mathbf{k}) := (E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \frac{|\partial_{j}\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}},$$

$$f_{\nu}^{(3)}(\mathbf{k}) := (E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{3/2}},$$

$$f_{\nu}^{(4)}(\mathbf{k}) := (E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + \omega_{\nu}(\mathbf{k}))^{-1} \frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}} |\partial_{j}\mathbf{e}^{(\lambda)}(\mathbf{k})|.$$

Lemma 6.9. Under the assumptions of Theorem 4.4,

(6.18)
$$\sup_{0 < \nu \le 1} \|f_{\nu}^{(j)}\|_{L^{p}(S_{R})} < \infty, \quad j = 1, 2, 3, 4, \, p \in [1, 2).$$

Proof. First we consider the case $\mathbf{p} \neq \mathbf{0}$. Let $b_{\nu}(\mathbf{p})$ be the constant defined in Theorem C.10. Since $b_{\nu}(\mathbf{p})$ is continuous in ν for fixed \mathbf{p} , Theorem C.10 guarantees $\sup_{0 < \nu < 1} b_{\nu}(\mathbf{p}) = \max_{0 \le \nu \le 1} b_{\nu}(\mathbf{p}) < 1$. By Theorem C.10, we have

$$(E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}|)^{-1} \le \frac{1}{1 - b_{\nu}(\mathbf{p})} \max \left\{ \frac{1}{|\mathbf{k}|}, \frac{1}{|\mathbf{p}|} \right\} \le C \max \left\{ \frac{1}{|\mathbf{k}|}, \frac{1}{|\mathbf{p}|} \right\},$$

where

$$C := \sup_{0 < \nu < 1} \frac{1}{1 - b_{\nu}(\mathbf{p})}$$

is a finite constant. Hence

$$f_{\nu}^{(1)}(\mathbf{k}) \le C^2 \left\{ \frac{1}{|\mathbf{k}|^2} + \frac{1}{|\mathbf{p}|^2} \right\} \frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}}.$$

Since S_R is a bounded region, by the assumption $|\mathbf{k}|^{-5/2}|\hat{\rho}(\mathbf{k})| \in L^p(S_R)$, we obtain

$$\sup_{0<\nu<1} \|f_{\nu}^{(1)}\|_{L^{p}(S_{R})} < \infty.$$

Similarly,

$$\sup_{0<\nu\leq 1} \|f_{\nu}^{(j)}\|_{L^2(S_R)} < \infty, \quad j=2,3.$$

By (6.17), we have

$$f_{\nu}^{(4)}(\mathbf{k}) \le C^2 \left\{ \frac{1}{|\mathbf{k}|} + \frac{1}{|\mathbf{p}|} \right\} \frac{1}{\sqrt{k_1^2 + k_2^2}} \frac{|\hat{\rho}(\mathbf{k})|}{|\mathbf{k}|^{1/2}}$$

By using polar coordinates, we have

$$\int_{S_R} f_{\nu}^{(4)}(\mathbf{k}) d\mathbf{k} \leq 2\pi C \int_{[0,\pi)} \sin\theta d\theta \left[\frac{1}{\sin\theta} \right]^p \int_{[0,R)} |\mathbf{k}|^{2-p} \left(\frac{|\mathbf{k}| + |\mathbf{p}|}{|\mathbf{k}| \cdot |\mathbf{p}|} \right)^p \frac{|\hat{\rho}(\mathbf{k})|^p}{|\mathbf{k}|} d|\mathbf{k}|$$

$$< \infty.$$

Next we consider the case $\mathbf{p} = 0$. By (C.4), we have

$$(E_{\nu}(-\mathbf{k}) - E_{\nu}(\mathbf{0}) + \omega_{\nu}(\mathbf{k}))^{-1} \le \begin{cases} \frac{P}{a_{\nu}(P)|\mathbf{k}|} & \text{if } |\mathbf{k}| \le P, \\ a_{\nu}(P)^{-1} & \text{if } |\mathbf{k}| > P, \end{cases}$$

for any P > 0. By similar arguments, one can prove (6.18).

Let $W^{1,p}(\mathcal{X})$ be the Sobolev space on the configuration space \mathcal{X} , i.e., the set of all L^p -functions with their first derivatives also in L^p .

Lemma 6.10. Under the assumptions of Theorem 4.4, the n-th component of the massive ground state satisfies $\Phi_{\nu}^{(n)} \in \bigoplus^4 W^{1,p}((\mathsf{X}_R \times \{1,2\})^n)$ for all $p \in [1,2)$ and all R > 0, and

$$\sup_{0<\nu<1} \|\Phi_{\nu}^{(n)}(\mathbf{p})\|_{\bigoplus^{4} W^{1,p}((\mathsf{X}_R\times\{1,2\})^n)} < \infty.$$

Proof. By Lemma 6.7, we have

$$(\nabla_{\mathbf{k}} a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p}))^{(n-1)} (X; \mathbf{k}_{1}, \lambda_{1}; \dots; \mathbf{k}_{n-1}, \lambda_{n-1})$$

$$= \sqrt{n} \nabla_{\mathbf{k}} \Phi_{\nu}^{(n)}(\mathbf{p}; X; \mathbf{k}, \lambda; \mathbf{k}_{1}, \lambda_{1}; \dots; \mathbf{k}_{n-1}, \lambda_{n-1}).$$

Using Hölder's inequality and making a change of variables, one has, for all p < 2,

$$(6.19) \sum_{X=1}^{4} \sum_{\lambda_{1},\dots,\lambda_{n} \in \{1,2\}} \int_{(\mathsf{X}_{R})^{n}} d\mathbf{k}_{1} \cdots d\mathbf{k}_{n} \sum_{i=1}^{n} |\nabla_{\mathbf{k}_{i}} \Phi_{\nu}^{(n)}(\mathbf{p}; X; \mathbf{k}_{1}, \lambda_{1}; \dots; \mathbf{k}_{n}, \lambda_{n})|^{p}$$

$$\leq C \int_{\mathsf{X}_{R}} d\mathbf{k} \|\nabla_{\mathbf{k}} a_{\lambda}(\mathbf{k}) \Phi_{\nu}(\mathbf{p})\|^{p},$$

where C is a constant independent of ν . By Lemmas 6.8 and 6.9, the right hand side of (6.19) is finite uniformly in $\nu > 0$.

Proof of Theorem 4.4. As shown in the proof of Theorem 4.1, there exists a sequence $\{\nu_j\}_{j=1}^{\infty}$ such that the limit $\Phi_0(\mathbf{p}) := \text{w-lim}_{j\to\infty} \Phi_{\nu_j}(\mathbf{p})$ exists, and $\Phi_0(\mathbf{p}) \in \text{Dom}(H_f^{1/2}) \cap \text{Dom}(N_f^{1/2})$. Thus, $\Phi_0 \in Q(H(\mathbf{p}))$. If $\Phi_0(\mathbf{p}) \neq 0$, then $\Phi_0(\mathbf{p})$ is a ground state of $H(\mathbf{p})$. In the following, we show that indeed $\Phi_0(\mathbf{p}) \neq 0$.

Any vector $\Psi \in \bigoplus^4 \mathcal{F}^n = \mathbb{C}^4 \otimes \mathcal{F}^n$ is a function of the particle helicity $X \in \{1, 2, 3, 4\}$, the *n*-photon wave number $(\mathbf{k}_1, \dots, \mathbf{k}_n) \in \mathbb{R}^{3n}$, and the photon polarization $\lambda_1, \dots, \lambda_n \in \{1, 2\}$. For simplicity, we set

$$\begin{split} &\Phi_j^{(n)}(\mathbf{k}_1,\ldots,\mathbf{k}_n) := \Phi_{\nu_j}(\mathbf{p})^{(n)}(X;\mathbf{k}_1,\lambda_1;\ldots;\mathbf{k}_n,\lambda_n), \\ &\Phi_0^{(n)}(\mathbf{k}_1,\ldots,\mathbf{k}_n) := \Phi_0(\mathbf{p})^{(n)}(X;\mathbf{k}_1,\lambda_1;\ldots;\mathbf{k}_n,\lambda_n) \end{split}$$

for $X \in \{1, 2, 3, 4\}$ and $\lambda_1, \dots, \lambda_n \in \{1, 2\}$. Note that $\Phi_j^{(n)}, \Phi_0^{(n)} \in L^2(\mathbb{R}^{3n})$. We show that s- $\lim_{j\to\infty} \Phi_j^{(n)} = \Phi_0^{(n)}$ for all $n \in \mathbb{N}, X \in \{1, 2, 3, 4\}$ and $\lambda_1, \dots, \lambda_n \in \{1, 2\}$.

By Lemma 6.10 and the Rellich-Kondrashov theorem,

(6.20)
$$\lim_{i \to \infty} \|\Phi_j^{(n)} - \Phi_0^{(n)}\|_{L^2(X_R^n)} = 0$$

for all R > 0 (see [GLL, p. 578] for details). We set $\Phi_j := (\Phi_j^{(n)})_{n=0}^{\infty}$ and $\Phi_0 := (\Phi_0^{(n)})_{n=0}^{\infty} \in \bigoplus^4 \mathcal{F}_{rad}$. Let χ_R be the characteristic function of the ball $\{\mathbf{k} \in \mathbb{R}^3 \mid \Phi_0^{(n)}\}_{n=0}^{\infty}$

 $|\mathbf{k}| < R$. We denote the orthogonal projection onto $\bigoplus_{j=0}^n \mathbb{C}^4 \otimes \mathcal{F}^j$ by P_n . Then we have

$$\|\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2 = \|P_n\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2 + \|(1 - P_n)\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2$$

$$\leq \|P_n\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2 + \frac{1}{n}\|N_f^{1/2}\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2.$$

Since each component $(\Gamma(\chi_R)\Phi_j)^{(n)}$ converges to $(\Gamma(\chi_R)\Phi_0)^{(n)}$ strongly as $j\to\infty$, we have

$$\limsup_{j \to \infty} \|\Gamma(\chi_R)(\Phi_j - \Phi_0)\|^2 \le \frac{1}{n} \limsup_{j \to \infty} \|N_f^{1/2}(\Phi_j - \Phi_0)\|^2$$

for all $n \in \mathbb{N}$. By Lemma 6.4, $\limsup_{j \to \infty} \|N_f^{1/2}(\Phi_j - \Phi_0)\|^2 < \infty$. Thus we obtain

(6.21)
$$\operatorname{s-lim}_{j \to \infty} \Gamma(\chi_R) \Phi_j = \Gamma(\chi_R) \Phi_0.$$

Therefore for all R > 0 we have

$$\begin{split} \|\Phi_{j} - \Phi_{0}\| &= \|\Gamma(\chi_{R})(\Phi_{j} - \Phi_{0})\| + \|(1 - P_{0})(\Gamma(\chi_{R}) - 1)(\Phi_{j} - \Phi_{0})\|^{2} \\ &\leq \|\Gamma(\chi_{R})(\Phi_{j} - \Phi_{0})\| + \|(1 - P_{0})(1 - \Gamma(\chi_{R}))H_{f}^{-1/2}\| \cdot \|H_{f}^{1/2}(\Phi_{j} - \Phi_{0})\| \\ &\leq \|\Gamma(\chi_{R})(\Phi_{j} - \Phi_{0})\| + \frac{C}{R^{1/2}}, \end{split}$$

where C is a constant independent of R > 0. By (6.21), we obtain

$$\operatorname{s-lim}_{j\to\infty}\Phi_j=\Phi_0,$$

which implies that Φ_0 is a normalized ground state of $\overline{H(\mathbf{p})}$.

§7. Proof of Theorem 5.2

Throughout this section we assume that the assumptions of Theorem 5.2 hold. By Appendices A and B, it suffices to prove Theorem 5.2 in the case $\mathbf{e} = \bar{\mathbf{e}}$. Here $\bar{\mathbf{e}}$ is the polarization vector defined in (B.1). Note that $\bar{\mathbf{e}}$ depends on \mathbf{j} . By assumption, there exists a non-negative constant t such that $\mathbf{p} = t\mathbf{j}$. We choose a matrix $T \in SO(3)$ such that $T^{-1}\mathbf{p} = (0,0,|\mathbf{p}|)$ and $T^{-1}\mathbf{j} = (0,0,1)$. Let U be the unitary operator defined in the proof of Proposition C.4. By (C.1), we obtain

$$U\overline{H(\mathbf{p})}U^* = \overline{(|\mathbf{p}|\alpha_3 + m\beta + H_f - \boldsymbol{\alpha} \cdot d\Gamma(\mathbf{k}) - q\boldsymbol{\alpha} \cdot \Phi_{\mathrm{S}}(\vec{\lambda}))},$$

where

$$\vec{\lambda} = (\lambda_1, \lambda_2, \lambda_3) = \frac{\hat{\rho}(T\mathbf{k})}{|\mathbf{k}|^{1/2}} (T^{-1}\bar{\mathbf{e}}^{(1)}(T\mathbf{k}), T^{-1}\bar{\mathbf{e}}^{(2)}(T\mathbf{k})) \in (L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1, 2\}))^3.$$

Since $T \in SO(3)$, we have

$$T^{-1}\bar{\mathbf{e}}^{(1)}(T\mathbf{k}) = \frac{T^{-1}[(T\mathbf{k}) \wedge \mathbf{j}]}{|(T\mathbf{k}) \wedge \mathbf{j}|} = \frac{\mathbf{k} \wedge (0, 0, 1)}{|\mathbf{k} \wedge (0, 0, 1)|},$$
$$T^{-1}\bar{\mathbf{e}}^{(2)}(T\mathbf{k}) = \frac{\mathbf{k}}{|\mathbf{k}|} \wedge (T^{-1}\bar{\mathbf{e}}^{(1)}(T\mathbf{k})).$$

It is easy to see that $\hat{\rho}(TR'\mathbf{k}) = \hat{\rho}(T\mathbf{k})$, $\mathbf{k} \in \mathbb{R}^3$, for all $R' \in O(3)$ such that R'(0,0,1) = (0,0,1). Since $\mathbf{S} = (i/4)\boldsymbol{\alpha} \wedge \boldsymbol{\alpha}$, we have

$$U(\mathbf{j} \cdot \mathbf{S})U^* = \frac{i}{4}\mathbf{j} \cdot [(T\alpha) \cdot (T\alpha)] = \frac{i}{4}\mathbf{j} \cdot [T(\alpha \wedge \alpha)] = \frac{i}{4}(\alpha \wedge \alpha)_3 = S_3.$$

Moreover, one can show that $U(\mathbf{j} \cdot d\Gamma(\vec{\ell}))U^* = d\Gamma(\ell_3)$. Therefore,

$$UJ_{\mathbf{j}}(\bar{\mathbf{e}})U^* = S_3 + d\Gamma(\ell_3),$$

and hence it is sufficient to prove Theorem 5.2 in the case

(7.1)
$$\mathbf{p} = (0, 0, |\mathbf{p}|), \quad \mathbf{j} = (0, 0, 1).$$

Proof of Theorem 5.2. We assume (7.1) holds. We put

$$\check{\mathbf{e}}^{(1)}(\mathbf{k}) := \frac{(k_2, -k_1, 0)}{\sqrt{k_1^2 + k_2^2}}, \quad \check{\mathbf{e}}^{(2)}(\mathbf{k}) := \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \check{\mathbf{e}}^{(1)}(\mathbf{k}).$$

For a real parameter $\theta \in \mathbb{R}$, we set

$$W := \exp[i\theta J_{\mathbf{j}}(\check{\mathbf{e}})], \quad \Theta := \begin{bmatrix} \cos \theta & -\sin \theta & 0 \\ \sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix}.$$

Then we obtain

(7.2)
$$W\alpha W^* = \Theta\alpha, \quad W\beta W^* = \beta,$$

(7.3)
$$Wd\Gamma(\mathbf{k})W^* = \Theta d\Gamma(\mathbf{k}), \quad WH_f(m)W^* = H_f(m),$$

$$(7.4) W\mathbf{A}W^* = \Theta\mathbf{A}.$$

Here, to show (7.4), we used the specific form of ě:

$$\check{\mathbf{e}}^{(\lambda)}(\Theta \mathbf{k}) = \Theta \check{\mathbf{e}}^{(\lambda)}(\mathbf{k}), \quad \lambda = 1, 2.$$

Since $\theta \in \mathbb{R}$ is arbitrary, (7.2)–(7.4) imply that $\overline{H(\mathbf{p})}$ strongly commutes with $J_{\mathbf{j}}(\check{\mathbf{e}})$. Thus, $\overline{H(\mathbf{p})}$ is reduced by the projection onto the eigenspace of $J_{\mathbf{j}}(\check{\mathbf{e}})$. In other words, $\overline{H(\mathbf{p})}$ decomposes as

$$\overline{H(\mathbf{p})} \cong \bigoplus_{z \in \mathbb{Z}_{1/2}} H(\mathbf{p}:z),$$

in the sense of (5.1). We furthermore define unitary operators η , τ and Υ by

$$(\eta f)(\mathbf{k}, \lambda) := \begin{cases} -f(k_1, -k_2, k_3, 1) & \text{if } \lambda = 1, \\ f(k_1, -k_2, k_3, 2) & \text{if } \lambda = 2, f \in L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1, 2\}), \end{cases}$$
$$\tau := \alpha_1 \alpha_2 \beta, \quad \Upsilon := \tau \cdot \Gamma(\eta).$$

It is easy to see that

$$\begin{split} \eta \ell_3 \eta^* &= -\ell_3, \quad \tau S_3 \tau^* = -S_3, \\ \eta k_1 \eta^* &= k_1, \quad \eta k_2 \eta^* = -k_2, \quad \eta k_3 \eta^* = k_3, \\ \tau \alpha_1 \tau^* &= \alpha_1, \quad \tau \alpha_2 \tau^* = -\alpha_2, \quad \tau \alpha_3 \tau^* = \alpha_3, \quad \tau \beta \tau^* = \beta, \\ \eta \check{\mathbf{e}}^{(1)}(\mathbf{k}) \eta^{-1} &= \frac{(k_2, -(-k_1), 0)}{\sqrt{k_1^2 + k_2^2}}, \quad \eta \check{\mathbf{e}}^{(2)}(\mathbf{k}) \eta^{-1} = \frac{(k_1 k_3, -k_2 k_3, -k_1^2 - k_2^2)}{|\mathbf{k}| \sqrt{k_1^2 + k_2^2}}. \end{split}$$

Hence

$$\Upsilon \overline{H(\mathbf{p})} \Upsilon^* = \overline{H(\mathbf{p})}, \quad \Upsilon J_{\mathbf{j}} \Upsilon^* = -J_{\mathbf{j}}.$$

Let E(z), $z \in \mathbb{Z}_{1/2}$, be the orthogonal projection onto $\ker(J_{\mathbf{j}} - z)$. Note that $\operatorname{Ran}(E(z)) = \mathcal{F}(z)$. Moreover $E(-z) \Upsilon E(z)$ is a unitary operator from $\operatorname{Ran}(E(z))$ to $\operatorname{Ran}(E(-z))$ and

$$E(-z)\Upsilon E(z)H(\mathbf{p}:z)E(z)\Upsilon^*E(-z) = E(-z)\Upsilon E(z)\Upsilon^*\overline{H(\mathbf{p})}\Upsilon E(z)\Upsilon^*E(-z)$$
$$= H(\mathbf{p}:-z).$$

Therefore $H(\mathbf{p}:z)$ is unitarily equivalent to $H(\mathbf{p}:-z)$ for all $z \in \mathbb{Z}_{1/2}$.

Appendix A. Remarks on polarization vectors

In this appendix, we show that quantum electrodynamics is independent of the choice of polarization vectors, i.e., the Hamiltonians defined by different polarization vectors are unitarily equivalent. We show the equivalence only for the Hamiltonians H and $H(\mathbf{p})$, but one can apply our proof to the Pauli–Fierz model and various QED models. The proof here is independent of the choice of $\hat{\rho}$ and ω .

We assume that the polarization vectors $\mathbf{e}^{(1)}(\mathbf{k})$, $\mathbf{e}^{(2)}(\mathbf{k})$ and \mathbf{k} form a right-handed system:

$$\mathbf{k} \cdot \mathbf{e}^{(1)}(\mathbf{k}) = 0, \quad \|\mathbf{e}^{(1)}(\mathbf{k})\|_{\mathbb{R}^3} = 1, \quad \mathbf{e}^{(2)}(\mathbf{k}) = \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}^{(1)}(\mathbf{k}), \quad \mathbf{k} \in \mathbb{R}^3.$$

Next, we take any polarization vectors $\mathbf{e}^{\prime(1)},\,\mathbf{e}^{\prime(2)}\colon$

$$\mathbf{k} \cdot \mathbf{e}^{\prime(\lambda)}(\mathbf{k}) = 0, \quad \mathbf{e}^{\prime(\lambda)}(\mathbf{k}) \cdot \mathbf{e}^{\prime(\mu)}(\mathbf{k}) = \delta_{\lambda,\mu}, \quad \mathbf{k} \in \mathbb{R}^3, \, \lambda, \mu \in \{1, 2\}.$$

Let H' and $H'(\mathbf{p})$ be the Hamiltonians H and $H(\mathbf{p})$ with $\mathbf{e}^{(\lambda)}$ replaced by $\mathbf{e}'^{(\lambda)}$, $\lambda = 1, 2$, respectively.

Theorem A.1. Assume that H is essentially self-adjoint. Then H' is also essentially self-adjoint and \bar{H} is unitarily equivalent to $\bar{H'}$ by means of a unitary operator $U(\mathbf{e} \leftarrow \mathbf{e'})$:

$$U(\mathbf{e} \leftarrow \mathbf{e}')\bar{H}'U(\mathbf{e} \leftarrow \mathbf{e}')^* = \bar{H}.$$

Theorem A.2. Assume that $H(\mathbf{p})$ is essentially self-adjoint. Then $H'(\mathbf{p})$ is also essentially self-adjoint and $\overline{H(\mathbf{p})}$ is unitarily equivalent to $\overline{H'(\mathbf{p})}$:

$$U(\mathbf{e} \leftarrow \mathbf{e}')\overline{H'(\mathbf{p})}U(\mathbf{e} \leftarrow \mathbf{e}')^* = \overline{H(\mathbf{p})}.$$

Remark A.3. The unitary operators $U(\mathbf{e} \leftarrow \mathbf{e}')$ defined below satisfy the chain rule:

$$U(\mathbf{e} \leftarrow \mathbf{e}') = U(\mathbf{e} \leftarrow \mathbf{e}'')U(\mathbf{e}'' \leftarrow \mathbf{e}'),$$

$$U(\mathbf{e} \leftarrow \mathbf{e}')^* = U(\mathbf{e}' \leftarrow \mathbf{e}).$$

Proofs of Theorems A.1 and A.2. By the definition of polarization vectors, for each $\mathbf{k} \in \mathbb{R}^3$ we have either $\mathbf{e}'^{(2)}(\mathbf{k}) = \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}'^{(1)}(\mathbf{k})$ or $\mathbf{e}'^{(2)}(\mathbf{k}) = -\frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}'^{(1)}(\mathbf{k})$. Let $O \subset \mathbb{R}^3$ be the set of all \mathbf{k} such that $\mathbf{e}'^{(2)}(\mathbf{k}) = -\frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}'^{(1)}(\mathbf{k})$. We define

$$\mathbf{e}''^{(1)}(\mathbf{k}) := \mathbf{e}'^{(1)}(\mathbf{k}), \quad \mathbf{e}''^{(2)}(\mathbf{k}) := \begin{cases} \mathbf{e}'^{(2)}(\mathbf{k}), & \mathbf{k} \in \mathbb{R}^3 \setminus O, \\ -\mathbf{e}'^{(2)}(\mathbf{k}), & \mathbf{k} \in O. \end{cases}$$

We define an operator H'' just as H with $\mathbf{e}^{(\lambda)}$ replaced by $\mathbf{e}''^{(\lambda)}$, $\lambda = 1, 2$. Let

$$\mathbf{g}'(\mathbf{k},\lambda;\mathbf{x}) := \frac{\hat{\rho}(\mathbf{k})}{|\mathbf{k}|^{1/2}} \mathbf{e}'^{(\lambda)}(\mathbf{k}) e^{-i\mathbf{k}\cdot\mathbf{x}}, \quad \mathbf{g}''(\mathbf{k},\lambda;\mathbf{x}) := \frac{\hat{\rho}(\mathbf{k})}{|\mathbf{k}|^{1/2}} \mathbf{e}''^{(\lambda)}(\mathbf{k}) e^{-i\mathbf{k}\cdot\mathbf{x}},$$

and we set

$$\mathbf{A}^{\sharp}(\hat{\mathbf{x}}) := \frac{1}{\sqrt{2}} \int_{\mathbb{R}^3}^{\oplus} \overline{[a(\mathbf{g}^{\sharp}(\cdot, \mathbf{x})) + a(\mathbf{g}^{\sharp}(\cdot, \mathbf{x}))^*]} \, d\mathbf{x},$$

where \sharp stands for ' or ". Since $(\mathbf{e}''^{(1)}(\mathbf{k}), \mathbf{e}''^{(2)}(\mathbf{k}), \mathbf{k})$ are right-handed vectors, i.e., $\mathbf{k} \cdot \mathbf{e}''^{(1)}(\mathbf{k}) = 0$, $\mathbf{e}''^{(2)}(\mathbf{k}) = \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \mathbf{e}''^{(1)}(\mathbf{k})$, there exists $\theta(\mathbf{k}) \in [0, 2\pi)$ such that

$$\begin{bmatrix} \mathbf{e}^{(1)}(\mathbf{k}) \\ \mathbf{e}^{(2)}(\mathbf{k}) \end{bmatrix} = \begin{bmatrix} \cos\theta(\mathbf{k}) & -\sin\theta(\mathbf{k}) \\ \sin\theta(\mathbf{k}) & \cos\theta(\mathbf{k}) \end{bmatrix} \begin{bmatrix} \mathbf{e}''^{(1)}(\mathbf{k}) \\ \mathbf{e}''^{(2)}(\mathbf{k}) \end{bmatrix}.$$

We define a unitary operator u_1 on $L^2(\mathbb{R}^3_k \times \{1,2\})$ by

$$\begin{bmatrix} (u_1 f)(\mathbf{k}, 1) \\ (u_1 f)(\mathbf{k}, 2) \end{bmatrix} := \begin{bmatrix} \cos \theta(\mathbf{k}) & -\sin \theta(\mathbf{k}) \\ \sin \theta(\mathbf{k}) & \cos \theta(\mathbf{k}) \end{bmatrix} \begin{bmatrix} f(\mathbf{k}, 1) \\ f(\mathbf{k}, 2) \end{bmatrix}, \quad \mathbf{k} \in \mathbb{R}^3.$$

The operator $U(\mathbf{e} \leftarrow \mathbf{e}'') := \Gamma(u_1)$ is a unitary operator on \mathcal{F}_{rad} . It is clear that

$$U(\mathbf{e} \leftarrow \mathbf{e}'')d\Gamma(\omega)U(\mathbf{e} \leftarrow \mathbf{e}'')^* = d\Gamma(\omega).$$

By the equality $u_1 \mathbf{g}''(\cdot, \mathbf{x}) = \mathbf{g}(\cdot, \mathbf{x})$, we have $U(\mathbf{e} \leftarrow \mathbf{e}'') \mathbf{A}''(\hat{\mathbf{x}}) U(\mathbf{e} \leftarrow \mathbf{e}'')^* = \mathbf{A}(\hat{\mathbf{x}})$. Therefore we get

$$U(\mathbf{e} \leftarrow \mathbf{e}'')\overline{H''}U(\mathbf{e} \leftarrow \mathbf{e}'')^* = \overline{U(\mathbf{e} \leftarrow \mathbf{e}'')H''U(\mathbf{e} \leftarrow \mathbf{e}'')^*} = \overline{H}.$$

This means that the operator H'' is essentially self-adjoint and $\overline{H''}$ is unitarily equivalent to \overline{H} . Next we show that $\overline{H''}$ is unitarily equivalent to $\overline{H'}$. Let u_2 be a unitary operator on $L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$ such that

$$(u_2 f)(\mathbf{k}, \lambda) := \begin{cases} -f(\mathbf{k}, 2), & \mathbf{k} \in O, \\ f(\mathbf{k}, \lambda), & \text{otherwise.} \end{cases}$$

It is easy to see that $u_1g'_j(\cdot, \mathbf{x}) = g''_j(\cdot, \mathbf{x}), j = 1, 2, 3$. Then $U(\mathbf{e}'' \leftarrow \mathbf{e}') := \Gamma(u_2)$ is a unitary transformation on \mathcal{F}_{rad} , and

$$U(\mathbf{e''} \leftarrow \mathbf{e'})d\Gamma(\omega)U(\mathbf{e''} \leftarrow \mathbf{e'})^* = d\Gamma(\omega).$$

By the definition of u_2 , the equality $U(\mathbf{e''} \leftarrow \mathbf{e'})\mathbf{A'}(\hat{\mathbf{x}})U(\mathbf{e''} \leftarrow \mathbf{e'})^* = \mathbf{A''}(\hat{\mathbf{x}})$ holds. Hence

$$U(\mathbf{e}'' \leftarrow \mathbf{e}')\overline{H'}U(\mathbf{e}'' \leftarrow \mathbf{e}')^* = \overline{U(\mathbf{e}'' \leftarrow \mathbf{e}')H'U(\mathbf{e}'' \leftarrow \mathbf{e}')^*} = \overline{H''},$$

which implies that H' is essentially self-adjoint and $\overline{H'}$ is unitarily equivalent to $\overline{H''}$. We set

$$U(\mathbf{e} \leftarrow \mathbf{e}') := U(\mathbf{e} \leftarrow \mathbf{e}'')U(\mathbf{e}'' \leftarrow \mathbf{e}').$$

Then $U(\mathbf{e} \leftarrow \mathbf{e}')\overline{H'}U(\mathbf{e} \leftarrow \mathbf{e}')^* = \overline{H}$. Thus Theorem A.1 is proved. The proof of Theorem A.2 is similar.

Appendix B. Remarks on the angular momentum

As is shown in Appendix A, spectral properties of QED models are independent of the choice of polarization vectors. Hence, in the definition of QED models, usually we do not need to specify them. However, the angular momentum of the electromagnetic field depends on the choice of polarization vectors, since the angular momentum does not commute with $U(\mathbf{e} \leftarrow \mathbf{e}')$. Therefore, when we discuss an angular momentum, we carefully specify the choice of polarization vectors. One can find the definition of an angular momentum for the electromagnetic field in the textbook [Sp, Section 13.5] (see also [Hi]). In this appendix, we propose an alternate definition.

Let (H, \mathbf{e}) be the pair of a Hamiltonian and polarization vectors.

For each unit vector $\mathbf{j} \in \mathbb{R}^3$, we can define specific polarization vectors $\mathbf{\bar{e}} = (\bar{\mathbf{e}}^{(1)}, \bar{\mathbf{e}}^{(2)})$ by

$$(\mathrm{B.1}) \qquad \qquad \bar{\mathbf{e}}^{(1)}(\mathbf{k}) := \frac{\mathbf{k} \wedge \mathbf{j}}{|\mathbf{k} \wedge \mathbf{j}|}, \quad \bar{\mathbf{e}}^{(2)}(\mathbf{k}) := \frac{\mathbf{k}}{|\mathbf{k}|} \wedge \bar{\mathbf{e}}^{(1)}(\mathbf{k}).$$

For the Dirac–Maxwell model $(H, \bar{\mathbf{e}})$, we define the angular momentum around the \mathbf{j} -axis by

$$L_{\mathbf{j}}(\bar{\mathbf{e}}) := d\Gamma(\overline{\mathbf{j} \cdot \ell}),$$

where

$$\vec{\ell} := (\ell_1, \ell_2, \ell_3) := i(\nabla_{\mathbf{k}} \wedge \mathbf{k})$$

is a triplet of self-adjoint operators acting on $L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\})$.

Let $\mathbf{e} = (\mathbf{e}^{(1)}, \mathbf{e}^{(2)})$ be any polarization vectors. The angular momentum around the **j**-axis in the Dirac–Maxwell model (H, \mathbf{e}) is defined by

$$L_{\mathbf{j}}(\mathbf{e}) := U(\mathbf{e} \leftarrow \bar{\mathbf{e}})L_{\mathbf{j}}(\bar{\mathbf{e}})U(\mathbf{e} \leftarrow \bar{\mathbf{e}})^*,$$

where $U(\bar{\mathbf{e}} \leftarrow \mathbf{e})$ is the unitary operator defined in Appendix A. By the chain rule for $U(\mathbf{e} \leftarrow \mathbf{e}')$, the angular momentums transform as

$$L_{\mathbf{j}}(\mathbf{e}) = U(\mathbf{e} \leftarrow \mathbf{e}')L_{\mathbf{j}}(\mathbf{e}')U(\mathbf{e} \leftarrow \mathbf{e}')^*,$$

where \mathbf{e} and \mathbf{e}' are arbitrary polarization vectors.

Appendix C. Some properties of the lowest energy

In this appendix, we show some properties of $E_{\nu}(\mathbf{p})$ which are used in the proofs of Theorems 4.1–4.4.

Proposition C.1 (Concavity). $E_{\nu}(\mathbf{p})$ is concave in $(\mathbf{p}, m, q) \in \mathbb{R}^3 \times \mathbb{R} \times \mathbb{R}$.

Proof. See [A2].
$$\Box$$

Proposition C.2 (Continuity). $E_{\nu}(\mathbf{p}, m)$ is Lipschitz continuous in (\mathbf{p}, m) , i.e.,

$$|E_{\nu}(\mathbf{p},m) - E_{\nu}(\mathbf{p}',m')| \le \sqrt{|\mathbf{p} - \mathbf{p}'|^2 + |m - m'|^2}, \quad \mathbf{p}, \mathbf{p}' \in \mathbb{R}^3, m, m' \in \mathbb{R}.$$

Proof. See [A2].
$$\Box$$

Proposition C.3 (Reflection symmetry in m). The Hamiltonian $\overline{H_{\nu}(\mathbf{p}, m)}$ is unitarily equivalent to $\overline{H_{\nu}(\mathbf{p}, -m)}$. In particular

$$E_{\nu}(\mathbf{p}, m) = E_{\nu}(\mathbf{p}, -m), \quad E_{\nu}(\mathbf{p}, m) \le E_{\nu}(\mathbf{p}, 0).$$

Proof. Let $\gamma_5 := -i\alpha_1\alpha_2\alpha_3$. Then γ_5 is a unitary operator and $\gamma_5\overline{H_{\nu}(\mathbf{p},m)}\gamma_5^* = \overline{H_{\nu}(\mathbf{p},-m)}$. Therefore $E_{\nu}(\mathbf{p},m) = E_{\nu}(\mathbf{p},-m)$. By Proposition C.1, $m \mapsto E_{\nu}(\mathbf{p},m)$ is concave. Hence $E_{\nu}(\mathbf{p},0) = E_{\nu}(\mathbf{p},\frac{1}{2}m-\frac{1}{2}m) \geq E_{\nu}(\mathbf{p},m)$.

Proposition C.4 (Rotation invariance of the total momentum). Let $T \in O(3)$. Assume that $|\hat{\rho}(\mathbf{k})| = |\hat{\rho}(T\mathbf{k})|$ for a.e. $\mathbf{k} \in \mathbb{R}^3$. Then $\overline{H_{\nu}(\mathbf{p})}$ is unitarily equivalent to $\overline{H_{\nu}(T\mathbf{p})}$. In particular, $E_{\nu}(\mathbf{p}) = E_{\nu}(T\mathbf{p})$.

Proof. For $T \in O(3)$, we define four 4×4 matrices by

$$\beta' := \beta, \quad \alpha'_j := \sum_{l=1}^3 T_{j,l} \alpha_l, \quad j = 1, 2, 3;$$

they obey $\{\alpha'_j, \beta'\} = 0$, $\{\alpha'_j, \alpha'_l\} = 2\delta_{j,l}$, j, l = 1, 2, 3. Then there exists a 4×4 unitary matrix u_T such that (see [T, Lemma 2.25])

$$u_T \alpha_j u_T^{-1} = \sum_{k=1}^3 T_{j,k} \alpha_k, \quad u_T \beta u_T^{-1} = \beta.$$

Therefore $u_T \boldsymbol{\alpha} \cdot \mathbf{p} u_T^{-1} = \sum_{k,l=1}^3 T_{l,k} \alpha_k p_l = \sum_{k,l=1}^3 \alpha_k (T^{-1})_{k,l} p_l = \boldsymbol{\alpha} \cdot (T^{-1} \mathbf{p})$. Similarly, we have

$$u_T(\boldsymbol{\alpha} \cdot d\Gamma(\mathbf{k}))u_T^{-1} = \boldsymbol{\alpha} \cdot (T^{-1}d\Gamma(\mathbf{k})), \quad u_T\boldsymbol{\alpha} \cdot \mathbf{A}u_T^{-1} = \boldsymbol{\alpha} \cdot (T^{-1}\mathbf{A}) = (T\boldsymbol{\alpha}) \cdot \mathbf{A}.$$

We define a rotation operator \hat{T} of photon momentum by

$$(\hat{T}f)(\mathbf{k},\lambda) = f(T^{-1}\mathbf{k},\lambda), \quad (\mathbf{k},\lambda) \in \mathbb{R}^3_{\mathbf{k}} \times \{1,2\}, f \in L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1,2\}).$$

Then for all $f \in \text{Dom}(k_i \hat{T})$,

$$\hat{T}^{-1}k_i\hat{T}f(\mathbf{k},\lambda) = (k_i\hat{T}f)(T\mathbf{k},\lambda) = (T\mathbf{k})_i(\hat{T}f)(T\mathbf{k},\lambda) = (T\mathbf{k})_if(\mathbf{k},\lambda).$$

Hence we obtain the operator equality $\hat{T}^{-1}k_j\hat{T}=(T\mathbf{k})_j,\ j=1,2,3.$ Thus

$$\Gamma(\hat{T}^{-1})d\Gamma(k_j)\Gamma(\hat{T}) = d\Gamma((T\mathbf{k})_j) = (T \cdot d\Gamma(\mathbf{k}))_j,$$

$$\Gamma(\hat{T}^{-1})H_f(\nu)\Gamma(\hat{T}) = H_f(\nu),$$

$$\Gamma(\hat{T}^{-1})A_j\Gamma(\hat{T}) = \Phi_S(\hat{T}^{-1}g_j), \quad j = 1, 2, 3,$$

where $\Phi_S(\cdot)$ is the Segal field operator (see [RS2, p. 209]) and $g_j(\cdot) := g_j(\cdot, \mathbf{x} = \mathbf{0}) \in L^2(\mathbb{R}^3_{\mathbf{k}} \times \{1, 2\})$. The operator $U := u_T \otimes \Gamma(\hat{T}^{-1})$ is a unitary operator on $\mathbb{C}^4 \otimes \mathcal{F}_{\text{rad}}$ and

(C.1)
$$U\overline{H_{\nu}(\mathbf{p})}U^{-1} = \overline{(\boldsymbol{\alpha}\cdot(T^{-1}\mathbf{p}) + m\beta + H_{f}(\nu) - \boldsymbol{\alpha}\cdot d\Gamma(\mathbf{k}) - q(T\boldsymbol{\alpha})\cdot\Phi_{S}(\hat{T}^{-1}\mathbf{g}))}.$$

Note that T is a 3×3 matrix and \hat{T} is unitary on $L^2(\mathbb{R}^3_{\mathbf{k}}\times\{1,2\})$. Since $T\in O(3)$, we have $(T\boldsymbol{\alpha})\cdot\Phi_S(\hat{T}^{-1}\mathbf{g})=\boldsymbol{\alpha}\cdot T^{-1}\Phi_S(\hat{T}^{-1}\mathbf{g})$, i.e.,

(C.2)
$$(T^{-1}\Phi_S(\hat{T}^{-1}\mathbf{g}))_j = \sum_{l=1}^3 (T^{-1})_{j,l} \Phi_S(\hat{T}^{-1}g_l), \quad j = 1, 2, 3.$$

We define

$$\mathbf{e}'^{(\lambda)}(\mathbf{k}) = T^{-1}\mathbf{e}^{(\lambda)}(T\mathbf{k}), \quad (\mathbf{k}, \lambda) \in \mathbb{R}^3 \times \{1, 2\}.$$

Then $\mathbf{e}'^{(1)}$ and $\mathbf{e}'^{(2)}$ are polarization vectors: $\mathbf{k} \cdot \mathbf{e}'^{(\lambda)}(\mathbf{k}) = 0$, $\mathbf{e}'^{(\lambda)}(\mathbf{k}) \cdot \mathbf{e}'^{(\mu)}(\mathbf{k}) = \delta_{\lambda,\mu}$. Since $|\hat{\rho}(\mathbf{k})| = |\hat{\rho}(T\mathbf{k})|$, there exists a Borel measurable function $\mathbf{k} \mapsto \kappa(\mathbf{k}) \in \mathbb{R}$ such that $\hat{\rho}(T\mathbf{k}) = e^{i\kappa(\mathbf{k})}\hat{\rho}(\mathbf{k})$ for a.e. $\mathbf{k} \in \mathbb{R}^3$. Therefore, we have

(C.3)
$$\sum_{l=1}^{3} (T^{-1})_{j,l} g_l(T\mathbf{k}, \lambda) = \frac{e^{i\kappa(\mathbf{k})} \hat{\rho}(\mathbf{k})}{|\mathbf{k}|^{1/2}} e_j^{\prime(\lambda)}(\mathbf{k}).$$

Let $H'_{\nu}(\mathbf{p})$ be defined just as $H_{\nu}(\mathbf{p})$ with $\mathbf{e}^{(\lambda)}$ replaced by $\mathbf{e}'^{(\lambda)}$. By (C.1)–(C.3), we have

$$U\overline{H_{\nu}(\mathbf{p})}U^* = V\overline{H_{\nu}'(T^{-1}\mathbf{p})}V^*,$$

where $V := \Gamma(e^{i\kappa(\cdot)})$. By Theorem A.2, $\overline{H'_{\nu}(T^{-1}\mathbf{p})}$ is unitarily equivalent to $\overline{H_{\nu}(T^{-1}\mathbf{p})}$. Since $\mathbf{p} \in \mathbb{R}^3$ is arbitrary, $\overline{H_{\nu}(\mathbf{p})}$ is unitarily equivalent to $\overline{H_{\nu}(T\mathbf{p})}$, and $E_{\nu}(\mathbf{p}) = E_{\nu}(T\mathbf{p})$.

If the cutoff function $|\hat{\rho}(\mathbf{k})|$ has reflection symmetry at the origin, the following important inequality holds.

Proposition C.5. Assume that $|\hat{\rho}(\mathbf{k})| = |\hat{\rho}(-\mathbf{k})|$ for almost every $\mathbf{k} \in \mathbb{R}^3$. Then

$$E_{\nu}(\mathbf{p}) \leq E_{\nu}(\mathbf{0}), \quad \mathbf{p} \in \mathbb{R}^3 \setminus \{\mathbf{0}\}.$$

Proof. By the assumption $\hat{\rho}(\mathbf{k}) = \hat{\rho}(-\mathbf{k})$ for a.e. $\mathbf{k} \in \mathbb{R}^3$ and Proposition C.4, we have $E_{\nu}(\mathbf{p}) = E_{\nu}(-\mathbf{p})$, $\mathbf{p} \in \mathbb{R}^3$. Using the concavity of $E_{\nu}(\mathbf{p})$ with respect to \mathbf{p} , we obtain

$$E_{\nu}(\mathbf{0}) = E_{\nu}(\frac{1}{2}\mathbf{p} - \frac{1}{2}\mathbf{p}) \ge \frac{1}{2}E_{\nu}(\mathbf{p}) + \frac{1}{2}E_{\nu}(-\mathbf{p}) = E_{\nu}(\mathbf{p})$$

for all $\mathbf{p} \in \mathbb{R}^3$.

Assuming that $H_{\nu}(\mathbf{0})$ has a ground state, we can obtain the following strict inverse energy inequality:

Proposition C.6. Assume that $|\hat{\rho}(\mathbf{k})| = |\hat{\rho}(-\mathbf{k})|$ for a.e. $\mathbf{k} \in \mathbb{R}^3$. If $\overline{H_{\nu}(\mathbf{0})}$ has a ground state, then

$$E_{\nu}(\mathbf{p}) < E_{\nu}(\mathbf{0})$$
 for all $\mathbf{p} \neq \mathbf{0}$.

Remark C.7. When $\nu > 0$, the massive Hamiltonian $H_{\nu}(\mathbf{0})$ has a ground state (Lemma 6.1). In the massless case $\nu = 0$, $H(\mathbf{0})$ has a ground state under suitable conditions (see Theorems 4.1, 4.2 and 4.4).

Proof of Proposition C.6. Assume that $E_{\nu}(\mathbf{p}) = E_{\nu}(\mathbf{0})$ for some $\mathbf{p} \in \mathbb{R}^3 \setminus \{\mathbf{0}\}$. Let $\Phi_{\nu}(\mathbf{0})$ be a normalized ground state of $H_{\nu}(\mathbf{0})$. For t = 1, -1, we have

$$E_{\nu}(\mathbf{p}) = E_{\nu}(t\mathbf{p}) \le \langle \Phi_{\nu}(\mathbf{0}), H_{\nu}(t\mathbf{p})\Phi_{\nu}(\mathbf{0}) \rangle = t\langle \Phi_{\nu}(\mathbf{0}), \boldsymbol{\alpha} \cdot \mathbf{p}\Phi_{\nu}(\mathbf{0}) \rangle + E_{\nu}(\mathbf{0}).$$

Therefore $\langle \Phi_{\nu}(\mathbf{0}), \boldsymbol{\alpha} \cdot \mathbf{p} \Phi_{\nu}(\mathbf{0}) \rangle = 0$, and hence $\langle \Phi_{\nu}(\mathbf{0}), H_{\nu}(\mathbf{p}) \Phi_{\nu}(\mathbf{0}) \rangle = E_{\nu}(\mathbf{0})$ = $E_{\nu}(\mathbf{p})$, which implies $\|(H_{\nu}(\mathbf{p}) - E_{\nu}(\mathbf{p}))^{1/2} \Phi_{\nu}(\mathbf{0})\| = 0$, and therefore $\Phi_{\nu}(\mathbf{0})$ is a ground state of $H_{\nu}(\mathbf{p})$. Thus $\boldsymbol{\alpha} \cdot \mathbf{p} \Phi_{\nu}(\mathbf{0}) = 0$, and we get a contradiction $|\mathbf{p}|^2 \Phi_{\nu}(\mathbf{0}) = 0$.

If the cutoff function $\hat{\rho}$ is spherically symmetric, the spectral properties of $\overline{H_{\nu}(\mathbf{p})}$ are independent of the direction of \mathbf{p} . The first part of the following proposition immediately follows from Proposition C.4, and the last part from Proposition C.1.

Proposition C.8 (Spherical symmetry in the total momentum). Assume that $|\hat{\rho}(\mathbf{k})|$ is a spherically symmetric function. Then $\overline{H_{\nu}(\mathbf{p})}$ is unitarily equivalent to $\overline{H_{\nu}(\mathbf{p}')}$ for all $\mathbf{p}' \in \mathbb{R}^3$ with $|\mathbf{p}| = |\mathbf{p}'|$. In particular $E_{\nu}(\mathbf{p})$ is spherically symmetric with respect to \mathbf{p} , and $E_{\nu}(\mathbf{p}) \geq E_{\nu}(\mathbf{p}')$ if $|\mathbf{p}| \leq |\mathbf{p}'|$.

Proposition C.9 (Massless limit). $E_{\nu}(\mathbf{p})$ is non-decreasing in $\nu \geq 0$ and

$$\lim_{\nu \to +0} E_{\nu}(\mathbf{p}) = E_0(\mathbf{p}).$$

Proof. Let $\nu \geq \nu' \geq 0$. Then $H_{\nu}(\mathbf{p}) \geq H_{\nu'}(\mathbf{p})$ in the sense of quadratic forms on $\mathcal{D} := \mathrm{Dom}(H_f) \cap \mathrm{Dom}(N_f)$. Therefore $\nu \mapsto E_{\nu}(\mathbf{p})$ is non-decreasing: $E_{\nu}(\mathbf{p}) \geq E_{\nu'}(\mathbf{p})$. It is easy to see that for all $\Psi \in \mathcal{D}$, $H_{\nu}(\mathbf{p})\Psi \to H(\mathbf{p})\Psi$ as $\nu \to 0$. Since \mathcal{D} is a common core for all $H_{\nu}(\mathbf{p})$, we have $H_{\nu}(\mathbf{p}) \to H(\mathbf{p})$ in the strong resolvent sense (see [RS1, Theorem VIII.25]). Using a property of strongly convergent operators [RS1, Theorem VIII.24], we conclude that $E_{\nu}(\mathbf{p}) \to E(\mathbf{p})$ as $\nu \to +0$.

By Proposition C.2,

$$0 \le E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}|, \quad \mathbf{p}, \mathbf{k} \in \mathbb{R}^3.$$

The function $\mathbf{k} \mapsto E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}|$ plays the role of a dispersion relation in the low-energy Dirac polaron.

Theorem C.10. Let $\nu \geq 0$. Assume that $\hat{\rho}$ is spherically symmetric. Suppose that $E_{\nu}(\mathbf{p}, m) < E_{\nu}(\mathbf{p}, 0)$. Then, for $\mathbf{p} \neq \mathbf{0}$,

$$E_{\nu}(\mathbf{p} - \mathbf{k}, m) - E_{\nu}(\mathbf{p}, m) + |\mathbf{k}| \ge \begin{cases} |\mathbf{k}| & \text{if } |\mathbf{p} - \mathbf{k}| \le |\mathbf{p}|, \\ (1 - b_{\nu}(\mathbf{p}))|\mathbf{k}| & \text{if } |\mathbf{p}| \le |\mathbf{p} - \mathbf{k}| \le 2|\mathbf{p}|, \\ (1 - b_{\nu}(\mathbf{p}))|\mathbf{p}| & \text{if } 2|\mathbf{p}| \le |\mathbf{p} - \mathbf{k}|, \end{cases}$$

where

$$b_{\nu}(\mathbf{p}) := \frac{E_{\nu}(\mathbf{p}, m) - E_{\nu}(2\mathbf{p}, m)}{|\mathbf{p}|} < 1.$$

In the case $\mathbf{p} = \mathbf{0}$, for all constant P > 0,

(C.4)
$$E_{\nu}(\mathbf{k}, m) - E_{\nu}(\mathbf{0}, m) + |\mathbf{k}| \ge \begin{cases} \frac{a_{\nu}(P)}{P} |\mathbf{k}| & \text{if } |\mathbf{k}| \le P, \\ a_{\nu}(P) & \text{if } |\mathbf{k}| > P, \end{cases}$$

where

$$a_{\nu}(P) := \left(E_{\nu}(\mathbf{k}, m) - E_{\nu}(\mathbf{0}, m) + |\mathbf{k}| \right) \Big|_{|\mathbf{k}| = P}$$

is a strictly positive constant.

Remark C.11. The idea of the proof of Theorem C.10 was developed in [LMS].

Proof of Theorem C.10. Before proving Theorem C.10, we prove the following lemma:

Lemma C.12. Let $\nu \geq 0$. Assume that $E_{\nu}(\mathbf{p}, m) < E_{\nu}(\mathbf{p}, 0)$. Then

(C.5)
$$E_{\nu}(\mathbf{p} - \mathbf{k}, m) - E_{\nu}(\mathbf{p}, m) + |\mathbf{k}| > 0, \quad \mathbf{k} \in \mathbb{R}^3 \setminus \{\mathbf{0}\}.$$

Proof. First we prove (C.5) for positive $\nu > 0$. We fix $m \neq 0$ and $\mathbf{p} \in \mathbb{R}^3$. Suppose that

(C.6)
$$E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| = 0$$

for some $\mathbf{k} \in \mathbb{R}^3 \setminus \{\mathbf{0}\}$. Let $\Phi_{\nu}(\mathbf{p} - \mathbf{k})$ be a normalized ground state of $H_{\nu}(\mathbf{p} - \mathbf{k})$ (see Lemma 6.1). Then

$$\begin{split} E_{\nu}(\mathbf{p} - \mathbf{k}) &= \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \overline{H_{\nu}(\mathbf{p} - \mathbf{k})} \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle \\ &= \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \overline{H_{\nu}(\mathbf{p})} \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle - \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \alpha \cdot \mathbf{k} \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle \\ &\geq E_{\nu}(\mathbf{p}) - |\mathbf{k}|. \end{split}$$

Hence, by assumption (C.6) we have $\langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \overline{H_{\nu}(\mathbf{p})} \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle = E_{\nu}(\mathbf{p})$ and $\langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \boldsymbol{\alpha} \cdot \mathbf{k} \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle = |\mathbf{k}|$, which implies that $\Phi_{\nu}(\mathbf{p} - \mathbf{k})$ is a ground state

of both $\overline{H_{\nu}(\mathbf{p})}$ and $-\boldsymbol{\alpha} \cdot \mathbf{k}$. Since $\mathbf{k} \neq \mathbf{0}$, we have $\langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}), \beta \Phi_{\nu}(\mathbf{p} - \mathbf{k}) \rangle = 0$, because $\boldsymbol{\alpha} \cdot \mathbf{k} \beta = -\beta \boldsymbol{\alpha} \cdot \mathbf{k}$. In what follows, to emphasize m-dependence, we write $H_{\nu}(\mathbf{p} - \mathbf{k}, m)$ and $\Phi_{\nu}(\mathbf{p} - \mathbf{k}, m)$ for $H_{\nu}(\mathbf{p} - \mathbf{k})$ and $\Phi_{\nu}(\mathbf{p} - \mathbf{k})$, respectively. By using the above facts, we have

$$E_{\nu}(\mathbf{p}, m) = \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m), \overline{H_{\nu}(\mathbf{p}, 0)} \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) \rangle \ge E_{\nu}(\mathbf{p}, 0),$$

which contradicts the inequality $E_{\nu}(\mathbf{p}, m) < E_{\nu}(\mathbf{p}, 0)$.

Next, we handle the case $\nu = 0$. Suppose that there exists a vector $\mathbf{k} \in \mathbb{R}^3 \setminus \{\mathbf{0}\}$ such that $E(\mathbf{p} - \mathbf{k}, m) - E(\mathbf{p}, m) + |\mathbf{k}| = 0$. It is not difficult to see that

$$\lim_{\nu \to +0} \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m), \overline{H(\mathbf{p} - \mathbf{k}, m)} \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) \rangle = E(\mathbf{p} - \mathbf{k}, m).$$

By these equations, we have

(C.7)
$$\lim_{\nu \to +0} \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m), \boldsymbol{\alpha} \cdot \mathbf{k} \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) \rangle = |\mathbf{k}|,$$

(C.8)
$$\lim_{\nu \to +0} \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m), \overline{H(\mathbf{p}, m)} \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) \rangle = E(\mathbf{p}, m).$$

Equation (C.7) implies that

$$\lim_{\nu \to +0} (|\mathbf{k}| - \boldsymbol{\alpha} \cdot \mathbf{k}) \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) = 0.$$

Therefore $\lim_{\nu \to +0} \langle \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m), \beta \Phi_{\nu}(\mathbf{p} - \mathbf{k}, m) \rangle = 0$. This fact and (C.8) imply $E(\mathbf{p}, m) = E(\mathbf{p}, 0)$, which contradicts $E(\mathbf{p}, m) < E(\mathbf{p}, 0)$.

We fix a vector \mathbf{p} such that $E_{\nu}(\mathbf{p}, m) < E_{\nu}(\mathbf{p}, 0)$. Since $\hat{\rho}$ is spherically symmetric, by Proposition C.8 the function

$$G_{\nu}(|\mathbf{k}|) := E_{\nu}(\mathbf{0}) - E_{\nu}(\mathbf{k}), \quad \mathbf{k} \in \mathbb{R}^3,$$

is non-decreasing, convex with respect to $|\mathbf{k}|$, and

(C.9)
$$0 < G_{\nu}(|\mathbf{k}|) < |\mathbf{k}|, \quad \mathbf{k} \in \mathbb{R}^3.$$

Since $G_{\nu}(s)$ is convex, $G_{\nu}(s)$ has a right derivative $G_{\nu}^{+\prime}(s)$:

$$G_{\nu}^{+\prime}(s) := \lim_{h \to +0} [G_{\nu}(s+h) - G_{\nu}(s)]/h.$$

First we show that

(C.10)
$$G_{\nu}^{+\prime}(s) < 1, \quad 0 \le s \le |\mathbf{p}|.$$

Since $G_{\nu}(s)$ is convex and $0 \le G_{\nu}(s) \le s$, $G_{\nu}^{+'}(s)$ is a non-decreasing function of s. If $G_{\nu}^{+'}(s_0) > 1$ for a constant $s_0 \ge 0$, then $G_{\nu}^{+'}(s) > 1$ for all $s \ge s_0$, and

$$G_{\nu}(s) = \int_{s_0}^{s} G_{\nu}^{+\prime}(t) dt + \int_{0}^{s_0} G_{\nu}^{+\prime}(t) dt \ge (s - s_0) G_{\nu}^{+\prime}(s_0) + \int_{0}^{s_0} G_{\nu}^{+\prime}(t) dt$$

for all $s > s_0$. This contradicts (C.9). Thus, $G_{\nu}^{+'}(s) \le 1$ for all $s \ge 0$. Let $s_1 \ge 0$ be a point such that $G_{\nu}^{+'}(s_1) = 1$ and $G_{\nu}^{+'}(s_1 - \epsilon) < 1$ for all $0 < \epsilon \le s_1$. If $|\mathbf{p}| < s_1$, (C.10) is trivial, so we consider the case $|\mathbf{p}| \ge s_1$. Note that $G_{\nu}^{+'}(s) = 1$ for all $s \ge s_1$. Hence $G_{\nu}(s)$ is a linear function of s if $s \ge s_1$:

$$G_{\nu}(s) = s + C, \quad s \ge s_1,$$

where C is a negative constant. By this equality, we have

$$E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| = -|\mathbf{p} - \mathbf{k}| + |\mathbf{p}| + |\mathbf{k}|,$$

for all \mathbf{p} and \mathbf{k} such that $|\mathbf{p} - \mathbf{k}| \ge s_1$ and $|\mathbf{p}| \ge s_1$. We choose $\mathbf{k} = -C\mathbf{p}$ for a constant $C > s_1/|\mathbf{p}|$. Then

$$E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| = 0,$$

contrary to Lemma C.12. Therefore $G_{\nu}^{+\prime}(s) < 1$ for all $0 \le s \le |\mathbf{p}|$.

Next, by using this inequality, we prove Theorem C.10. By (C.10) and convexity of G_{ν} ,

$$c_{\nu}(\mathbf{p}) := \frac{G_{\nu}(|\mathbf{p}|)}{|\mathbf{p}|} \le b_{\nu}(\mathbf{p}) < 1.$$

We define

$$C := \{ J : \mathbb{R}_+ \to \mathbb{R}_+ \mid J \text{ is convex, } 0 \le J(s) \le s \ (s \ge 0),$$
$$J(|\mathbf{p}|) = G_{\nu}(|\mathbf{p}|), \ J(2|\mathbf{p}|) = G_{\nu}(2|\mathbf{p}|) \}.$$

Then

$$E_{\nu}(\mathbf{p} - \mathbf{k}) - E_{\nu}(\mathbf{p}) + |\mathbf{k}| = |\mathbf{k}| + G_{\nu}(\mathbf{p}) - G_{\nu}(\mathbf{p} - \mathbf{k})$$

$$\geq |\mathbf{k}| + G_{\nu}(\mathbf{p}) - \sup_{J \in C} J(\mathbf{p} - \mathbf{k}) =: I.$$

The maximal function in \mathcal{C} is given by the following linear interpolation:

$$J_{\max}(s) := \begin{cases} c_{\nu}(\mathbf{p})s & \text{if } s \leq |\mathbf{p}|, \\ b_{\nu}(\mathbf{p})(s - |\mathbf{p}|) + G_{\nu}(|\mathbf{p}|) & \text{if } |\mathbf{p}| \leq s \leq 2|\mathbf{p}|, \\ s - 2|\mathbf{p}| + G_{\nu}(2|\mathbf{p}|) & \text{if } 2|\mathbf{p}| \leq |\mathbf{p} - \mathbf{k}|. \end{cases}$$

Hence

Therefore
$$I \geq |\mathbf{k}| + G_{\nu}(|\mathbf{p}|) - \begin{cases} c_{\nu}(\mathbf{p})|\mathbf{p} - \mathbf{k}| & \text{if } |\mathbf{p} - \mathbf{k}| \leq |\mathbf{p}|, \\ b_{\nu}(\mathbf{p})(|\mathbf{p} - \mathbf{k}| - |\mathbf{p}|) + G_{\nu}(|\mathbf{p}|) & \text{if } |\mathbf{p}| \leq |\mathbf{p} - \mathbf{k}| \leq 2|\mathbf{p}|, \\ |\mathbf{p} - \mathbf{k}| - 2|\mathbf{p}| + G_{\nu}(2|\mathbf{p}|) & \text{if } 2|\mathbf{p}| \leq |\mathbf{p} - \mathbf{k}|. \end{cases}$$

$$= \begin{cases} |\mathbf{k}| + c_{\nu}(\mathbf{p})(|\mathbf{p}| - |\mathbf{p} - \mathbf{k}|) & \text{if } |\mathbf{p} - \mathbf{k}| \leq |\mathbf{p}|, \\ |\mathbf{k}| - b_{\nu}(\mathbf{p})(|\mathbf{p} - \mathbf{k}| - |\mathbf{p}|) & \text{if } |\mathbf{p}| \leq |\mathbf{p} - \mathbf{k}| \leq 2|\mathbf{p}|, \\ |\mathbf{k}| - |\mathbf{p} - \mathbf{k}| + (2 - b_{\nu}(\mathbf{p}))|\mathbf{p}| & \text{if } 2|\mathbf{p}| \leq |\mathbf{p} - \mathbf{k}|. \end{cases}$$

Using the triangle inequality, one can obtain the desired estimate.

Finally, we prove (C.4). Since $G_{\nu}^{+\prime}(0) < 1$ and G_{ν} is convex, the constant $a_{\nu}(P)$ is strictly positive for all P > 0. It is easy to see that

$$G_{\nu}^{+'}(s) \le \frac{G_{\nu}(P)}{P} = \frac{-a_{\nu}(P) + P}{P}, \quad s \le P.$$

Hence

$$E_{\nu}(\mathbf{k}) - E_{\nu}(\mathbf{0}) + |\mathbf{k}| = |\mathbf{k}| - G_{\nu}(|\mathbf{k}|) = \int_{0}^{|\mathbf{k}|} (1 - G_{\nu}^{+\prime}(s)) ds$$

$$\geq \begin{cases} \int_{0}^{|\mathbf{k}|} \left(1 - \frac{G_{\nu}(P)}{P}\right) ds & \text{if } |\mathbf{k}| \leq P, \\ \int_{0}^{P} \left(1 - \frac{G_{\nu}(P)}{P}\right) ds + \int_{P}^{|\mathbf{k}|} (1 - G_{\nu}^{+\prime}(s)) ds & \text{if } |\mathbf{k}| > P, \end{cases}$$

$$\geq \begin{cases} (a_{\nu}(P)/P)|\mathbf{k}| & \text{if } |\mathbf{k}| \leq P, \\ a_{\nu}(P) & \text{if } |\mathbf{k}| > P. \end{cases}$$

This completes the proof.

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